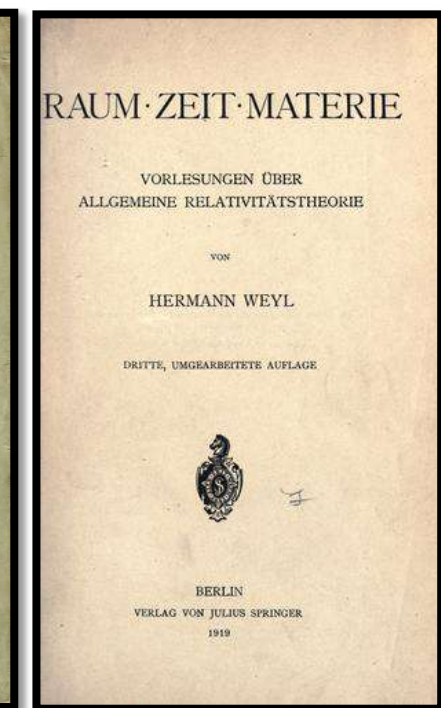
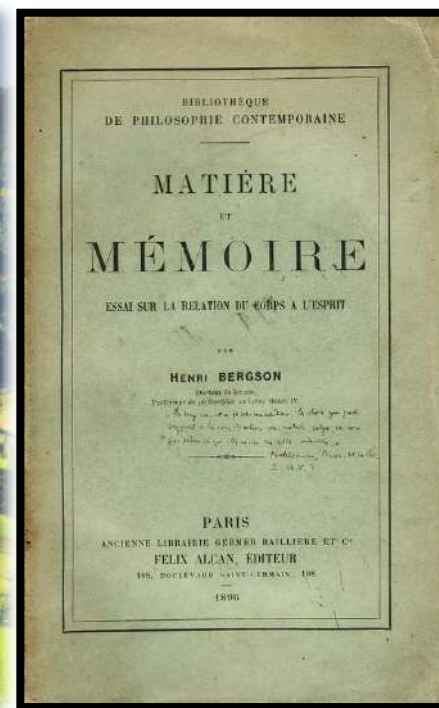


# **Cos'è la materia per la fisica delle particelle e perché tentare di osservarne la creazione**

Francesco Vissani  
INFN Gran Sasso & GSSI  
[vissani@lngs.infn.it](mailto:vissani@lngs.infn.it)



Possiamo far esperienza della meraviglia del creato, e ragionare sulla natura delle cose percettibili.

Un argomento importante di dibattiti anche recenti ruota intorno alla domanda: **cosa è la “materia”**.

È innegabile che la chimica prima e la fisica poi abbiano dato contributi essenziali a questa discussione; esamineremo allora certe questioni sollevate dalla fisica delle particelle, che credo meritino attenzione.

# di cosa ragioneremo

- cosa sostiene la fisica delle particelle sulla natura della materia
- quali quadri concettuali ci fornisce per ordinare le osservazioni
- quali sono quelli più credibili, evidenziando quelle che suggeriscono che la materia sia in una qualche misura, impermanente

# di cosa ragioneremo

- cosa sostiene la fisica delle particelle sulla natura della materia
- quali quadri concettuali ci fornisce per ordinare le osservazioni
- quali sono quelli più credibili, evidenziando quelle che suggeriscono che la materia sia in una qualche misura, impermanente

***è possibile osservare la creazione di particelle di materia in laboratorio?***

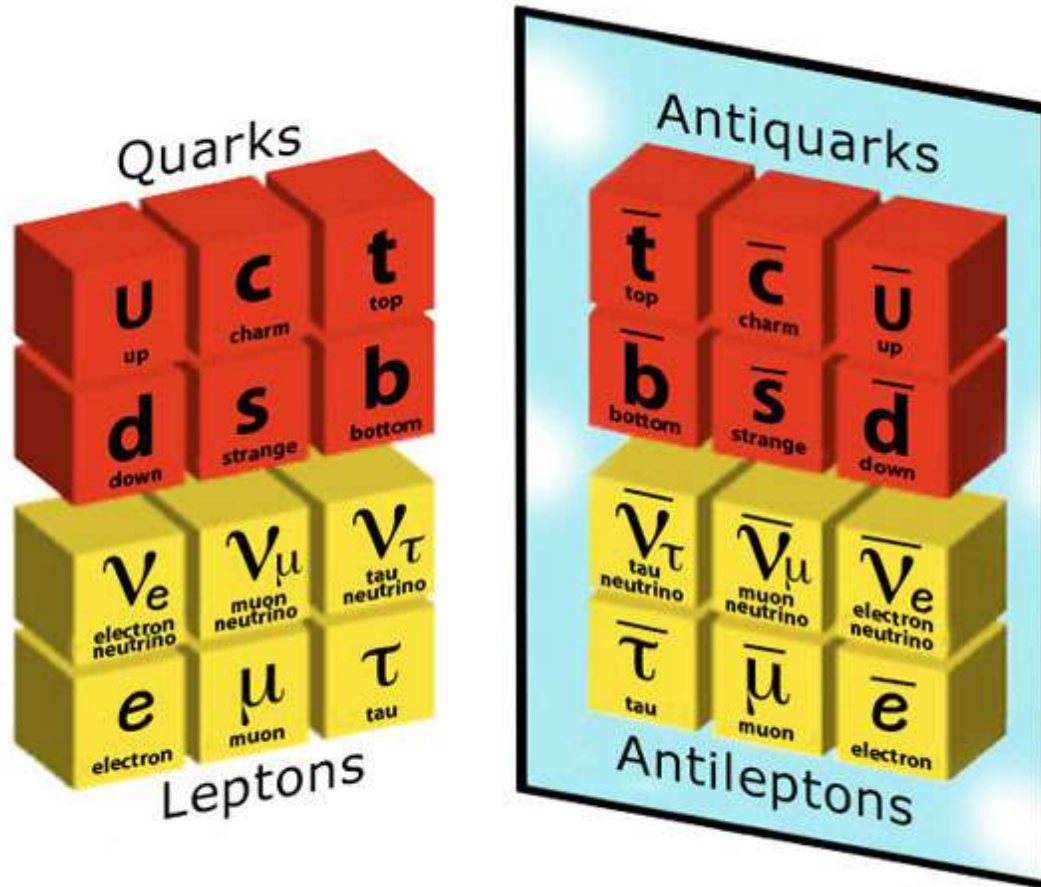
***cosa c'entrano i processi (altrettanto ipotetici) di disintegrazione della materia?***

# la materia per la fisica delle particelle [1]

- la materia sensibile è composta di parti ben definite, dette particelle elementari, ognuna caratterizzata dalle sue interazioni
- le particelle son descritte da rappresentazioni irriducibili e di dimensione finita del gruppo di Poincarè con spin e massa assegnati
- le particelle di materia (note) hanno spin  $\frac{1}{2}$

# elettroni e nuclei atomici → leptoni e quark

da Thomson & Rutherford ai giorni nostri



# meccanica ondulatoria e antiparticelle

$$(P_a)^2 \varphi = m^2 \varphi \quad \text{dove } P_a = i \hbar d/dx_a \quad [\text{Schrodinger, Klein, Fock}]$$

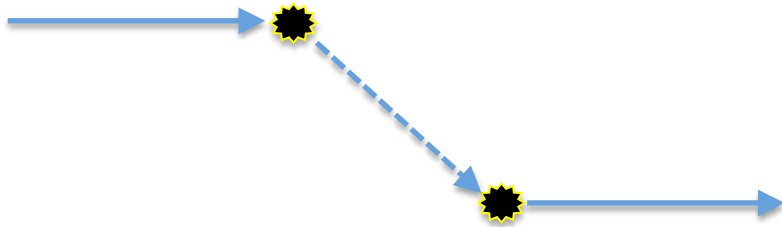
ha soluzioni  $\varphi \sim \exp(\pm i E t / \hbar)$

Gli stati con energie “negative” possono essere pensati come stati finali di una interazione, ovvero come  $(\varphi_-)^*$

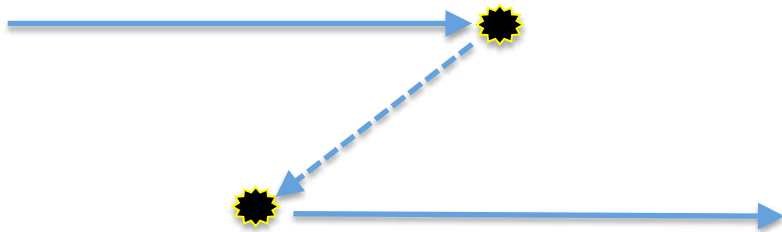
Ma hanno carica opposta alle particelle, in quanto

$$P_a \rightarrow P_a - q A_a \quad [\text{Fock, Weyl, Yang-Mills}]$$

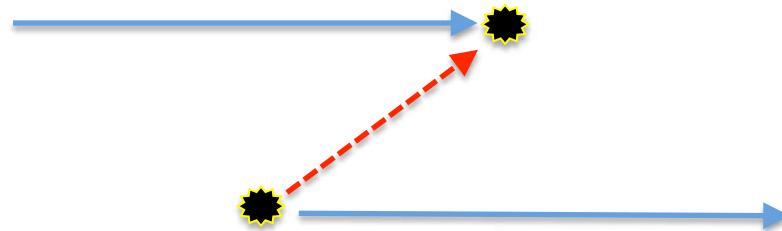
# un altro argomento sulle antiparticelle (Feynman)



un elettrone che cambia direzione due volte interagendo con un potenziale esterno



per certe trasformazioni di Lorentz, lo stato intermedio si muoverebbe all'indietro nel tempo!



ammettendo che esista una particella di carica opposta si evita questa fastidiosa conclusione



# equazione d'onda relativistica dell'elettrone

*The Quantum Theory of the Electron.*

By P. A. M. DIRAC, St. John's College, Cambridge.



(Communicated by R. H. Fowler, F.R.S.—Received January 2, 1928.)

The new quantum mechanics, when applied to the problem of the structure of the atom with point-charge electrons, does not give results in agreement with experiment. The discrepancies consist of “duplexity” phenomena, the observed number of stationary states for an electron in an atom being twice the number given by the theory. To meet the difficulty, Goudsmit and Uhlenbeck have introduced the idea of an electron with a spin angular momentum of half a quantum and a magnetic moment of one Bohr magneton. This model for the electron has been fitted into the new mechanics by Pauli,\* and Darwin,†

## Dirac pensò ben tre anni prima di parlare così:

*“Questo sarebbe un nuovo tipo di particella, sconosciuto alla fisica sperimentale, con la stessa massa e carica opposta dell'elettrone. Possiamo chiamare una tale particella un antielettrone”*



la morale che ne traggo: il fatto che ne facciamo tecnologia non significa che il concetto di antiparticella (e di antimateria) sia banale, o che non meriti discussione.

## in effetti una osservazione apparentemente semplice ma importante fu formulata solo dopo nove anni

$$\gamma^a P_a \psi = m \psi \quad \text{dove } P_a = i \hbar d/dx_a$$

[Dirac '28]

ha soluzioni  $\psi \sim \exp(\pm i E t/\hbar)$

Le soluzioni con energie “negative” possono essere pensate come stati finali di una interazione, ovvero come  $(\psi_-)^*$

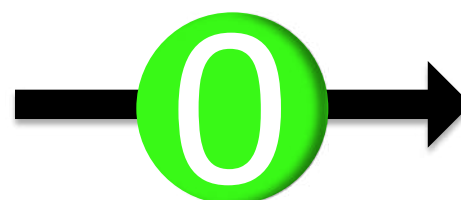
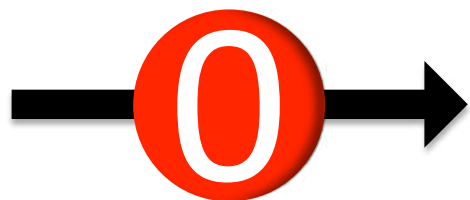
$\psi_-$  e  $\psi_+$  si possono identificare se **la carica è nulla**. Questa risulta chiaro nella base in cui le matrici  $\gamma^a$  sono immaginarie:

$$\psi_- = (\psi_+)^*$$

[Majorana '37]

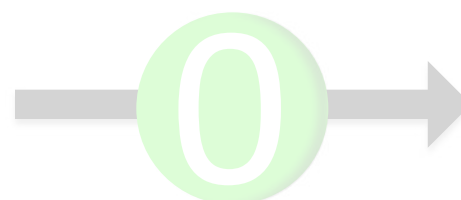
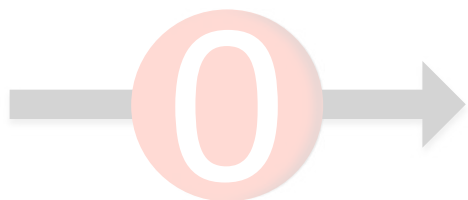
## la materia per la fisica delle particelle [2]

- la meccanica ondulatoria, applicata alle particelle, porta a pensare che ognuna di esse debba essere necessariamente accompagnata da una antiparticella
- alcune particelle neutre coincidono con le loro antiparticelle: p.e. il fotone, la  $Z^0$  oppure il  $\pi^0$
- tra le particelle di materia (ovvero, quelle con spin  $\frac{1}{2}$ ) l'unica che potrebbe obbedire a questa condizione e conformarsi all'ipotesi di Majorana è il **neutrino**

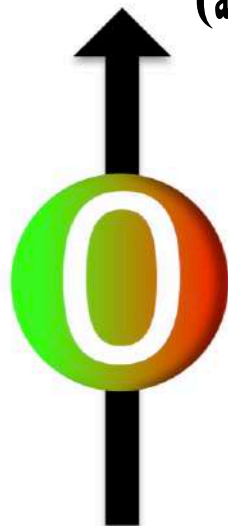


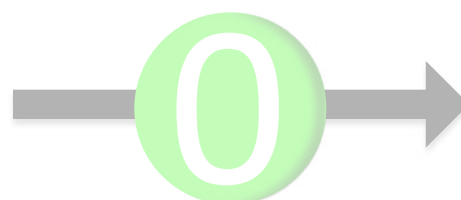
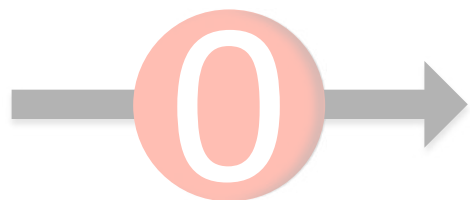
**l'elicità distingue neutrini e antineutrini relativistici**





**ma se son fermi, come possiamo distinguerli?**  
(a meno di ipotesi ad hoc)





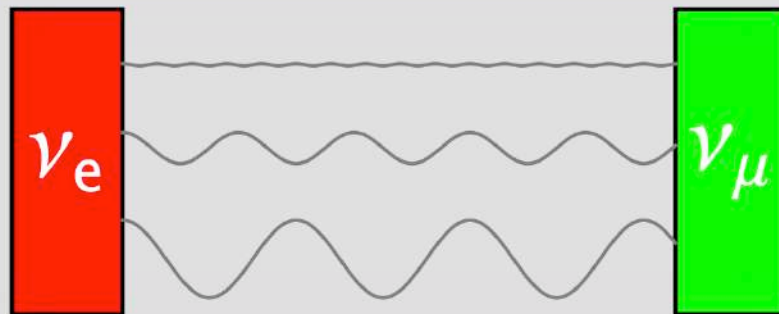
**ipotesi: i neutrini sono materia e antimateria (Majorana)**



## dove abbiamo considerato che:

**l'elicità** dei neutrini, risp. antineutrini, è negativa, risp. positiva in accordo con la struttura V-A delle interazioni deboli  
(notare che come regola si parla solo di particelle ultrarelativistiche)

**la massa** dei neutrini è piccola ma non zero, come dimostrato usando il metodo suggerito da Pontecorvo (intendo le trasformazioni di sapore o oscillazioni, Nobel in fisica 2015)

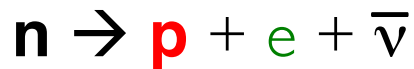
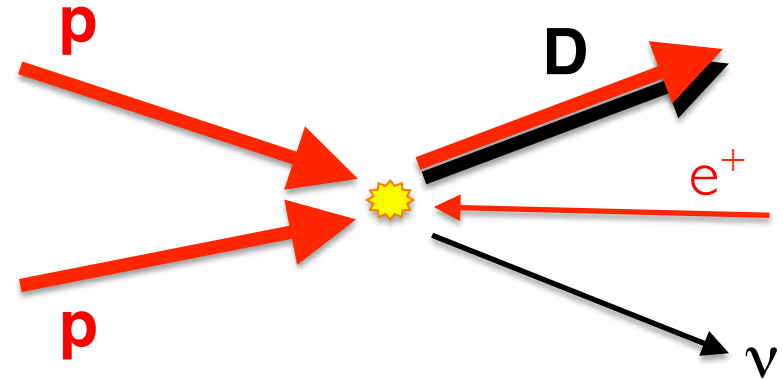




