

July 14-19, 2017

*Recent developments
in
shell model studies of atomic nuclei*

Takaharu Otsuka



This work has been supported by MEXT and JICFuS as a priority issue (Elucidation of the fundamental laws and evolution of the universe) to be tackled by using Post ‘K’ Computer

From the shell model perspectives,
three (possible) pillars combined for future

computation

Monte Carlo
Shell Model
(MCSM)

(almost)
unlimited
dimensionality

massive
parallel
computers

Hamiltonian

pf
pfg9d5 (A3DA) (Ni)
8+8 on ^{56}Ni core (Zr)
8+8 on ^{80}Zr core (Sn)
8+10 on ^{132}Sn core
(Sm)

...
island of stability
+
 χ EFT based
(multi-)shell int.

many-body dynamics

Shell evolution
(Type I & II)

Quantum Phase
Transition

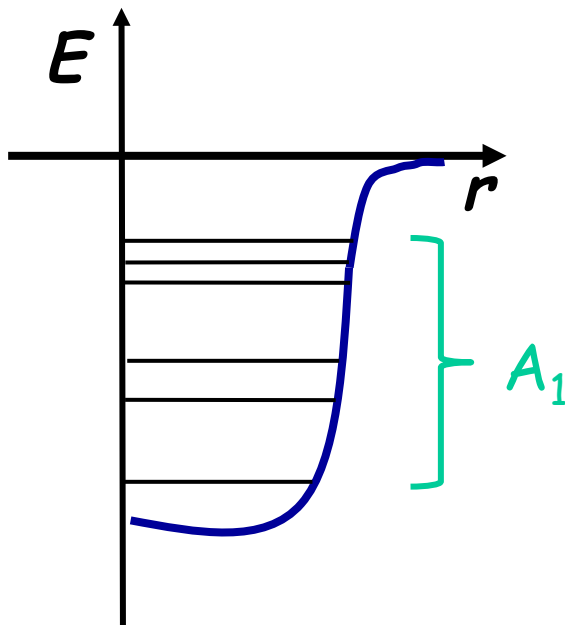
Shape coexistence

Quantum
Self-organization

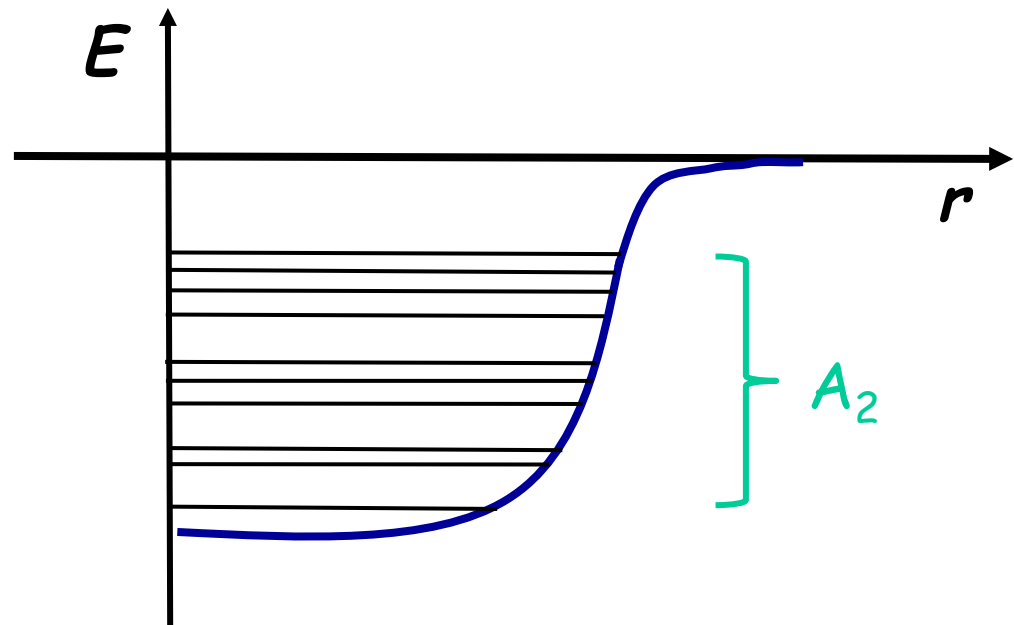
Single-particle states - starting point -