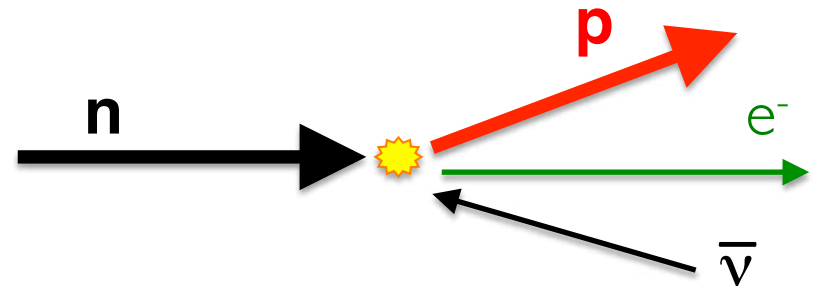
di solito vale  $E_\nu \gg m_\nu c^2$



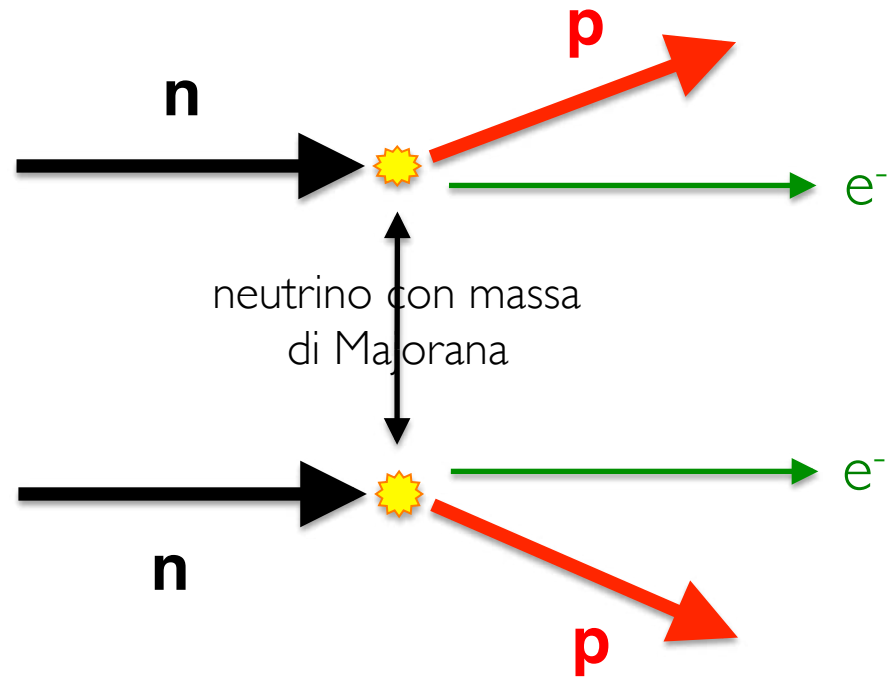
*il neutrino va considerato come la particella di materia che accompagna l'antielettrone*



*l'antineutrino va considerato come la particella di antimateria che accompagna l'elettrone*

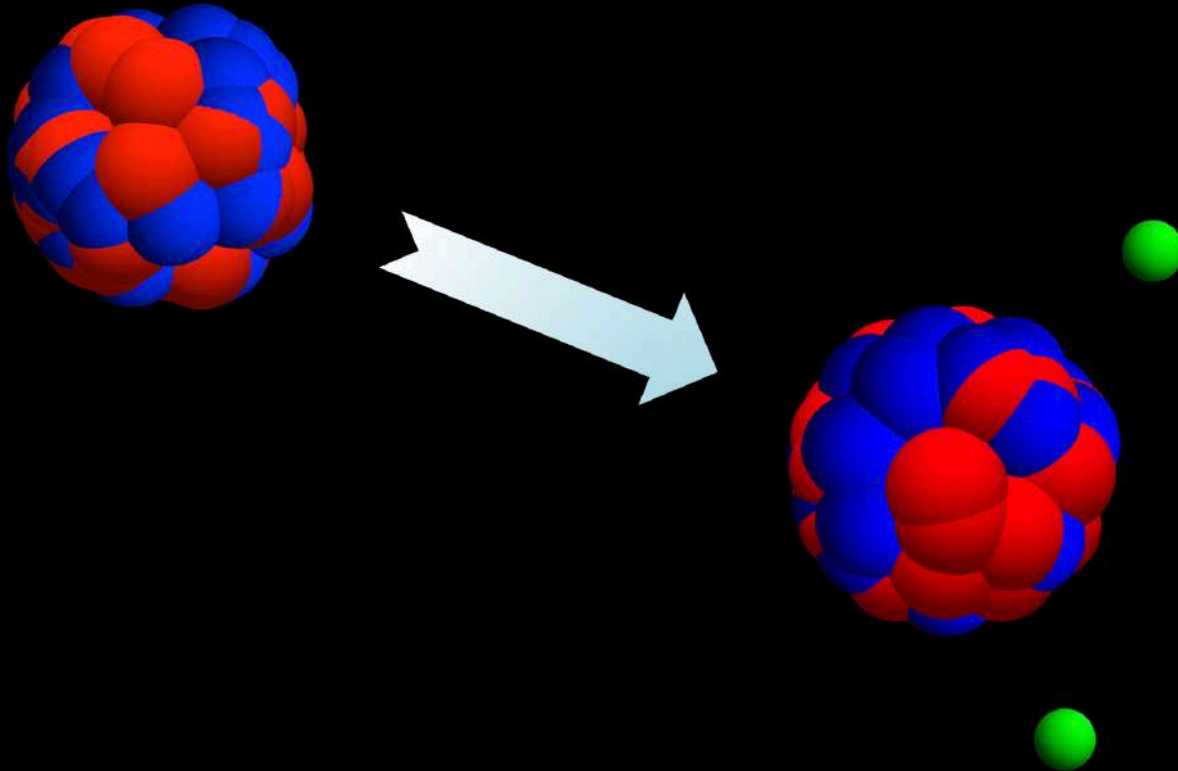


# i neutrini di Majorana rendono possibili nuovi processi



comunemente chiamato “decadimento doppio  $\beta$  senza emissione di neutrini” [Furry '39]

# Il neutrino di Majorana fa da ponte tra materia e antimateria



**si possono creare 2 elettroni - 2 particelle di materia !!!**

## come chiarire le aspettative su $2n \rightarrow 2p + 2e$

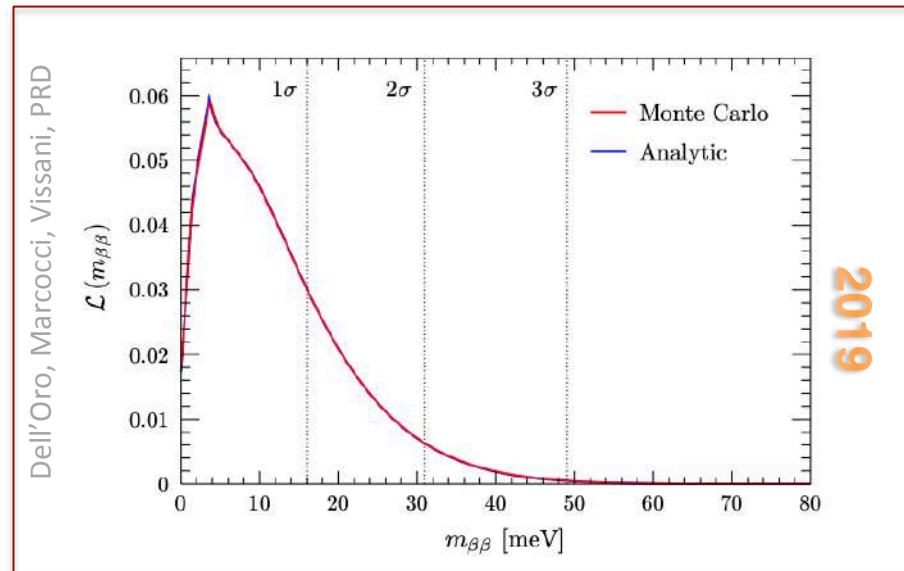
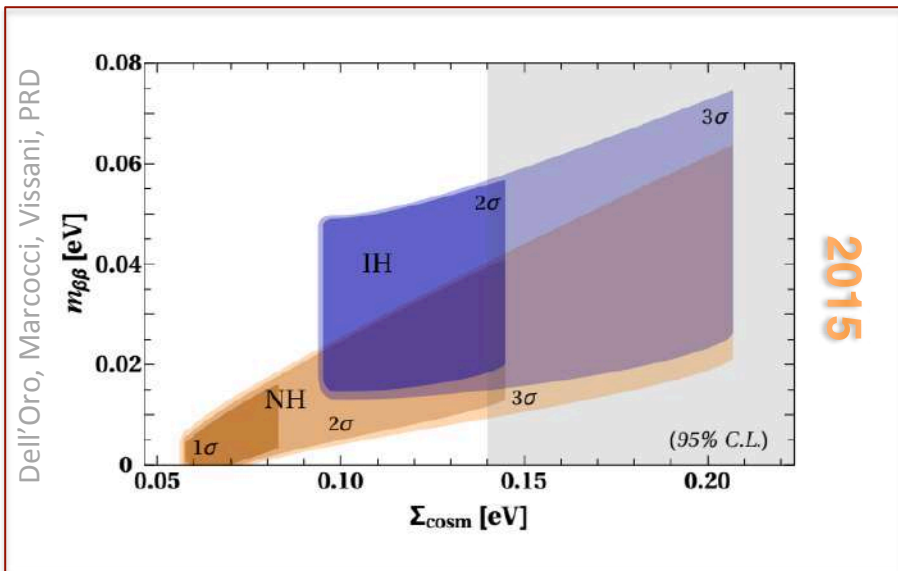
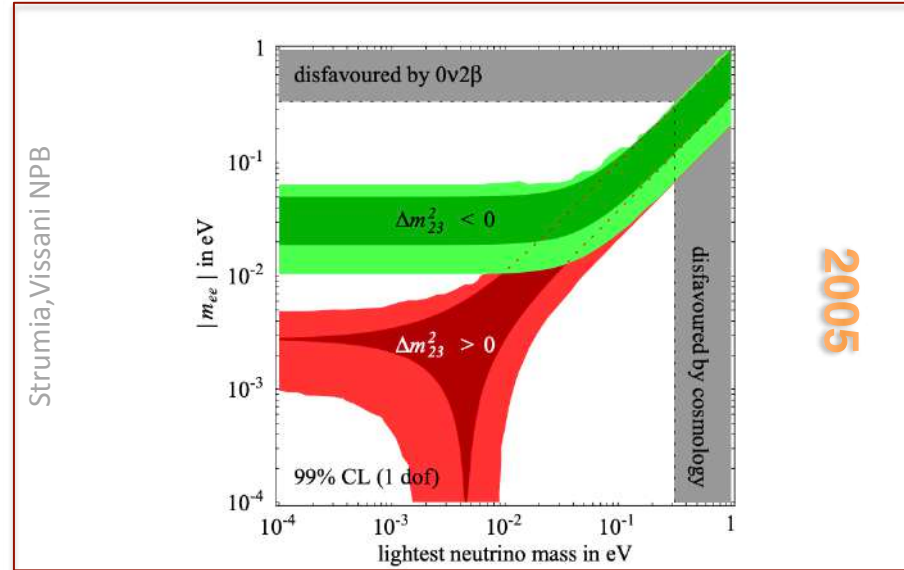
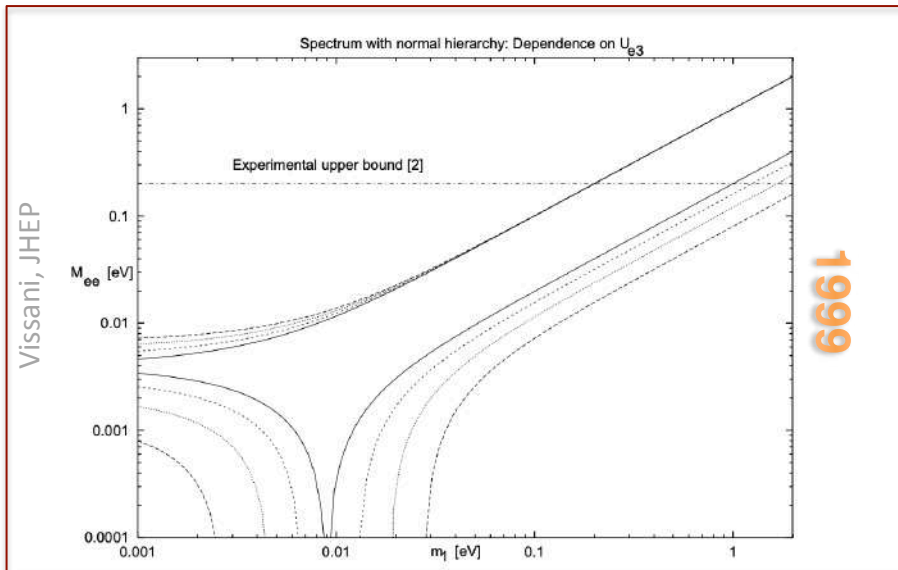
La massa di Majorana che bisogna quantificare è questa,

$$m_{\beta\beta} \equiv |(M_\nu)_{ee}| = \left| \sum_{i=1}^3 |U_{ei}^2| e^{i\xi_i} m_i \right|$$

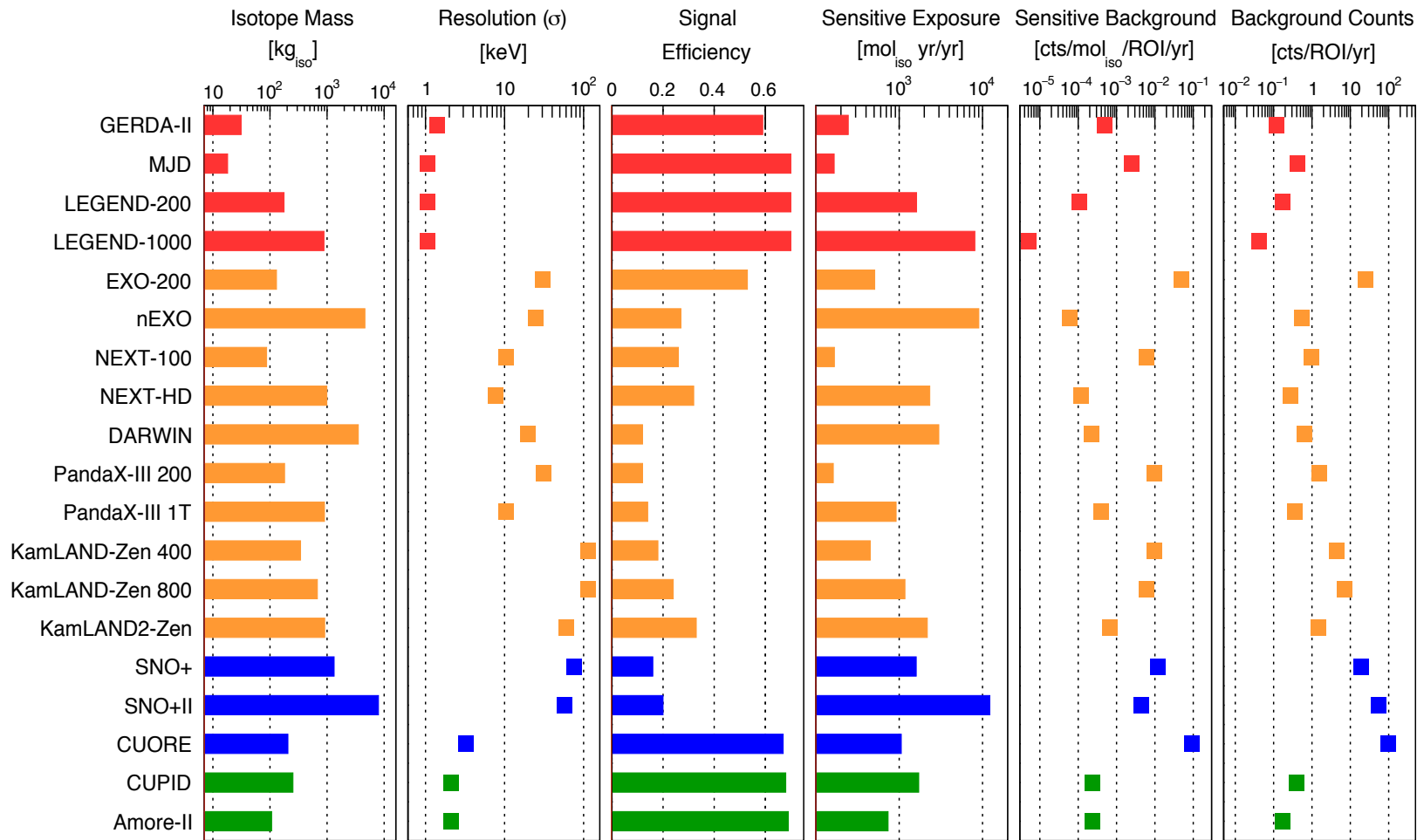
Simboli: il 1° è quello tradizionale, il 2° è l'elemento  $e-e$  della matrice di massa di Majorana

Le oscillazioni misurano solo i valori assoluti del mescolamento leptonic  $|U_{ei}^2|$  e le differenze di massa quadrate, ma non la scala di **massa assoluta** e le **fasi**  $\xi_j$ .

# vincoli sulla massa di Majorana rilevante per $2n \rightarrow 2p + 2e$



# in corso d'opera: studio sistematico delle sensibilità sperimentali



M. Agostini, G. Benato, J. A. Detwiler, J. Menendez, and F. Vissani, article in preparation

## la materia per la fisica delle particelle [3]

- se il neutrino è particella ed antiparticella allo stesso tempo, secondo l'ipotesi di Majorana, esistono processi in cui **si creano elettroni**
- sappiamo che tali transizioni son molto rare. Non siamo ancora in grado di dire precisamente quanto, ma c'è margine per raffinare le aspettative

## la materia per la fisica delle particelle [3]

- se il neutrino è particella ed antiparticella allo stesso tempo, secondo l'ipotesi di Majorana, esistono processi in cui **si creano elettroni**
- sappiamo che tali transizioni son molto rare. Non siamo ancora in grado di dire precisamente quanto, ma c'è margine per raffinare le aspettative

*esaminiamo da ultimo come si inseriscono queste aspettative nel quadro teorico di cui oggi disponiamo: il modello standard*



## [modello standard 1/2] definizione

- gruppo di gauge  $SU(3)_c \times SU(2)_L \times U(1)_Y$  (Weyl, Yang-Mills... GWS)
- 3 famiglie di spinori:  $q_L, l_L, u_R, d_R, e_R$  [6+2+3+3+1=15 s. di Weyl]
- 1 doppietto scalare: H, la cui massa quadra negativa rompe spontaneamente  $SU(2)_L \times U(1)_Y \rightarrow U(1)_{em}$  (Higgs; Brout-Englert, ...)

*si scrivano tutti i termini rinormalizzabili e si fissino i valori dei parametri operando tante misure quanti sono i parametri*

# [modello standard 2/2] implicazioni

- ☑ Una sola scala di massa (o due se contiamo anche la densità di energia)  
17 parametri senza dimensioni (o 18, contando  $\theta_{CP}$ )
- ☑ teoria calcolabile ad ogni ordine perturbativo
- ☑ regime non perturbativo per  $SU(3)_c$  a basse energie (glueballs?).  
gli adroni più leggeri hanno vita media infinita come gli elettroni
- ☑ assenza di correnti neutre fondamentali che violino il sapore adronico;  
CP violata solo debolmente dalle correnti cariche
- ☑ 4 numeri globali,  $\mathbf{B}$ ,  $\mathbf{L}_e$ ,  $\mathbf{L}_\mu$ ,  $\mathbf{L}_\tau$  conservati in teoria delle perturbazioni.

# [modello standard 2/2] implicazioni

- ☑ Una sola scala di massa (o due se contiamo anche la densità di energia)  
17 parametri senza dimensioni (o 18, contando  $\theta_{CP}$ )
- ☑ teoria calcolabile ad ogni ordine perturbativo
- ☑ regime non perturbativo per  $SU(3)_c$  a basse energie (glueballs?).  
gli adroni più leggeri hanno vita media infinita come gli elettroni
- ☑ assenza di correnti neutre fondamentali che violino il sapore adronico;  
CP violata solo debolmente dalle correnti cariche
- ☑ 4 numeri globali,  $\mathbf{B}$ ,  $\mathbf{L}_e$ ,  $\mathbf{L}_\mu$ ,  $\mathbf{L}_\tau$  conservati in teoria delle perturbazioni.  
 $\mathbf{B-L}$ ,  $\mathbf{L}_e-\mathbf{L}_\mu$ ,  $\mathbf{L}_e-\mathbf{L}_\tau$  conservati esattamente, dove  $\mathbf{L}=\mathbf{L}_e+\mathbf{L}_\mu+\mathbf{L}_\tau$

# [modello standard 2/2] implicazioni

- ☑ Una sola scala di massa (o due se contiamo anche la densità di energia)  
17 parametri senza dimensioni (o 18, contando  $\theta_{CP}$ )
- ☑ teoria calcolabile ad ogni ordine perturbativo
- ☑ regime non perturbativo per  $SU(3)_c$  a basse energie (glueballs?).  
gli adroni più leggeri hanno vita media infinita come gli elettroni
- ☑ assenza di correnti neutre fondamentali che violino il sapore adronico;  
CP violata solo debolmente dalle correnti cariche
- ☑ 4 numeri globali,  $\mathbf{B}$ ,  $\mathbf{L}_e$ ,  $\mathbf{L}_\mu$ ,  $\mathbf{L}_\tau$  conservati in teoria delle perturbazioni.  
 $\mathbf{B-L}$ ,  $\mathbf{L}_e-\mathbf{L}_\mu$ ,  $\mathbf{L}_e-\mathbf{L}_\tau$  conservati esattamente, dove  $\mathbf{L}=\mathbf{L}_e+\mathbf{L}_\mu+\mathbf{L}_\tau$
- ☑ **i neutrini non hanno massa, fatto che contraddice le oscillazioni**

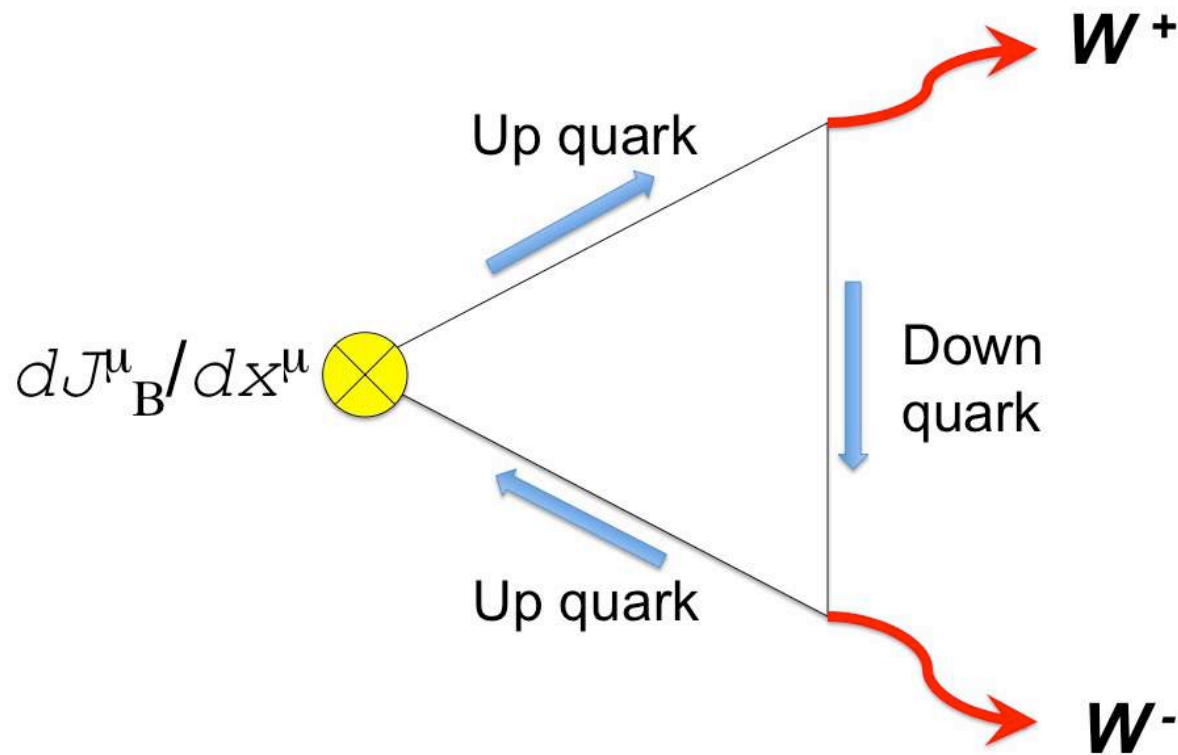
# c'è solo 1 tipo di leptone fondamentale

(=T2K, NOvA, OPERA, SK, DeepCore hanno provato che si conserva solo il numero totale L)

|                                | $\Delta L_e$ | $\Delta L_\mu$ | $\Delta L_\tau$ | $\Delta L$ |
|--------------------------------|--------------|----------------|-----------------|------------|
| $\nu_\mu \rightarrow \nu_e$    | +1           | -1             | 0               | 0          |
| $\nu_\mu \rightarrow \nu_\tau$ | 0            | -1             | +1              | 0          |

Da questo punto di vista, ci sono 2 due tipi di **particelle di materia**: leptoni e quark.  
Le sole simmetrie globali residue sono **L** e **B**.

**nel “modello standard” B e L non sono conservati esattamente!**  
**invece B-L lo è, fatto che collega leptoni & quarks**



[Steinberger '49; Adler; Bell, Jackiw '69]

# c'è solo 1 tipo di leptone fermione fondamentale

(=T2K, NOvA, OPERA, SK, DeepCore hanno provato che si conserva solo il numero totale L)

|                                | $\Delta(L_e - L_\mu)$ | $\Delta(L_\mu - L_\tau)$ | $\Delta(L_\tau - L_e)$ | $\Delta(B - L)$ |
|--------------------------------|-----------------------|--------------------------|------------------------|-----------------|
| $\nu_\mu \rightarrow \nu_e$    | +2                    | -1                       | -1                     | 0               |
| $\nu_\mu \rightarrow \nu_\tau$ | +1                    | -2                       | +1                     | 0               |

Unendo esperimenti+teoria, concludiamo che tutte le simmetrie globali sono violate tranne una

Osservare di  $2n \rightarrow 2p + 2e$  significherebbe che è violata anche l'ultima simmetria globale, **B-L**

# oscillazioni e frontiere del modello standard

Weinberg 79

$$\delta\mathcal{L} = \frac{(\ell H)^2}{M} + \frac{\ell q q q}{M'^2} + \frac{(\ell q d^c)^2}{M''^5} \quad \text{with} \quad \begin{cases} M < 10^{11} \text{ TeV} & \text{for dim.5} \\ M' > 10^{12} \text{ TeV} & \text{for dim.6} \\ M'' > 5 \text{ TeV} & \text{for dim.9} \end{cases}$$

- il 1° (dim.5) **descrive masse di Majorana e viola B-L**
- il 2° (dim.6) causa l'instabilità del protone ma conserva B-L
- il 3° (dim.9) viola il numero leptonico e contribuisce a  $2n \rightarrow 2p + 2e$
- (A dim.7 la simmetria B-L è rotta; a dim.9 anche B è violata)



## [discussione 1/3]: la prospettiva storica

- La fisica delle particelle elementari si configura come un significativo sviluppo della dottrina dell'atomismo, che costituisce una risposta al «problema dell'essere»: gli atomi, pur troppo piccoli da esser visti, sono i *baluardi esistenziali* della materia che ci costituisce.
- Gli aspetti ondulatori della meccanica quantistica invece possono essere idealmente collegati ai precetti delle dottrine pitagoriche: lo stesso Schrödinger diceva *all is waves*.
- Tali aspetti, nella teoria dello spazio-tempo di Einstein con la descrizione di gauge delle interazioni, implicano l'esistenza delle antiparticelle. L'antimateria da esse costituita risulta ben distinta dalla materia nella teoria di gauge di riferimento: il *modello standard*.

## [discussione 1/3]: la prospettiva storica

- La fisica delle particelle elementari si configura come un significativo sviluppo della dottrina dell'atomismo, che costituisce una risposta al «problema dell'essere»: gli atomi, pur troppo piccoli da esser visti, sono i *baluardi esistenziali* della materia che ci costituisce.
- Gli aspetti ondulatori della meccanica quantistica invece possono essere idealmente collegati ai precetti delle dottrine pitagoriche: lo stesso Schrödinger diceva *all is waves*.
- Tali aspetti, nella teoria dello spazio-tempo di Einstein con la descrizione di gauge delle interazioni, implicano l'esistenza delle antiparticelle. L'antimateria da esse costituita risulta ben distinta dalla materia nella teoria di gauge di riferimento: il *modello standard*.

***La particolare natura del neutrino di Majorana implica che la materia, definita secondo il modello standard, può apparire o scomparire. Si riapre il dibattito sul «problema dell'essere», ma nasce una dinamica tra materia ed antimateria che sembra interessante se non addirittura desiderabile.***



## [discussione 2/3]: un tarlo della fisica moderna

*come abbiamo visto, il matrimonio tra meccanica ondulatoria e relatività porta a credere nella inevitabilità dell'esistenza dell'antimateria. Inoltre, le idee di Friedmann (+LRW) e le osservazioni implicano che l'universo fosse un tempo caldissimo.*

- ① certamente, l'antimateria pervadeva l'Universo nel lontano passato
- ② la cosmologia osservativa dice che l'antimateria in pratica è assente
- ③ come armonizzare queste storie? (Sakharov 1967)
- ④ il modello standard non è in grado di rispondere (nonostante B+L sia violata)
- ⑤ bisogna pensare a ulteriori violazioni di **B** e/o di **L**. Si noti che secondo caso è fortemente motivato dall'ipotesi di Majorana

## [discussione 3/3]: aspettative e dubbi

- Dall'unificazione delle costanti di accoppiamento abbiamo degli indizi a favore di una teoria di gauge estesa, non una vera predizione. Sono ignoti i campi scalari, le scale di intermedie, gli effetti di soglia, ecc. Le masse dei fermioni sono uno dei pochi vincoli, la genesi della materia una domanda che può trovare risposte in questi contesti.
- Esistono nuove particelle? Qual'è la natura della materia oscura? Cosa pensare della supersimmetria? La guerra lampo del '70 è diventata una guerra di trincea. Non disponiamo di un «modello standard grande-unificato», ma abbiamo un'importante dato: sappiamo che i neutrini hanno massa. Sarebbe essenziale conoscerla meglio.
- Sembra ragionevole continuare a investigare se la materia possa essere creata, una possibilità suggerita molto tempo fa e che le migliori teorie oggi disponibili, nelle quali il neutrino ha massa di Majorana, concorrono ad indicare per plausibile.



*Grazie per  
l'attenzione!*

## Di cosa è fatta la “materia”?

| componenti elementari | caratteristica identificativa                                                        | vigenza teoria - esp. | ragione della inadeguatezza               |
|-----------------------|--------------------------------------------------------------------------------------|-----------------------|-------------------------------------------|
| atomi                 | <b>specie</b> , massa                                                                | fino 1838 - 1909      | “atomi di elettricità”                    |
| nuclei & elettroni    | <b>massa</b> , <b>carica</b>                                                         | fino 1933 - 1956      | neutroni & neutrini                       |
| p, n, e, $\nu_e$      | <b>B</b> , <b>L<sub>e</sub></b> , spin                                               | fino 1961 - 1968      | quark/SM                                  |
| quark, leptoni        | <b>B</b> , <b>L<sub>e</sub></b> , <b>L<sub>u</sub></b> , <b>L<sub>t</sub></b> , spin | fino 1967 - 2010      | conversione sapore leptonic               |
| quark-antileptoni     | <b>B-L</b> , spin                                                                    | fino 1937/1969 - ?    | $\nu$ Majorana / $2n \rightarrow 2p + 2e$ |
| fermioni              | <b>spin</b>                                                                          | fino ?? - ???         | ???                                       |

# referenze

*Signal of neutrinoless double beta decay, neutrino spectrum and oscillation scenarios*  
[JHEP 06 \(1999\) 022](#), 9906525

*Implications of neutrino data circa 2005*  
con A. Strumia, [Nucl.Phys.B 726 \(2005\) 294](#), 0503246

*Neutrinoless double beta decay: 2015 review* con S. Dell'Oro, S. Marcocci, M. Viel  
[Adv.High Energy Phys. 2016 \(2016\) 2162659](#), 1601.07512

*Empirical inference on the Majorana mass of the ordinary neutrinos*  
con S. Dell'Oro e S. Marcocci, [Phys.Rev.D 100 \(2019\) 7, 073003](#)

*Da Pitagora all'antimateria*, di Francesco Vissani, Published on July 28, 2019  
<https://www.linkedin.com/pulse/da-pitagora-allantimateria-francesco-vissani-phd/>

*A history of some recent attempts to go beyond the standard model*, opening talk at the meeting  
*Gravity, Information and Fundamental Symmetries*, Munich, November 4-6, 2019,  
<https://www.mpi-hd.mpg.de/lin/events/ra-workshop/index.html>

M. Agostini, G. Benato, J. Detwiler, J. Menendez, F. Vissani, rassegna in preparazione per RMP

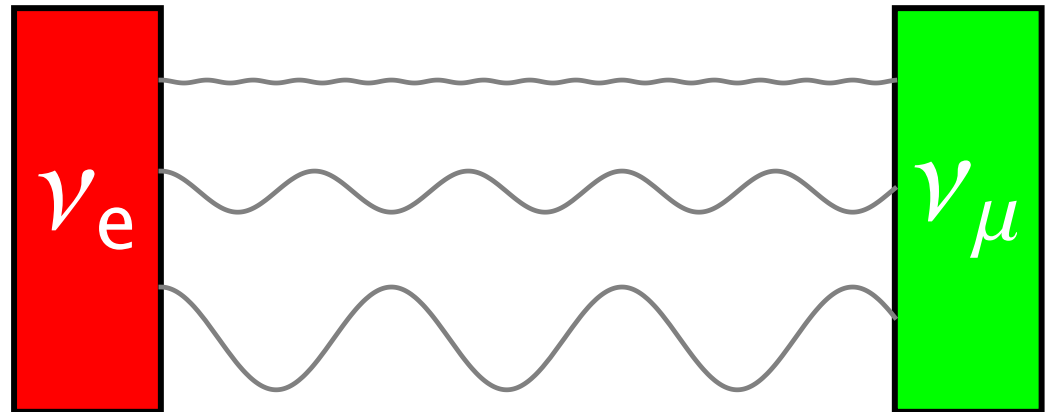
# **materiale a suffragio**

masse dei neutrini in cosmologia,  
oscillazioni, limiti sulla massa di  
Majorana, seesaw, storia