Mean potential becomes wider so as to cast A nucleons with the same separation energy, keeping its depth.

light nuclei



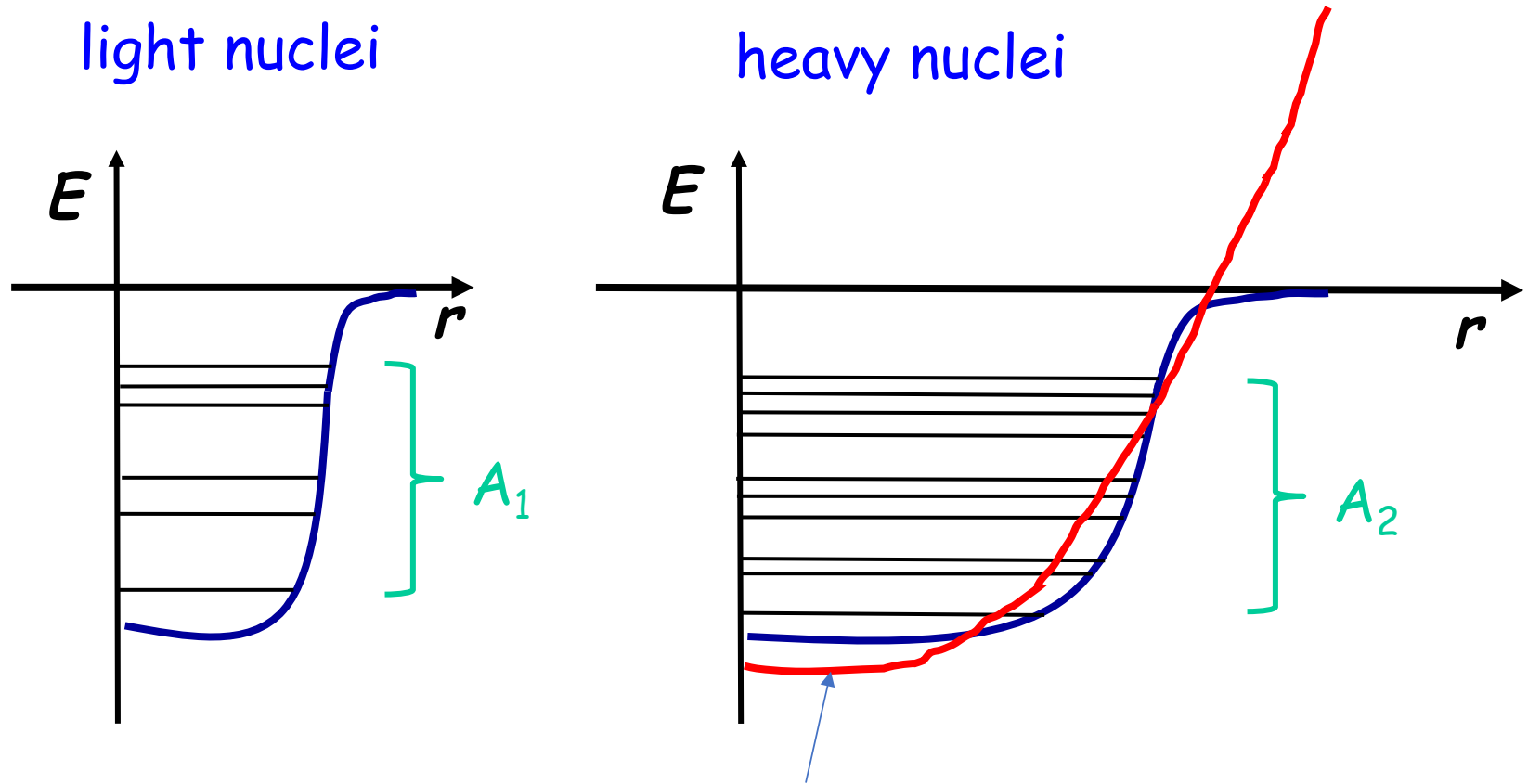
heavy nuclei



This potential can be approximated by a Harmonic Oscillator potential.