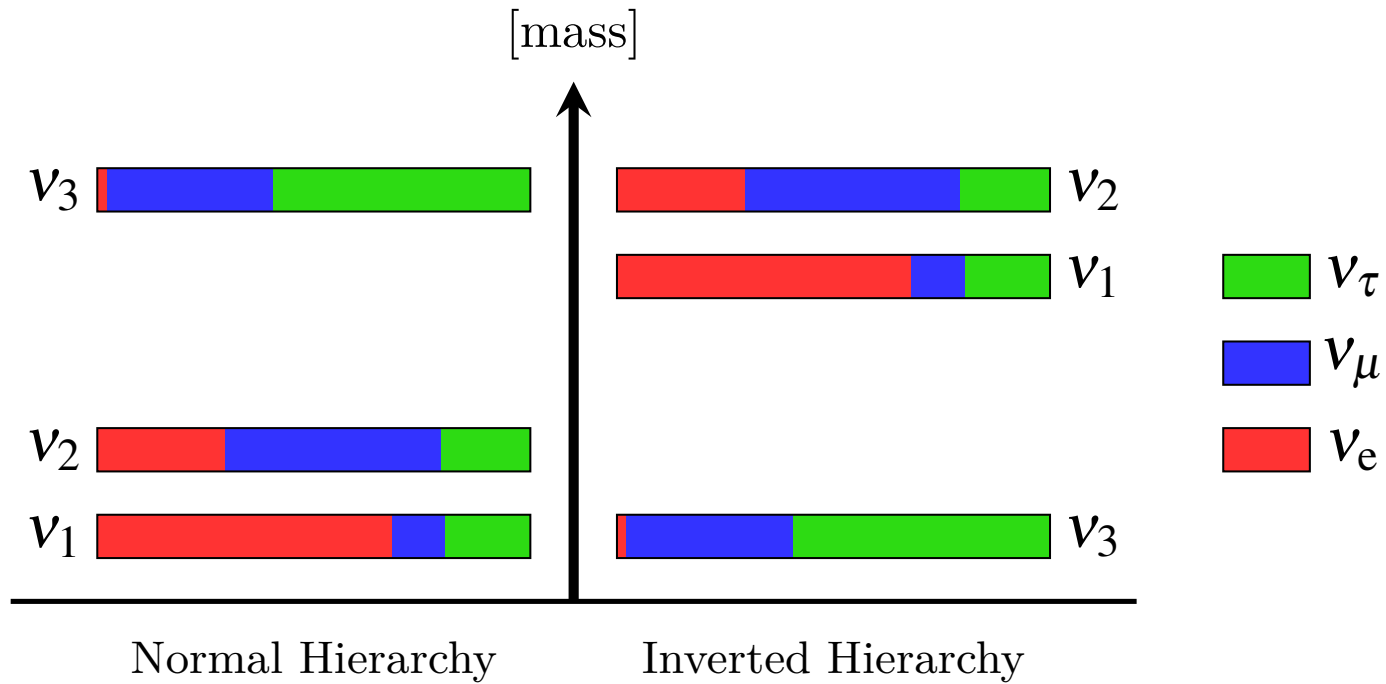
# come funzionano le oscillazioni

Ogni neutrino vien prodotto come miscela di più onde, che descrivono particelle con masse diverse. Dunque le onde hanno velocità diverse, e così, quando si propagano, i neutrini cambiano natura.



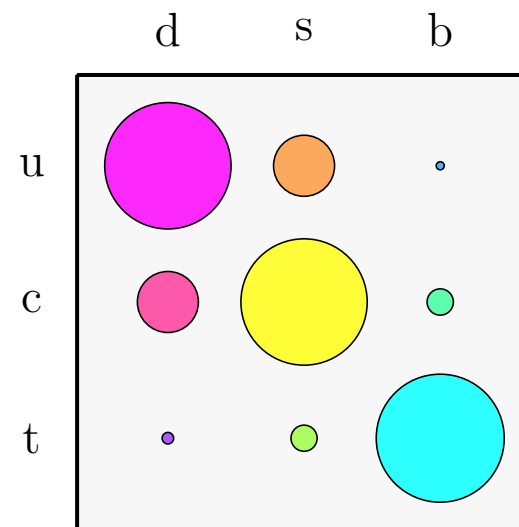
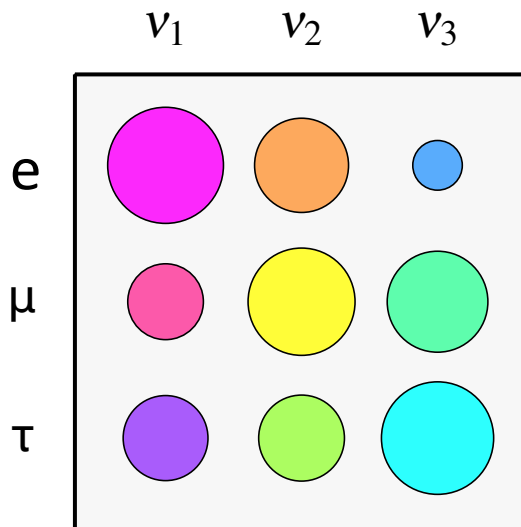
# modello a tre sapori [2]

Fantini, Gallo Rosso et al, 2018

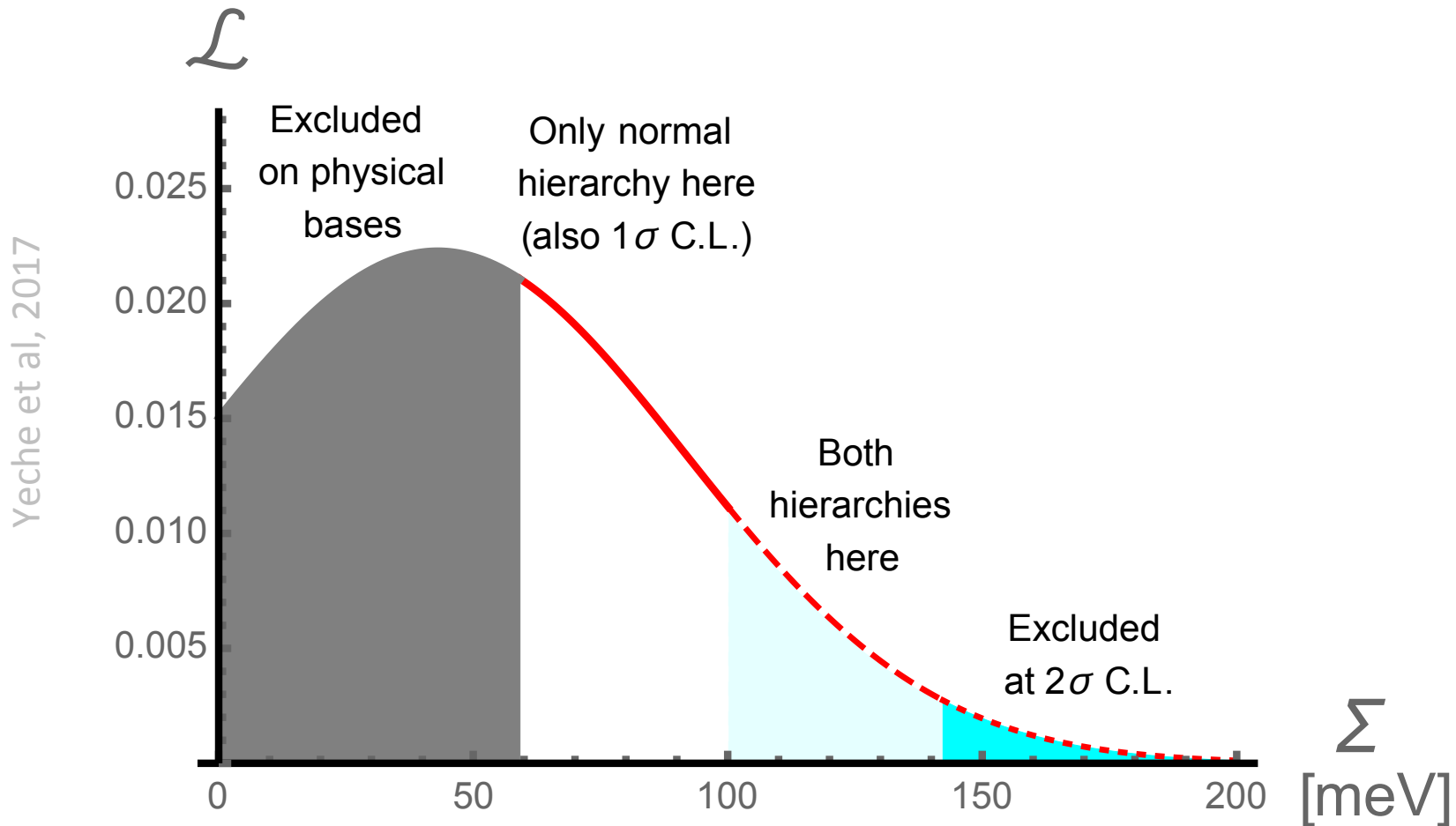


# modello a tre sapori [1]

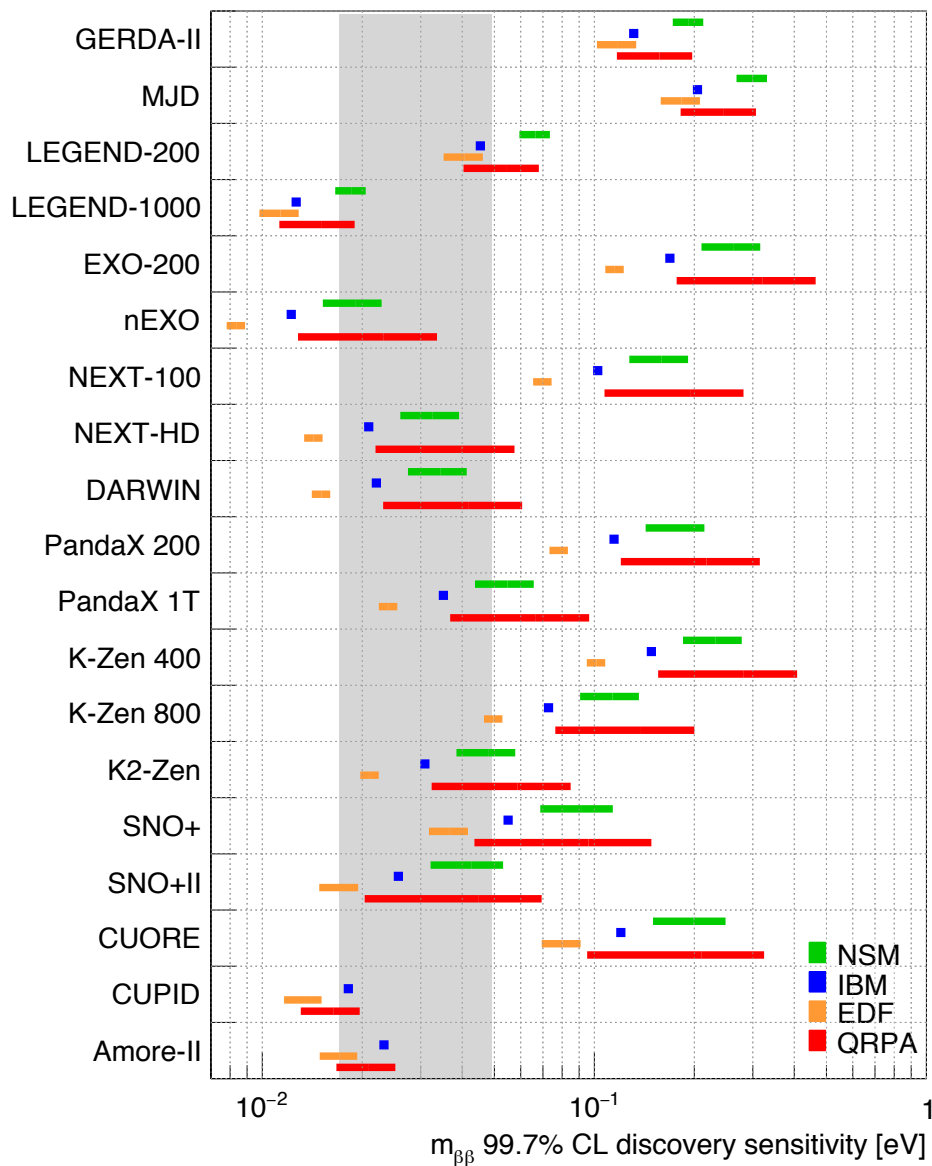
Fantini, Gallo Rosso et al, 2018



# CMB permette di sondare $\Sigma = m_1 + m_2 + m_3$

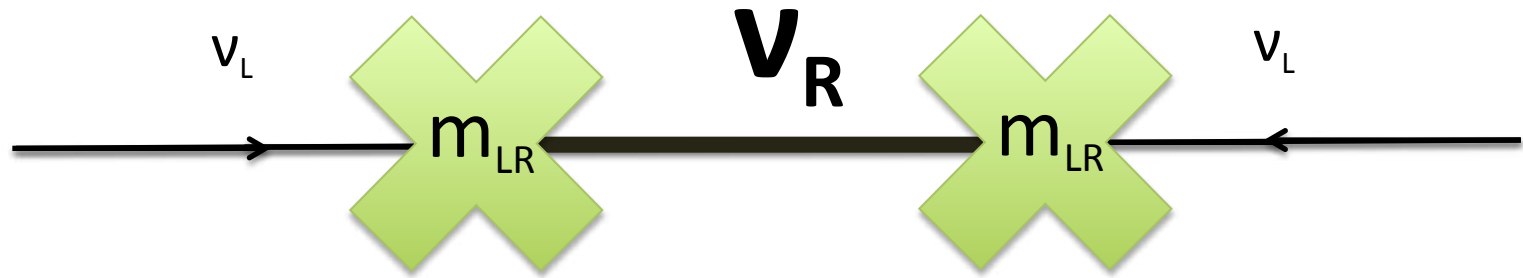


# sensibilità alla massa di Majorana

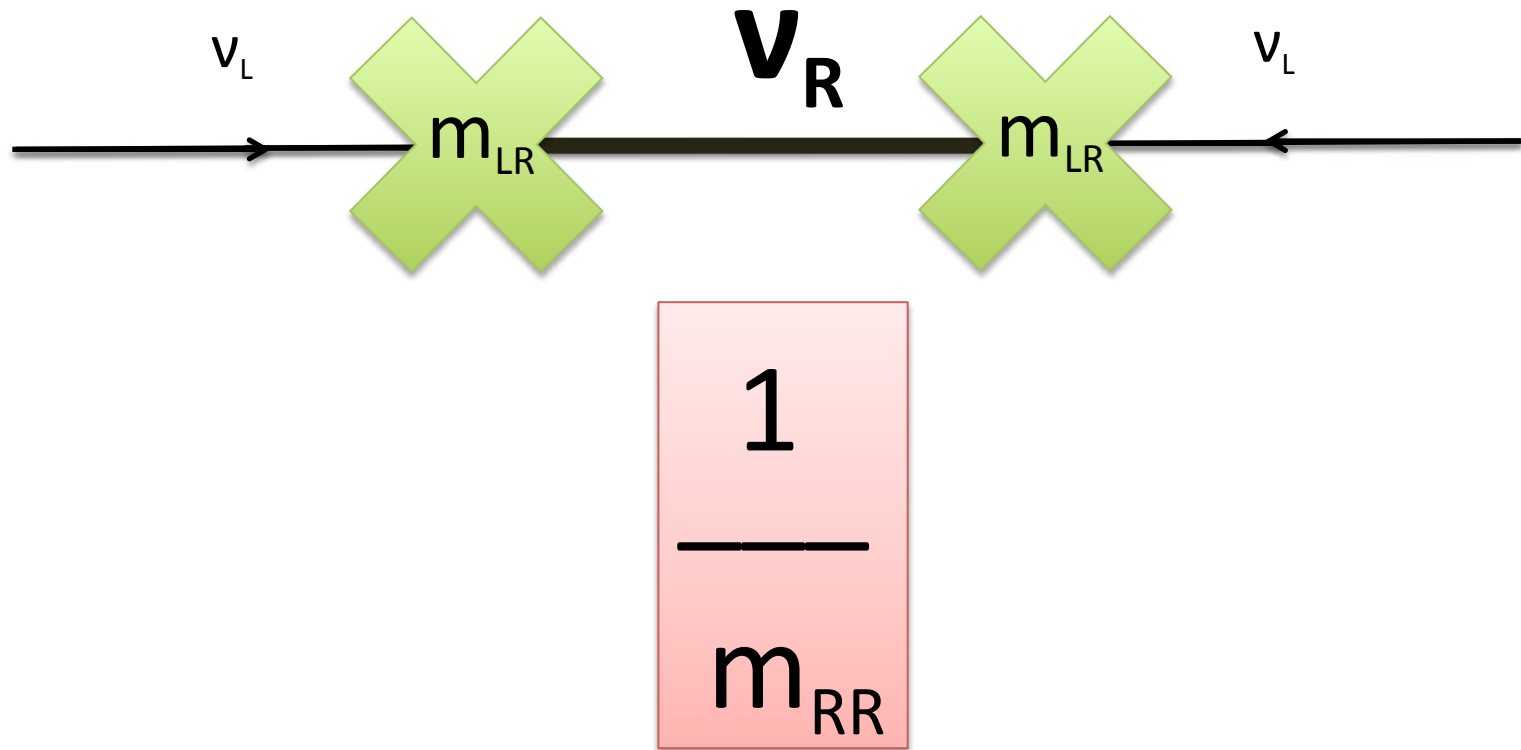


M. Agostini, G. Benato, J. A. DeWitler, J. Menendez, and F. Vissani, article in preparation

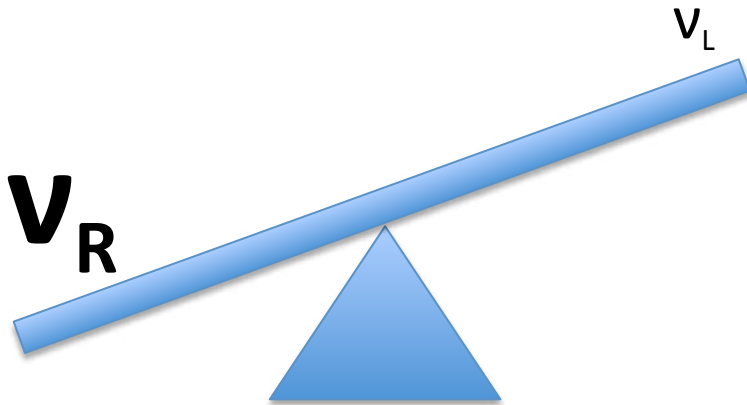
# una spiegazione della piccolezza delle masse