Single-particle states - starting point -

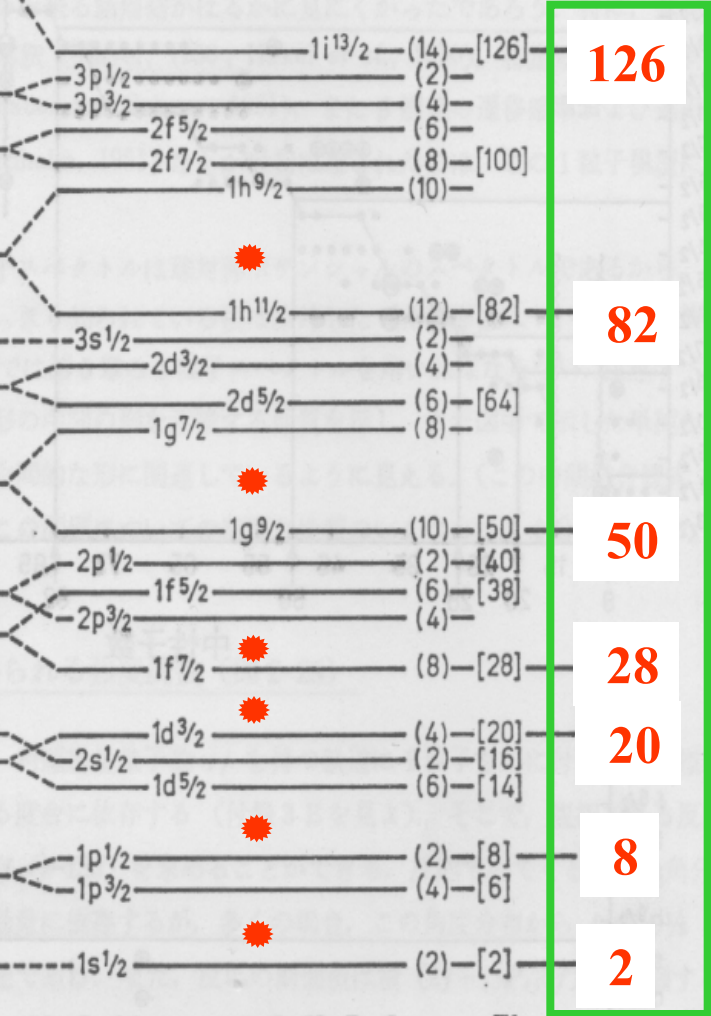
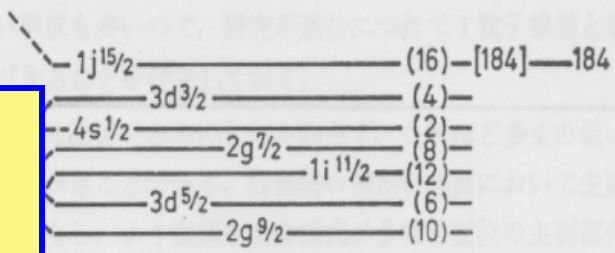
Mean potential becomes wider so as to cast A nucleons with the same separation energy, keeping its depth.



This potential can be approximated by a Harmonic Oscillator potential.

Eigenvalues of
HO potential

5ħω
4ħω
3ħω
2ħω
1ħω
0



Magic numbers
by
Mayer and
Jensen (1949)