# una spiegazione della piccolezza delle masse



questo è chiamato “seesaw” (altalena)





# cronologia

49

- 1930-1933 scoperta neutrini
- 1937 peculiarità dei neutrini (Majorana)
- 1933-1958 interazioni deboli
- 1957-1998 “oscillazioni” di neutrino
- 1961-1974 modello standard
- 1939-... doppio  $\beta$  senza neutrini
- 1974-... estensione modello standard
- 1990-... massa e cosmologia

# **ancora sulle masse dei neutrini**

dall'idea delle oscillazioni al Nobel  
2015, l'odissea dei neutrini solari,  
analisi globali

**the idea that worked – neutrino oscillations -  
was proposed early (late fifties / early sixties)**



# full recognition 1/2 a century later

Nobelpriset i fysik 2015

The Nobel Prize in Physics 2015



## Nobelpriset i fysik 2015



KUNGL. VETENSKAPS-  
AKADEMIEN  
THE ROYAL SWEDISH ACADEMY OF SCIENCES



**Takaaki Kajita**  
Super-Kamiokande Collaboration  
University of Tokyo, Kashiwa, Japan



**Arthur B. McDonald**  
Sudbury Neutrino Observatory Collaboration  
Queen's University, Kingston, Canada

*"för upptäckten av neutrinooscillationer, som visar att neutriner har massa"*  
*"for the discovery of neutrino oscillations, which shows that neutrinos have mass"*

# solar neutrinos odyssey

- ★ lot of resistance to the idea of solar neutrino experiments, see Bahcall
- ★ then, the indications from solar neutrino experiment has long been ignored. It took 20 years before a test was conducted
  - in part it was distrust of astrophysics and nuclear physics
  - in part this was due to widespread prejudice against neutrino mass
  - acceptance begun when it was realized that small mixing could work

# solar neutrinos odyssey

- ★ lot of resistance to the idea of solar neutrino experiments, see Bahcall
- ★ then, the indications from solar neutrino experiment has long been ignored. It took 20 years before a test was conducted
  - in part it was distrust of astrophysics and nuclear physics
  - in part this was due to widespread prejudice against neutrino mass
  - acceptance begun when it was realized that small mixing could work
- ★ today, oscillation of solar neutrinos is considered an obvious thing
- ★ (even if there are still doubts and things to do)



## where we are now

terrestrial experiments, from reactor to accelerator; tested oscillations

overall consistency of the indications

important role of “global fits” – i.e., of taking seriously the hypothesis of massive neutrinos

also appearance was seen (more on that later)

first hints of “normal spectrum” - aka ordering, aka hierarchy

now, we would like to probe neutrino mass, beginning with KATRIN



Review

### Current unknowns in the three-neutrino framework

F. Capozzi<sup>a</sup>, E. Lisi<sup>b,\*,</sup>, A. Marrone<sup>c, b</sup>, A. Palazzo<sup>c, b</sup>

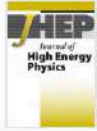
[Show more](#)

<https://doi.org/10.1016/j.pnpnp.2018.05.005>

[Get rights and content](#)

#### Abstract


We present an up-to-date global analysis of data coming from [neutrino oscillation](#) and non-oscillation experiments, as available in April 2018, within the standard framework including three massive and mixed [neutrinos](#). We discuss in detail the status of the three-neutrino ( $3\nu$ ) mass-mixing parameters, both known and unknown. Concerning the latter, we find that: normal ordering (NO) is favored over inverted ordering (IO) at  $3\sigma$  level; the Dirac CP phase is constrained within  $\sim 15\%$  ( $\sim 9\%$ ) uncertainty in NO (IO) around nearly-maximal CP-violating values; the octant of the largest mixing angle and the absolute [neutrino masses](#) remain undetermined. We briefly comment on other unknowns related to theoretical and experimental uncertainties (within  $3\nu$ ) or possible new states and interactions (beyond  $3\nu$ ).



## Updated global analysis of neutrino oscillations in the presence of eV-scale sterile neutrinos

Authors

[Authors and affiliations](#)

Mona Dentler , Álvaro Hernández-Cabezudo, Joachim Kopp, Pedro Machado, Michele Maltoni, Ivan Martínez-Soler, Thomas Schwetz

**Open Access** | Regular Article - Theoretical Physics

First Online: 03 August 2018

49 322 38  
Shares Downloads Citations

### ABSTRACT

We discuss the possibility to explain the anomalies in short-baseline neutrino oscillation experiments in terms of sterile neutrinos. We work in a  $3 + 1$  framework and pay special attention to recent new data from reactor experiments, IceCube and MINOS+. We find that results from the DANSS and NEOS reactor experiments support the sterile neutrino explanation of the reactor anomaly, based on an analysis that relies solely on the relative comparison of measured reactor spectra. Global data from the  $\nu_e$  disappearance channel favour sterile neutrino oscillations at the  $3\sigma$  level with  $\Delta m_{41}^2 \approx 1.3 \text{ eV}^2$  and  $|U_{e4}| = 0.1$ , even without any assumptions on predicted reactor fluxes. In contrast, the anomalies in the  $\nu_e$  appearance channel (dominated by LSND) are in strong tension with improved bounds on  $\nu_\mu$  disappearance, mostly driven by MINOS+ and IceCube. Under the sterile neutrino oscillation hypothesis, the  $p$ -value for those data sets being consistent is less than  $2.6 \times 10^{-6}$ . Therefore, an explanation of the LSND anomaly in terms of sterile neutrino oscillations in the  $3 + 1$  scenario is excluded at the  $4.7\sigma$  level. This result is robust with respect to variations in the analysis and used data, in particular it depends neither on the theoretically predicted reactor neutrino fluxes, nor on constraints from any single experiment. Irrespective of the anomalies, we provide updated constraints on the allowed mixing strengths  $|U_{\alpha 4}|$  ( $\alpha = e, \mu, \tau$ ) of active neutrinos with a fourth neutrino mass state in the eV range.

## more light neutrinos?

various neutrino experiments found anomalous results

individually, many of them could be explained invoking new light neutrinos

however, when one takes this hypothesis seriously, contradictions emerge and this interpretation is not supported by the experiments

the current situation with cosmology is also contradictory

(on top of that, while new light neutrinos might exist, there is no strong theoretical argument in their favor)



# teorie di gauge e decadimento del protone

teorie di gauge, la discussione negli  
anni 70-80, le ricerche del  
decadimento del protone

# under the spell of gauge principle

- ① Yang+Mills 1954,  
*non abelian gauge theories*
- ② Bludman, Leite Lopes 1958,  
*neutral currents interactions*
- ③ Glashow 1961,  
 $SU(2)_L \times U(1)_Y$  and the  $Z^0$  boson
- ④ Anderson, Higgs, Englert+Brout,  
Guralnik+Hagen+Kibble, 1962-67,  
*broken symmetries & gauge bosons mass*
- ⑤ Salam+Ward, Weinberg, 1964-68,  
*a model of leptons*
- ⑥ Glashow+Iliopoulos+Maiani 1970,  
*consistency and the 4<sup>th</sup> quark*
- ⑦ 't Hooft+Veltman, Lee, 1971,  
*renormalizability*
- ⑧ Bouchiat+Iliopoulos+Meyer,  
Gross+Jackiw 1972,  
*anomaly cancellation*

....

Nobel prizes in physics 1979, 1984, 1999, 2008, 2013

# the age of gauge (theories)

APS physics Journals ▾ Physics Magazine Help/Feedback

## PHYSICAL REVIEW LETTERS

Highlights Recent Accepted Collections Authors Referees

### Unity of All Elementary-Particle Forces

Howard Georgi and S. L. Glashow  
Phys. Rev. Lett. **32**, 438 – Published 25 February 1974

Article PDF Export Citation

#### ABSTRACT

Strong, electromagnetic, and weak forces are conjectured to arise from a single fundamental interaction based on the gauge group  $SU(5)$ .

Received 10 January 1974

DOI: <https://doi.org/10.1103/PhysRevLett.32.438>

©1974 American Physical Society

Get Access Share Export Search 5

## Annals of Physics

Volume 93, Issues 1–2, 5 September 1975,  
Pages 193–266

### Unified interactions of leptons and hadrons ☆

Harald Fritzsch, Peter Minkowski

Show more

[https://doi.org/10.1016/0003-4916\(75\)90211-0](https://doi.org/10.1016/0003-4916(75)90211-0) Get rights and content

#### Abstract

It is suggested that a unifying description of leptons and hadrons can be obtained within a nonabelian gauge theory where the gauge group is a symmetry group of a set of massless elementary fermions (leptons, quarks). We investigate the consequences of such an approach for the strong, electromagnetic, and weak interactions. We study both gauge theories with and without fermion number conservation, e.g., theories based on the groups  $SU_n \times SU_n$  ( $n = 8, 12, 16$ ) and  $SO_n$  ( $n = 10, 14$ ).

# from 1979 Nobel lectures

## **Salam:**

That summer [1973, ed] Jogesh Pati and I had predicted proton decay within the context of what is now called GUT.

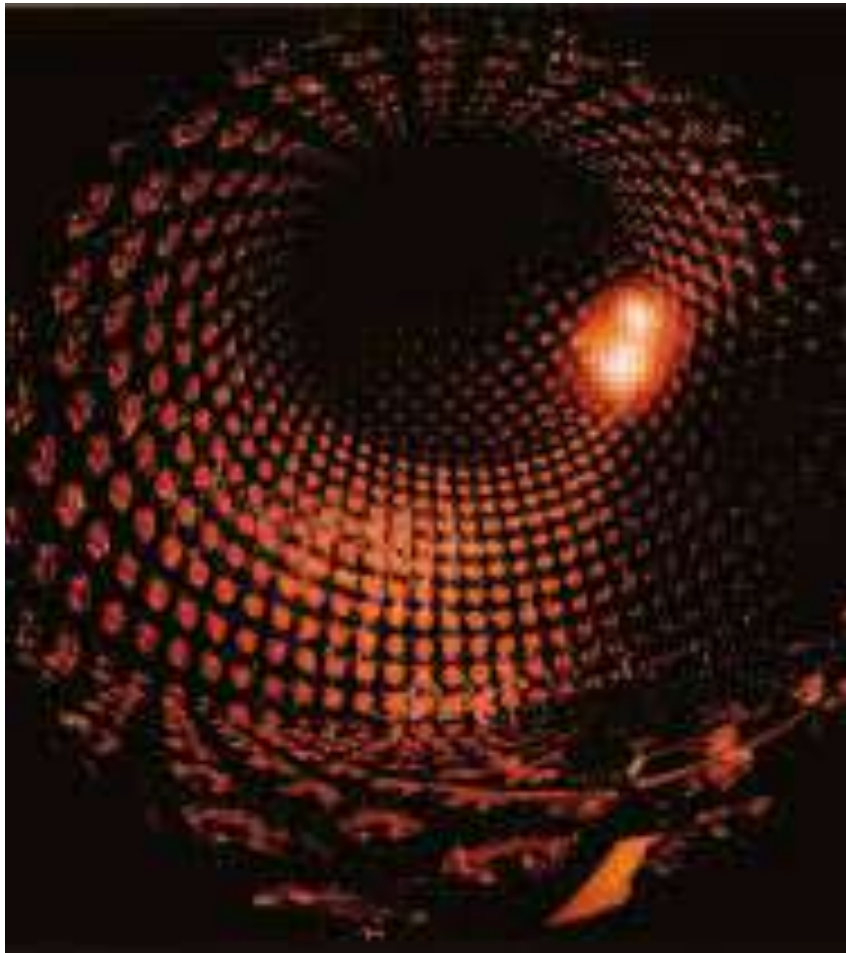
## **Glashow:**

GUT - perhaps along the lines of the original  $SU(5)$  theory of Georgi and me - must be essentially correct. This implies that the proton, and indeed all nuclear matter, must be inherently unstable.

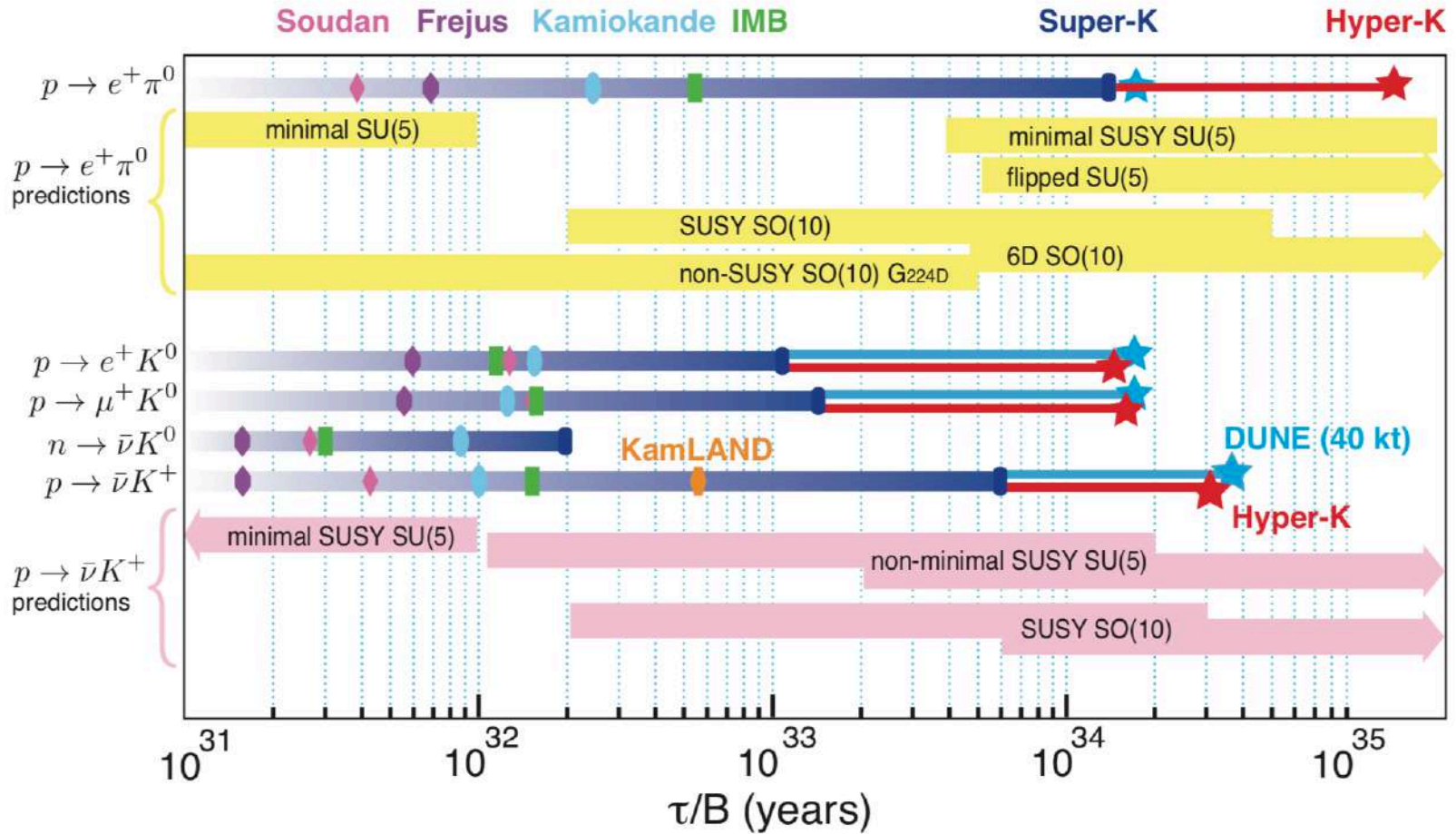
## **Weinberg:**

If effects of a tiny non-conservation of baryon or lepton number such as proton decay or neutrino masses are discovered experimentally, we will then be left with gauge symmetries as the only true internal symmetries of nature, a conclusion that I would regard as most satisfactory.

# KGF, IMB, NUSEX, KAMIOKANDE, ICARUS...

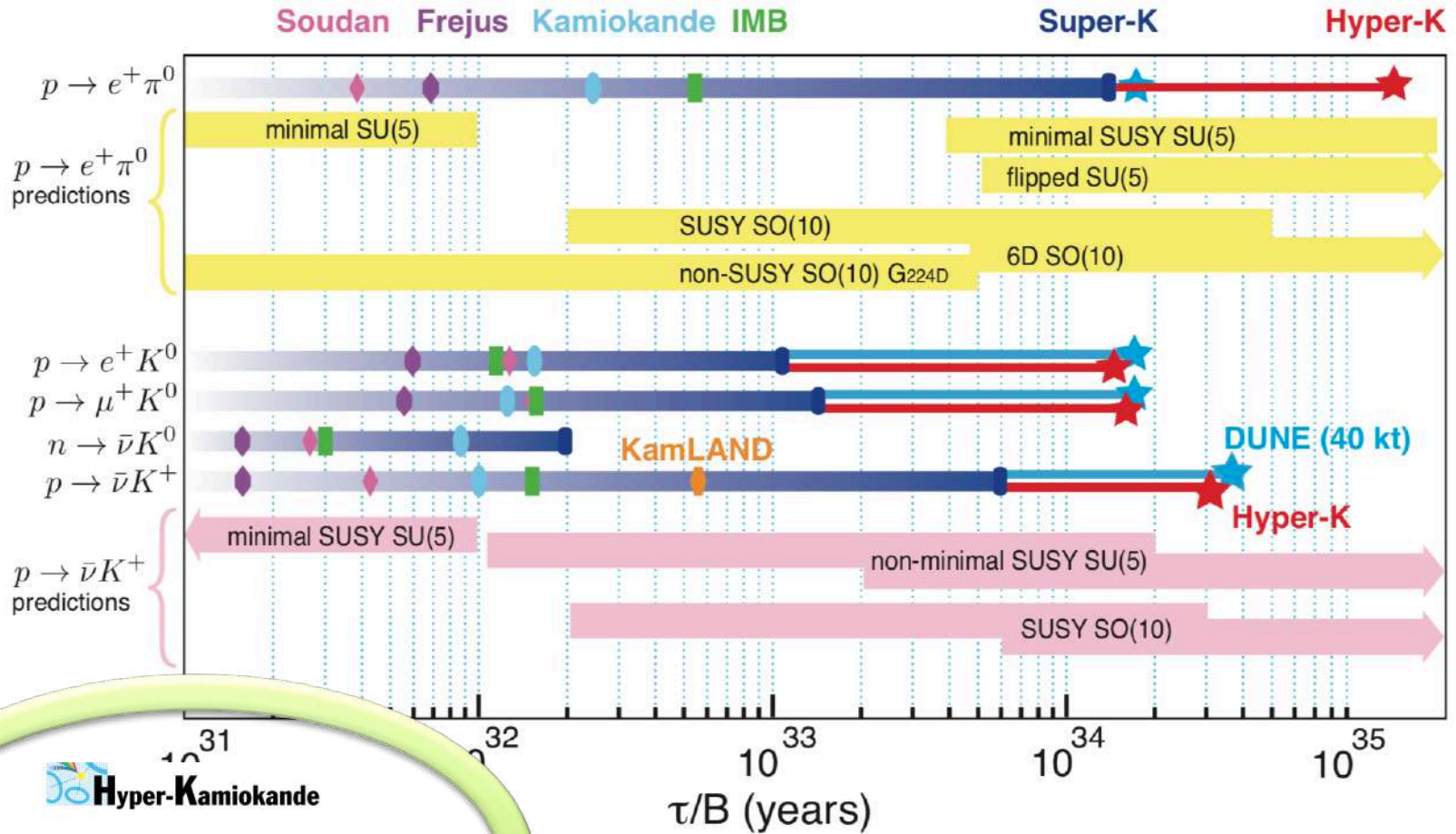


# ... JUNO, DUNE, HYPER-KAMIOKANDE





# ... JUNO, DUNE, HYPER-KAMIOKANDE



**Hyper-Kamiokande**

Design Report  
(Dated: November 30, 2018)

[nessuna referenza alla teoria]



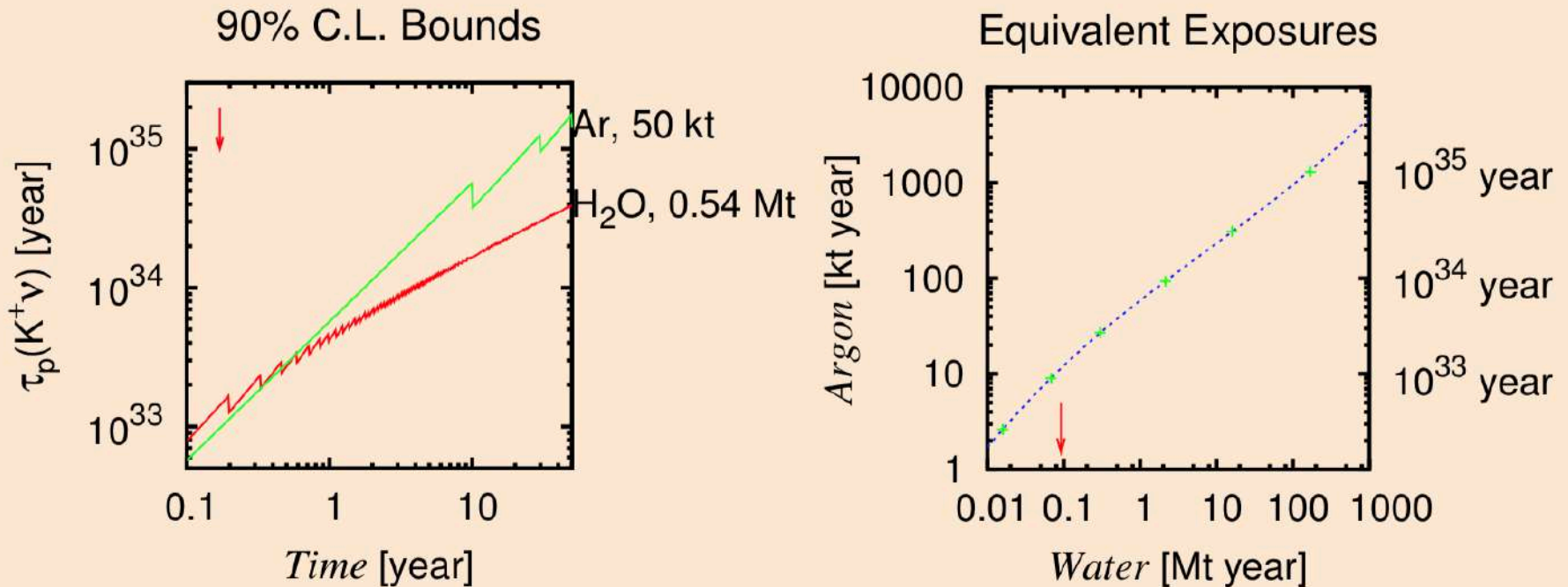
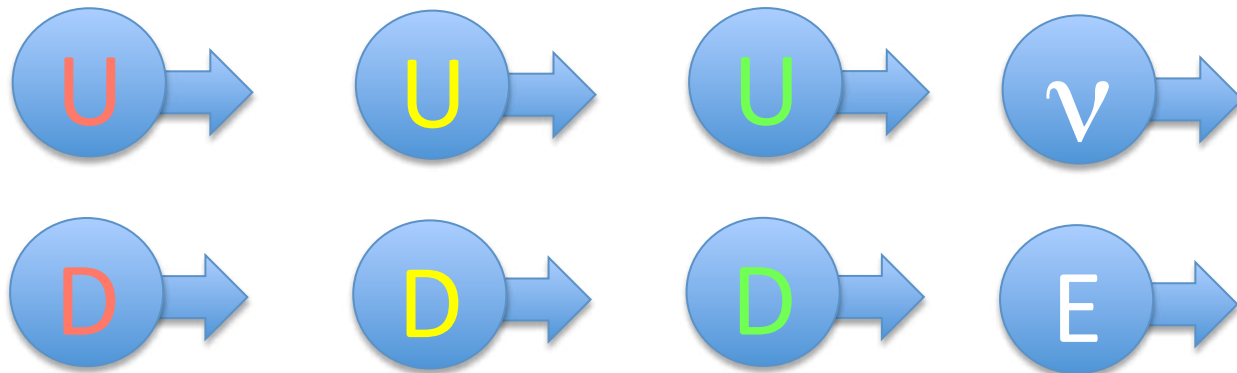
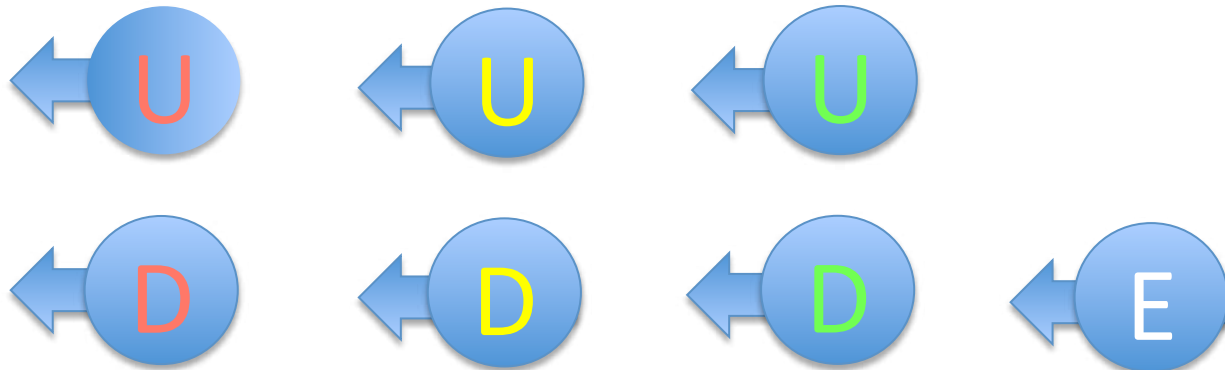


Figure 1: Sensitivity to  $p \rightarrow K\bar{\nu}$ ; syst. not included. Water,  $\epsilon = 14.6\%$  and  $b = 14/(\text{Mton y})$  (2 methods, summed); Argon,  $\epsilon = 97\%$  and  $b = 1/(\text{Mton y})$ . Impact of stat. fluctuations  $\approx 2$ .

# annotazioni teoriche

“manca qualcosa”, sui neutrini RH,  
sulla supersimmetria, sulle energie  
piu' alte

# matter particles in one family



# Glancing beyond SM

- High dim. operators, invariant under SM symmetry, summarize new physics at ultra-high scales
- (They play exactly the same role of Fermi interactions)
- The one with lowest dimension describes **Majorana neutrino masses**
- Oscillations are matched by a huge mass, say, of GUT

$$m_{\text{overall}}^{\nu} \sim \frac{M_W^2}{M_{\text{GUT}}} = 65 \text{ meV} \times \frac{10^{14} \text{ GeV}}{M_{\text{GUT}}}$$

# on right handed neutrinos

- it is plausible that right handed neutrinos exist
- if lighter  $< 10^{14-15}$  GeV, light neutrino masses can be explained
- they can also give reason of baryon asymmetry
- if heavier  $> 10^{7-8}$  GeV, they mean “hierarchy problem”
- (but plenty of similar problems with cosmological constant)
- with them, B-L can be promoted to a gauge symmetry and renewed hopes for grander gauge theory

## **SUPERSYMMETRY, SUPERGRAVITY AND PARTICLE PHYSICS**

**H.P. NILLES\***

*Département de Physique Théorique, Université de Genève, 1211 Genève 4, Switzerland*

and

*CERN, Genève, Switzerland*

Received 16 February 1984

future. Experiments within the next five to ten years will enable us to decide whether supersymmetry as a solution of the naturalness problem of the weak interactions scale is a myth or reality.

This far for the 100 GeV region. The extension of the models to higher energies is of course only

## SUPERSYMMETRY, SUPERGRAVITY AND PARTICLE PHYSICS

H.P. NILLES\*

*Département de Physique Théorique, Université de Genève, 1211 Genève 4, Switzerland*

and

*CERN, Genève, Switzerland*

Received 16 February 1984

future. Experiments within the next five to ten years will enable us to decide whether supersymmetry as a solution of the naturalness problem of the weak interactions scale is a myth or reality.

This far for the 100 GeV region. The extension of the models to higher energies is of course only

**2020-1984=36 y and  $36/(5 \text{ to } 10) \geq 3.6$ , thus  
the CL of the 2<sup>nd</sup> option seems  $> 3.6\sigma$  (=0.03%) 😊**



# quosque tandem, gravity?

consider the mass that realizes

$$\frac{e^2}{r} \equiv G_N \frac{m_s^2}{r}$$

this defines the Stoney mass (1881)

$$\Rightarrow m_s \equiv \sqrt{\frac{e^2}{G_N}} \approx 2\mu g$$

energy at which gravity equates electric force

$$\Rightarrow E_s \equiv m_s \times c^2 \approx 10^{18} \text{ GeV}$$

## **more info** (and a couple of jokes too)

Thomson, Weyl & Yang; on “matter”; on  $\nu$  mass;  
physics at ultrahigh masses; on chosen names; the  
story of V-A

# 1<sup>st</sup> elementary matter particle (Thomson)

Hertz showed, however, that cathode particles possess another property which seemed inconsistent with the idea that they are particles of matter, for he found that they were able to penetrate very thin sheets of metal, e.g.

to the conclusion that the mass of the corpuscle is only about 1/1,700 of that of the hydrogen atom. Thus the atom is not the ultimate limit to the subdivision of matter; we may go further and get to the corpuscle, and at this stage the corpuscle is the same from whatever source it may be derived.

The corpuscle appear to form a part of all kinds of matter under the most diverse conditions; it seems natural therefore to regard it as one of the bricks of which atoms are built up.

## Hermann Weyl

“in our time, there has been unloosed a cataclysm which has swept away space, time, and matter hitherto regarded as the firmest pillars of natural science, but only to make place for a view of things of wider scope, and entailing a deeper vision.”

From *Space Time Matter*, 1918



# C N Yang at the Centennial of MIT, 1961

our present knowledge is sufficient to enable us to say with some certainty that great clarification will come in the field of weak interactions in the next few years. With luck on our side we might even hope to see some integration of the various manifestations of the weak interactions.

Beyond that we are on very uncertain grounds.

- ★ Is the continuum concept of space time extrapolatable to regions of space  $10^{-14}$  cm to  $10^{-17}$  cm, and to regions smaller than  $10^{-17}$  cm?
- ★ What is the unifying basis of the strong, the electromagnetic and the weak interactions?
- ★ What is the role of the gravitational field relative to all these?

If it is difficult to locate singularities of functions by extrapolation, it is as difficult to predict revolutionary changes in physical concepts by forecasting.

## round table “Einstein and the physics of the future” published in *Some Strangeness in the Proportion*, ed. H. Woolf, 1980

### **Weinberg:**

[...] the lifetime of the proton (this has been worked on by a number of people now) comes out to be of the order  $10^{30}$  to  $10^{32}$  years. The present experimental lower bound is  $10^{29}$  years. Thus the time is ripe for an assault on the next few orders of magnitude in the proton lifetime.

### **Dyson:**

[...] the modern view of particle theory, with the sub-nuclear world a playground of interlocking broken and unbroken symmetries, had its roots in Felix Klein's Erlanger Program of 1872 [...]

I predict that in the next 25 years we shall see the emergence of unified physical theories in which general relativity, group theory, and field theory are tied together with bonds of rigorous maths.

### **Yang:**

beautiful mathematics is the language of fundamental physics [...]

Maybe it is my prejudice - maybe it is my ignorance - but I do not believe that any of these graded Lie algebras has the intrinsic and fundamental beauty of Lie algebras and Lie groups, not as yet!