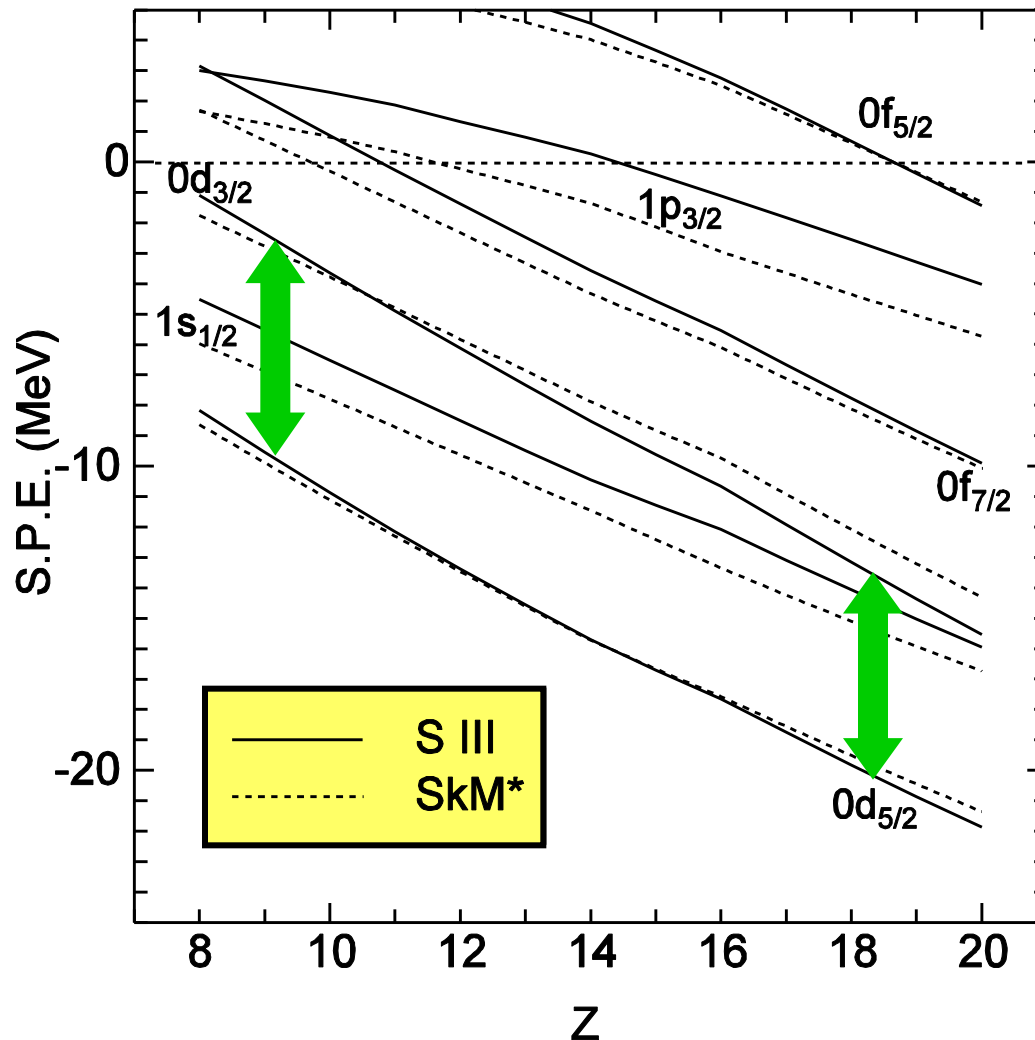


R SHELL MODEL

図 2-23 1 粒子軌道の順序. 図は M. G. Mayer and J. H. D. Jensen, *Elementary Theory of Nuclear Shell Structure*, p. 58, Wiley, New York, 1955 からとった.