[Main page](#)

[Contents](#)

[Featured content](#)

[Current events](#)

[Random article](#)

[Donate to Wikipedia](#)

[Wikipedia store](#)

Interaction

[Help](#)

[About Wikipedia](#)

[Community portal](#)

[Recent changes](#)

[Contact page](#)

Tools

[What links here](#)

[Related changes](#)

[Upload file](#)

[Special pages](#)

[Permanent link](#)

[Page information](#)

[Wikidata item](#)

[Cite this page](#)

Print/export

[Create a book](#)

[Download as PDF](#)

# Matter

From Wikipedia, the free encyclopedia

*This article is about the concept in the physical sciences. For other uses, see [Matter \(disambiguation\)](#).*

In [classical physics](#) and general chemistry, **matter** is any substance that has [mass](#) and takes up space by being ultimately composed of [atoms](#), which are made up of interacting [subatomic particles](#), and in everyday use we think of matter as being made up of them, and any particles (or [combination of particles](#)) that act as if they have both [rest mass](#) and [charge](#), or other energy phenomena or waves such as [light](#) or [sound](#).<sup>[1][2]</sup> Matter exists in various [states](#) (all of which are made up of matter) such as [solid](#), [liquid](#), and [gas](#) – for example [water](#) exists as ice, liquid water, and gaseous steam – but other states include [plasma](#), [Bose–Einstein condensate](#), [fermionic condensates](#), and [quark–gluon plasma](#).<sup>[3]</sup>

Usually atoms can be imagined as a [nucleus](#) of [protons](#) and [neutrons](#), and a surrounding "cloud" of orbiting [electrons](#). This is not strictly correct, because subatomic particles and their properties are governed by their [quantum nature](#), which means they behave more like [waves as well as particles](#) and they do not have well-defined sizes or positions. In the [Standard Model of particle physics](#), the [elementary constituents](#) of atoms are [quantum entities](#) which do not have an inherent "size" or "volume" in a classical sense. In other [fundamental interactions](#), some "point particles" known as [fermions](#) ([quarks](#), [leptons](#)), and many composite particles exist under everyday conditions; this creates the property of matter which appears to us as matter taking up space.

For much of the history of the [natural sciences](#) people have contemplated the exact nature of matter. The idea of [particulate theory of matter](#), was first put forward by the Greek philosophers [Leucippus](#) (~490 BC) and [Democritus](#).

## Contents [hide]

- [Comparison with mass](#)
- [Definition](#)
  - [Based on atoms](#)
  - [Based on protons, neutrons and electrons](#)
  - [Based on quarks and leptons](#)
  - [Based on elementary fermions \(mass, volume, and space\)](#)



# Matter

From Wikipedia, the free encyclopedia

*This article is about the concept in the physical sciences. For other uses, see [Matter \(disambiguation\)](#).*

In [classical physics](#) and general chemistry, **matter** is any substance that has [mass](#) and takes up space by being ultimately composed of [atoms](#), which are made up of interacting [subatomic particles](#), and in everyday use we think of matter as being made up of them, and any particles (or [combination of particles](#)) that act as if they have both [rest mass](#) and [momentum](#), or other energy phenomena or waves such as [light](#) or [sound](#).<sup>[1][2]</sup> Matter exists in various [states](#) (all as [solid](#), [liquid](#), and [gas](#) – for example [water](#) exists as ice, liquid water, and gaseous steam – but other states include [plasma](#), [fermionic condensates](#), and [quark–gluon plasma](#).<sup>[3]</sup>

Usually atoms can be imagined as a [nucleus](#) of [protons](#) and [neutrons](#), and a surrounding "cloud" of orbiting [electrons](#), but this is not strictly correct, because subatomic particles and their properties are governed by their [quantum nature](#), which means they behave more like [waves as well as particles](#) and they do not have well-defined sizes or positions. In the [Standard Model of particle physics](#), the [elementary constituents](#) of atoms are [quantum entities](#) which do not have an inherent "size" or "volume" in a classical sense. In other [fundamental interactions](#), some "point particles" known as [fermions](#) ([quarks](#), [leptons](#)), and many composite particles exist under everyday conditions; this creates the property of matter which appears to us as matter taking up space.

For much of the history of the [natural sciences](#) people have contemplated the exact nature of matter. The idea of a [particulate theory of matter](#), was first put forward by the Greek philosophers [Leucippus](#) (~490 BC) and [Democritus](#).

## Comparison with mass

### 2 Definition

- 2.1 Based on atoms
- 2.2 Based on protons, neutrons and electrons
- 2.3 Based on quarks and leptons
- 2.4 Based on elementary fermions (mass, volume, and space)

- ⊙ In ordinary language we say:  
*matter does not disappear*
- ⊙ In this panel, the main reaction that allows the Sun to work, that means that *neutrinos are matter particles* (leptons)
- ⊙ Known physics says: *numbers of baryons and leptons do not change*
- ⊙ Only B-L exactly conserved (SM)

**proton fusion:  $p+p \rightarrow D+e^++\nu$**

Electric charge is conserved:  
 $1+1=1+1+0$

Baryon number is conserved:  
 $1+1=2+0+0$

Lepton number is conserved:  
 $0+0=0-1+1$

$$(A, Z) \rightarrow (A, Z+1) + e^- + \bar{\nu}_e \quad \text{All OK}$$

$$(A, Z) \rightarrow (A, Z+2) + 2 e^- \quad \text{L changes}$$

$$n \rightarrow e^- + p \quad \text{L and spin change}$$

$$p \rightarrow e^+ + \pi^0 \quad \text{B and L change}$$

$$n \rightarrow e^- + K^+ \quad \text{B and L change}$$

$$e^- \rightarrow \gamma + \nu_e \quad \text{Charge changes}$$

$$n \rightarrow \bar{n} \quad \text{B changes}$$

# direct search of big-bang neutrinos

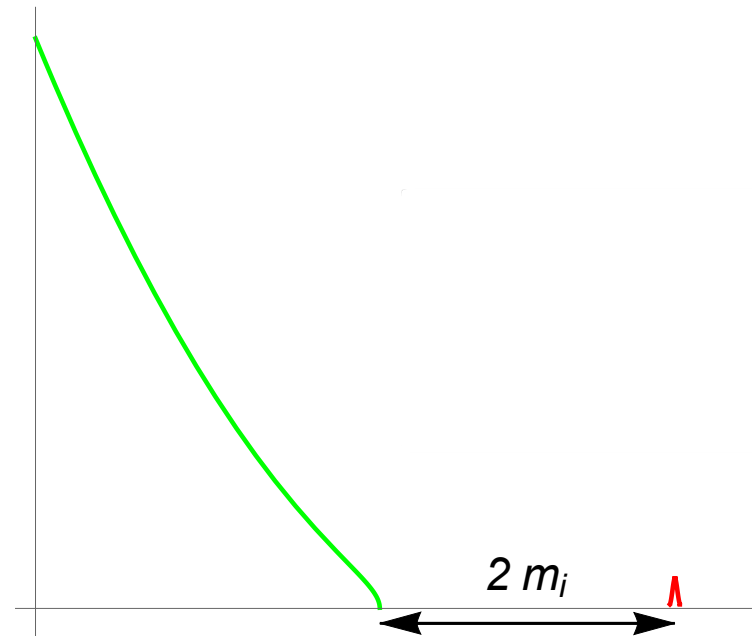
big-bang neutrinos produce 3 **neutrino-capture** lines for a radioactive target

their positions depend on  $m_i$ ; their intensity on  $|U_{ei}|^2$

lightest neutrino gives the most intense line for normal hierarchy

Needs

- great energy resolution
- big target mass,  $\geq 100\text{g}$  of tritium



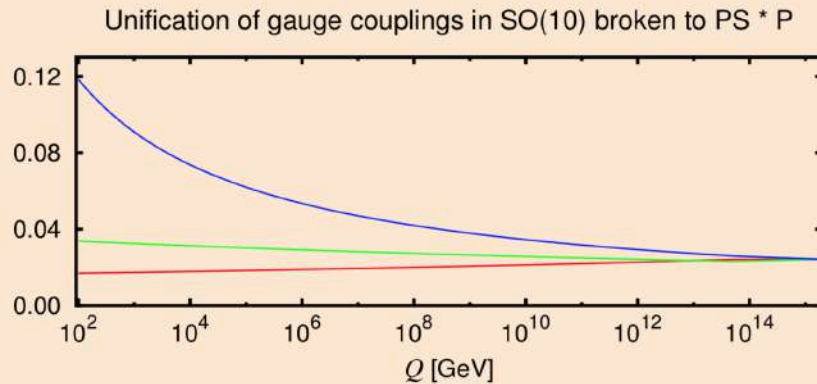


Figure 2: Evolution of the gauge coupling constants in a GUT model with intermediate scale. Here,  $M_{\text{interm.}} \approx 5 \times 10^{13}$  GeV.

Bajc et al 2005; Bertolini et al 2009-2011; Jshipura et al 2011; Bucciella et al 2012; Dueck et al 2013; Altarelli et al 2013; Ohlsson et al 2019

## minimal SO(10)

(principled model)

- ★ 16-plet coupled to 10 and 126 higgs: heavy right-handed neutrinos
- ★ (Peccei Quinn symmetry to address strong CP and dark matter)



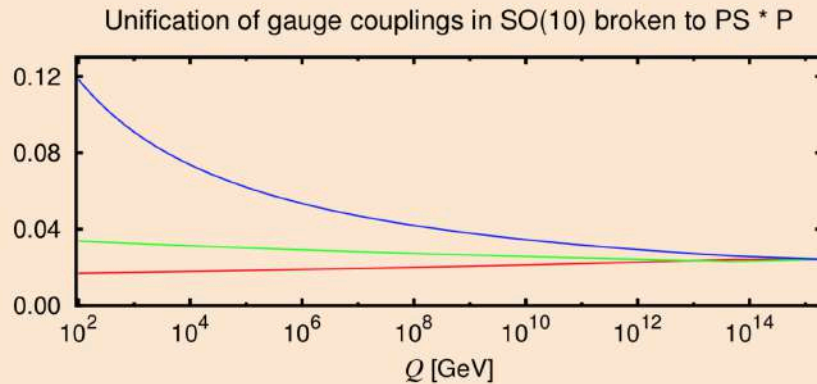


Figure 2: Evolution of the gauge coupling constants in a GUT model with intermediate scale. Here,  $M_{\text{interm.}} \approx 5 \times 10^{13}$  GeV.

Bajc et al 2005; Bertolini et al 2009-2011; Jshipura et al 2011; Bucciella et al 2012; Dueck et al 2013; Altarelli et al 2013; Ohlsson et al 2019

## minimal SO(10)

(principled model)

- ★ 16-plet coupled to 10 and 126 higgs: heavy right-handed neutrinos
- ★ (Peccei Quinn symmetry to address strong CP and dark matter)
- ★ neutrinos are massive and fermion masses constrained

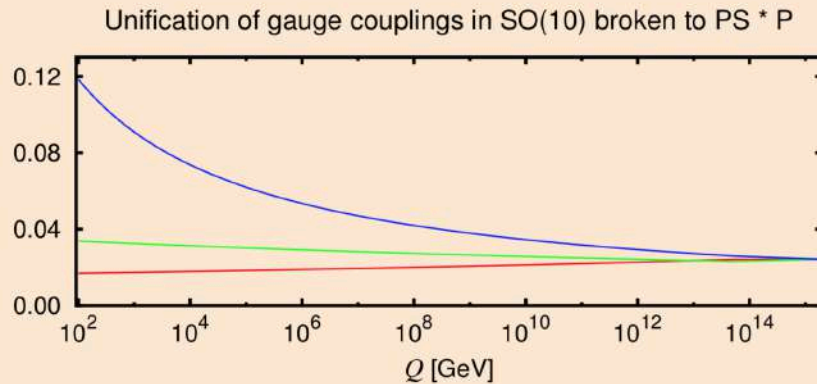


Figure 2: Evolution of the gauge coupling constants in a GUT model with intermediate scale. Here,  $M_{\text{interm.}} \approx 5 \times 10^{13}$  GeV.

Bajc et al 2005; Bertolini et al 2009-2011; Jshipura et al 2011; Bucciella et al 2012; Dueck et al 2013; Altarelli et al 2013; Ohlsson et al 2019

## minimal SO(10)

(principled model)

- ★ 16-plet coupled to 10 and 126 higgs: heavy right-handed neutrinos
- ★ (Peccei Quinn symmetry to address strong CP and dark matter)
- ★ neutrinos are massive and fermion masses constrained
- ★ normal mass hierarchy;  $m_{\beta\beta}$  in the few meV range



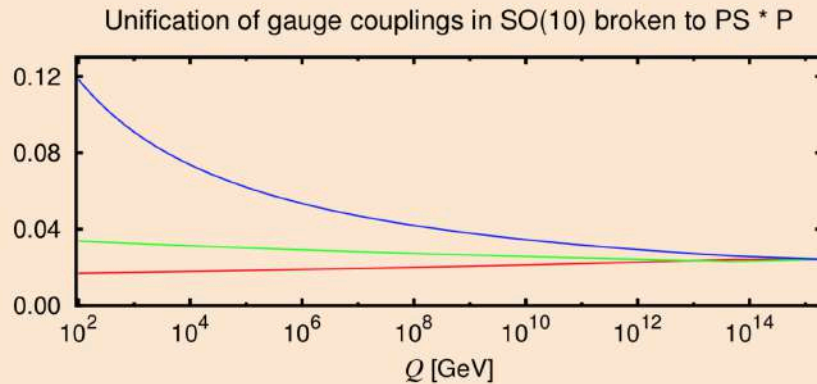


Figure 2: Evolution of the gauge coupling constants in a GUT model with intermediate scale. Here,  $M_{\text{interm.}} \approx 5 \times 10^{13}$  GeV.

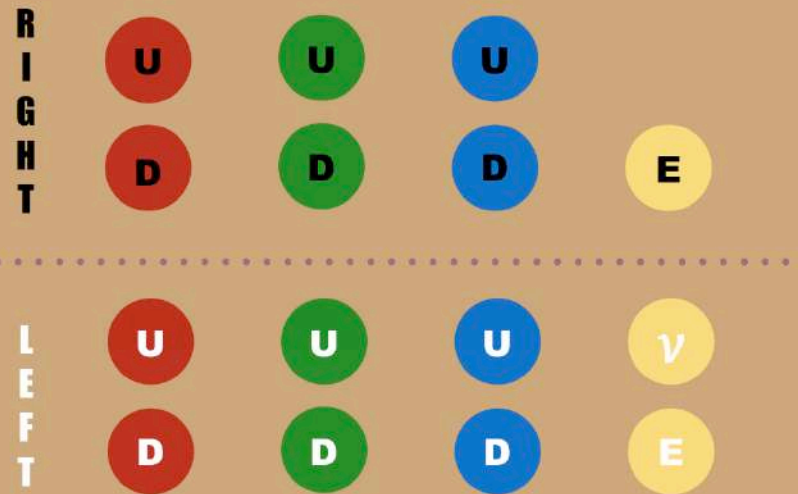
Bajc et al 2005; Bertolini et al 2009-2011; Joshipura et al 2011; Bucciella et al 2012; Dueck et al 2013; Altarelli et al 2013; Ohlsson et al 2019

## minimal SO(10)

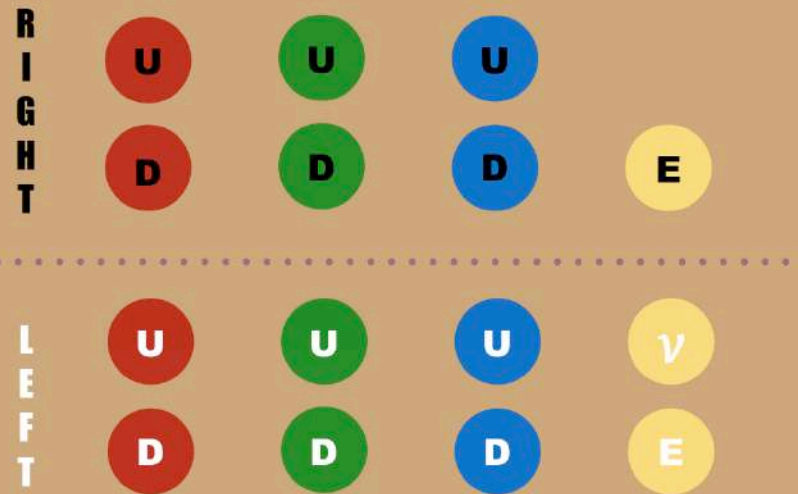
(principled model)

- ★ 16-plet coupled to 10 and 126 higgs: heavy right-handed neutrinos
- ★ (Peccei Quinn symmetry to address strong CP and dark matter)
- ★ neutrinos are massive and fermion masses constrained
- ★ normal mass hierarchy;  $m_{\nu\beta\beta}$  in the few meV range
- ★ (potentially interesting expectations for proton decay)

# 15 particles per family

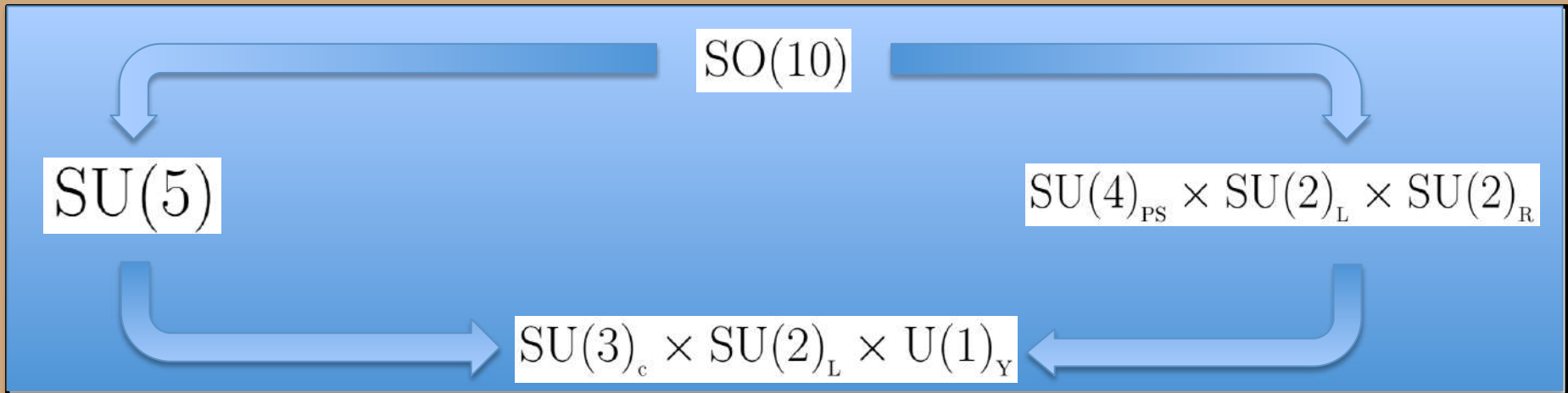
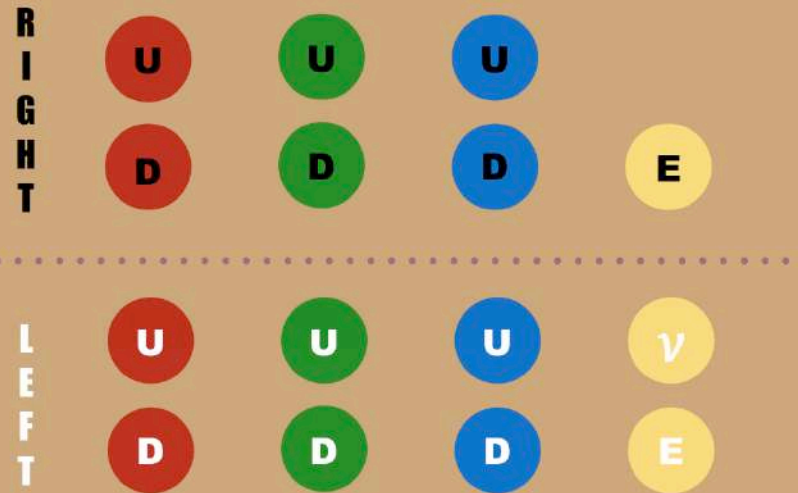


# 15 particles per family

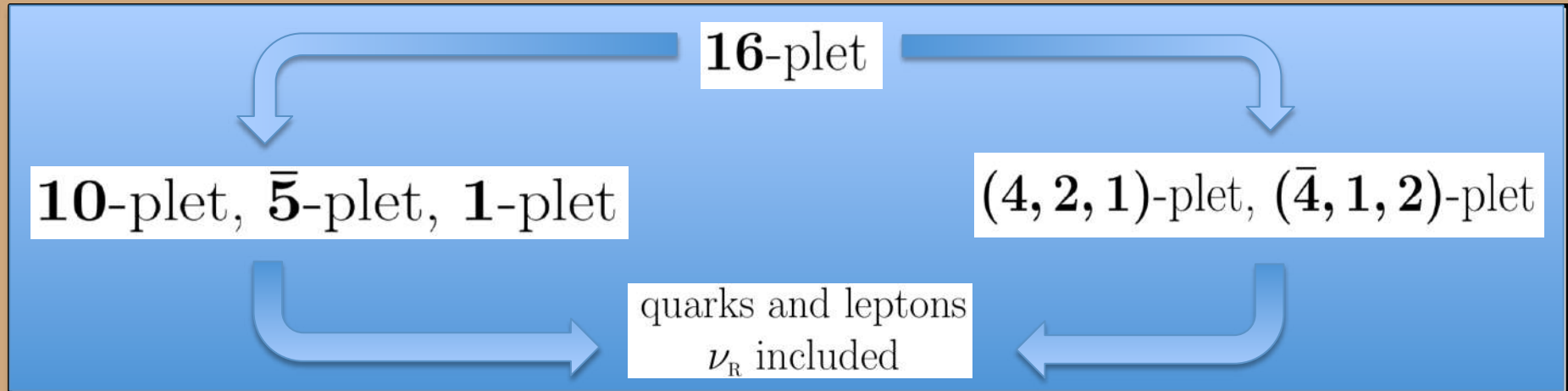
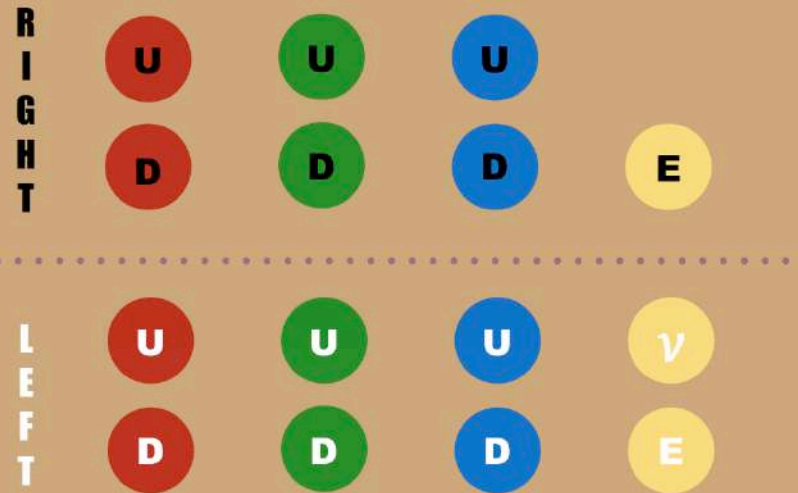


$SU(2)_L$  acts on  $\begin{pmatrix} \nu_L \\ e_L \end{pmatrix}$  while  $SU(2)_R$  acts on  $\begin{pmatrix} \nu_R \\ e_R \end{pmatrix}$

# 15 particles per family



# 15 particles per family



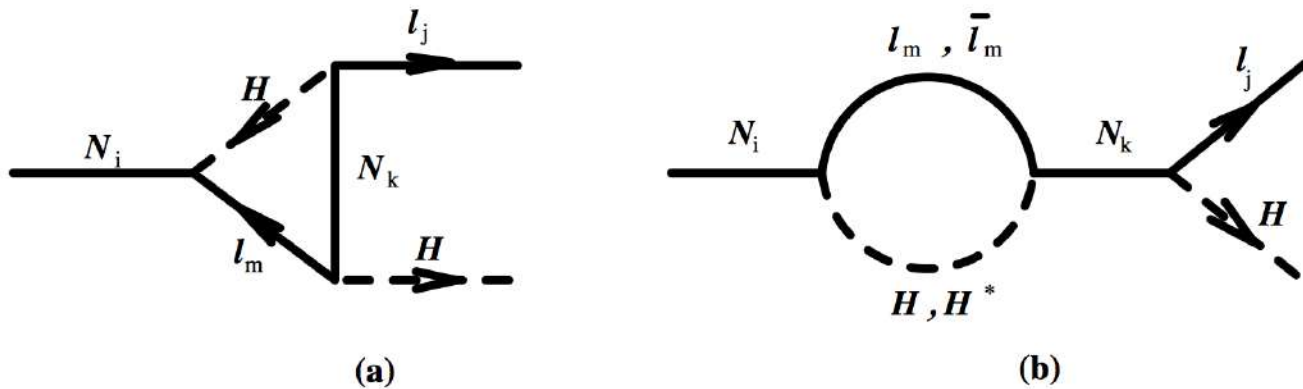


Figure 1: Diagrams contributing to the vertex (Fig. 1a) and wave function (Fig. 1b) CP violation in the heavy singlet neutrino decay.

Covi et al. '96

# Heavy RH neutrinos

can prefer to decay into antileptons; then, converted into baryons by the SM effects described previously

## Do experiments suggest a hierarchy problem?

Francesco Vissani

*International Centre for Theoretical Physics, Strada Costiera 11, I-34013 Trieste, Italy*

(Received 18 September 1997; published 14 April 1998)

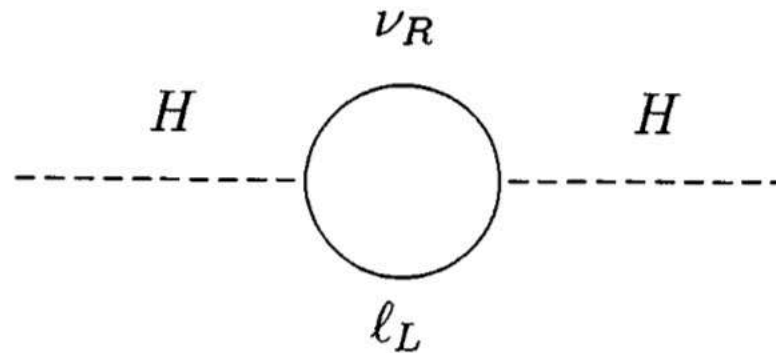


FIG. 1. The Feynman diagram originating the corrections in Eq. (1);  $\nu_R$  denotes the right-handed neutrino of mass  $M_R$ ,  $\ell_L = (\nu_L, e_L)$  the leptonic and  $H$  the Higgs doublets.

*OK, but what about the cosmological constant?*



the standard model. So (if a joke is permitted) we present a prediction for LHC:

$$m_H = 135 \pm 5 \text{ GeV} \tag{14}$$

the reason is that this value will increase the entropy in the minds of several theorists. Note, however, that the decay of a standard and supersymmetric HIGGS particles with the same mass (or also the production rate–“cross-sections”) could be rather different; thence, these measurements would offer a possibility to distinguish between the SM and its supersymmetric extension even in this tricky case.

### 3 (Not quite a) conclusion

We would like to close this pages by spending few words of caution, to remind that failures of the standard model have been often claimed in the past years (today, several of them are considered dubious or simply wrong tracks). Here is an arbitrary selection:

| THEORETICAL<br>INTERPRETATION | EXPERIMENTAL<br>ANOMALY                |
|-------------------------------|----------------------------------------|
| leptoquark .....              | High $x$ and $Q^2$ events at HERA      |
| compositeness .....           | Excess of 4-jet events at ALEPH        |
| light gravitino .....         | $ee\gamma\cancel{E}$ event at CDF      |
| 17 keV neutrino .....         | bump in $\beta$ spectra (SIMPSON, ...) |
| monopole .....                | induced currents (CABRERA)             |
| proton decay .....            | contained multitrack events at KGF     |
| ...                           | ...                                    |

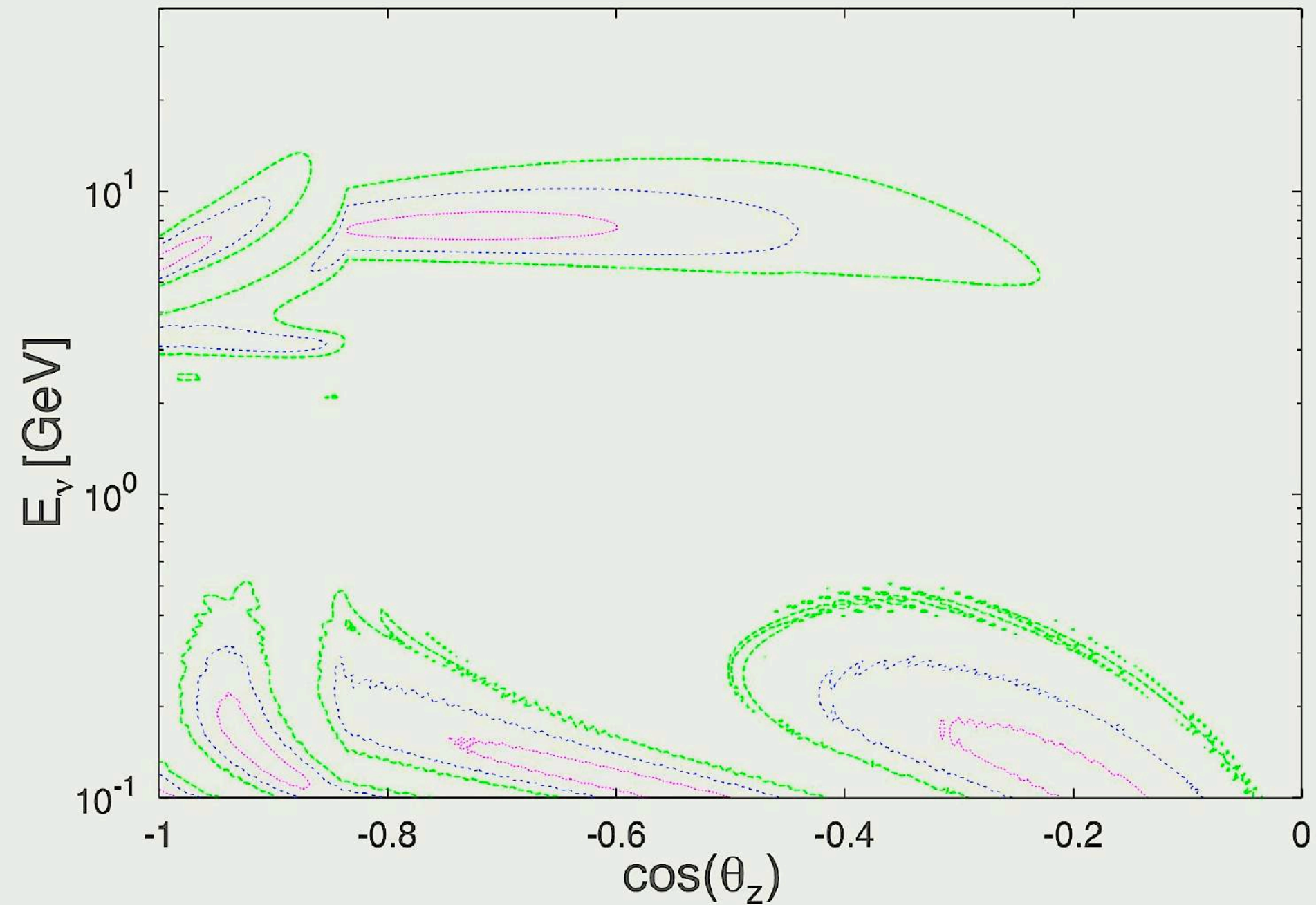
Is there any moral behind these stories? Maybe not; however:

- 1) they suggest to go slowly and carefully from data to theories and back (because of possible pitfalls of interpretation, of suggestion, *etc.*);
- 2) they witness how hard is to reach the frontiers of standard model; and, also, how strong is the desire of particle physicists to find them!

**NH** → **NO**

***Normal hierarchy*** → ***Normal ordering***

$P_{ee}=0.7, 0.5, 0.3$  through the Earth (La Thuile 2003)



**NO** → **YES**

***Normal ordering*** → ***Yearningly Expected Spectrum***

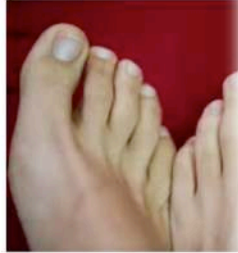
webbed feet hallux rigidus foot toe bro infection broken toe dal shabet beyond naturalness naturalness theory meaning food dict



Why are my toes red? Causes, medicalnewstoday.com



toe - Wiktionary en.wiktionary.org



Toe Pictures I Download From unsplash.com



Naturalness as a success factor phys.org



Naturalness.com is for sale squadhelp.com

ball jute twine micro loyalty bruno banani thongs



Emely - String I Strings I brunobanani.com



Tapered End 0.95" Strandbergguitars.com

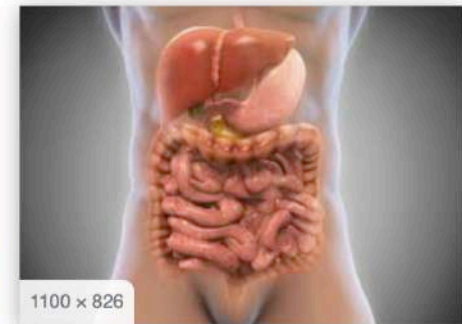


COSABELLA SWEET cosabella.com

leaky microbiota bacteria leaky gut syndrome gut brain axis



gut health - latest news, breaking stories a... independent.co.uk



1100 x 826

How gut bacteria may help to spot and add... medicalnewstoday.com

# universal V-A

## **Ruderman+Finkelstein 1949**

Predictions of  $R(\pi \rightarrow e + \nu)/R(\pi \rightarrow \mu + \nu)$  in various hypotheses

## **Durbin+Loar+Steinberger 1951**

pion parity from deuterium photodissociation

## **Lokanathan+Steinberger 1955 & Anderson+Lattes 1957**

apparently  $R(\pi \rightarrow e + \nu)$  is just absent, ruling out V-A ☹️

## **Sudarshan+Marshak 1957 & Feynman+Gell-Mann 1958**

theory first! V-A implies that previous result is inaccurate

## **Fazzini et al. 1958**

$R(\pi \rightarrow e + \nu)/R(\pi \rightarrow \mu + \nu)$  confirms V-A 😊