Realization in Hartree-Fock energies by Skyrme model

Neutron Single-Particle Energies at $N=20$



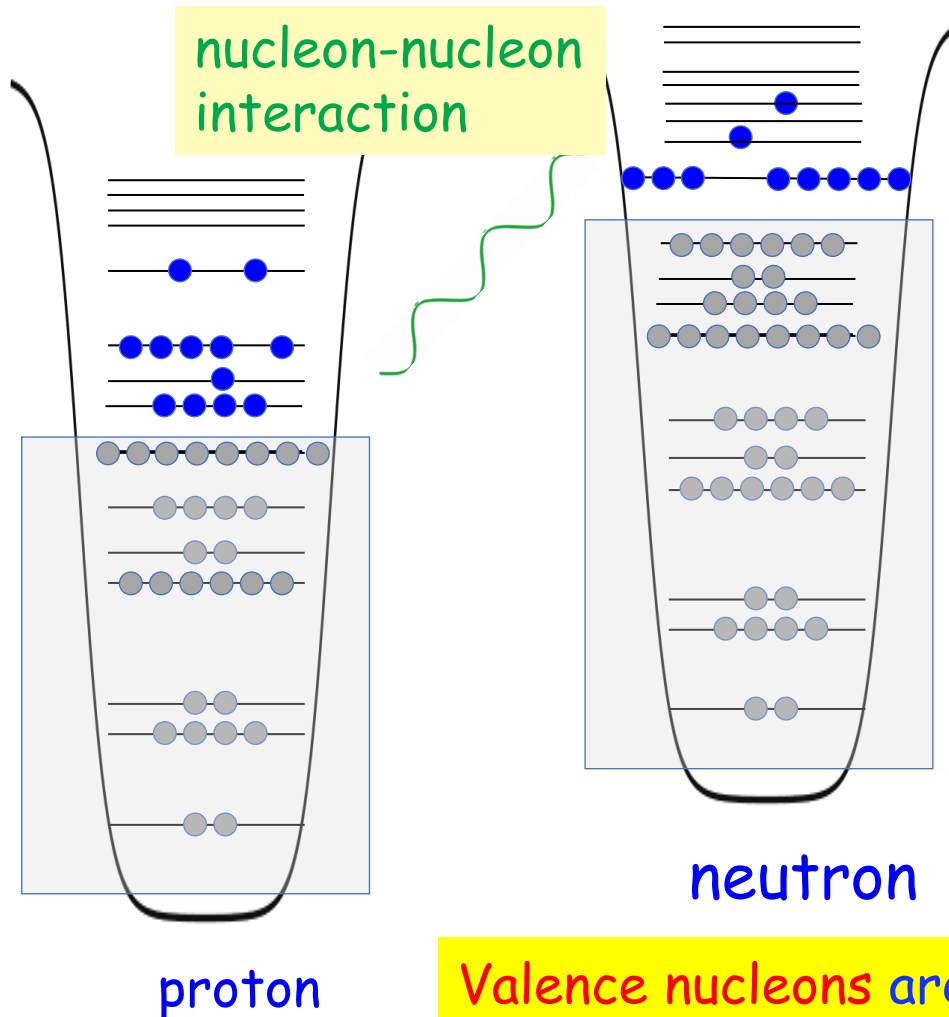
The shell structure remain rather unchanged

-- orbitals shifting together

-- change of potential depth

~ Woods-Saxon.

shell structure and nucleon-nucleon interaction



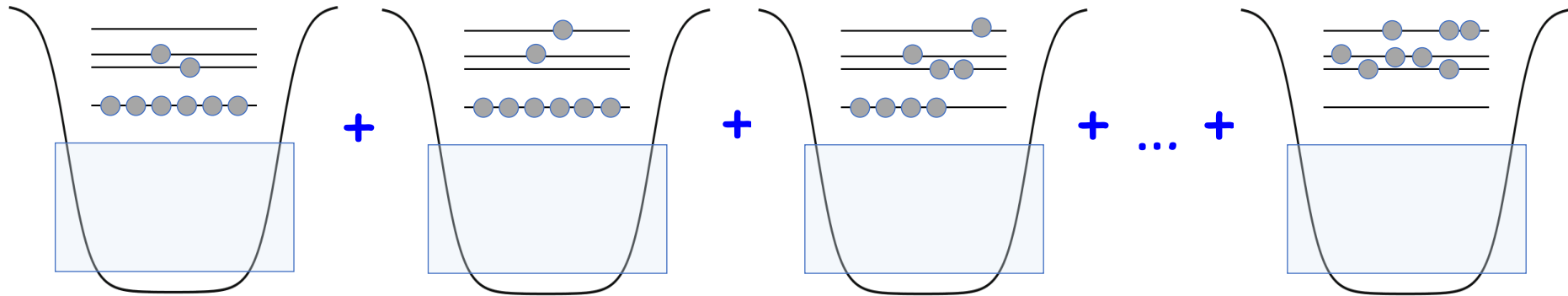
Protons and neutrons are orbiting in the mean potential like a "vase"
→ single-particle energies

Lower orbits form the **inert core** (or closed shell)
(*shaded parts in the figure*)

Upper orbits are only partially occupied
(**valence orbits and nucleons**).

Valence nucleons are the major source of nuclear dynamics at low excitation energy, because the **inert core** is frozen (*implicitly taken into account in terms of effective interaction and operators*).

Possible configurations : **dimension** of the shell-model calculation



How can we find the solution of this problem ?

Hamiltonian in shell model calculations

ϵ_i : Single Particle Energy (SPE)

$\langle j1, j2, J, T | V | j3, j4, J, T \rangle$

: Two-Body Matrix Element (TBME)

$$H = \sum_i \epsilon_i n_i + \sum_{i,j,k,l} v_{ij,kl} a_i^\dagger a_j^\dagger a_l a_k$$

Step 1: Calculate matrix elements

$$\langle \phi_1 | H | \phi_1 \rangle, \quad \langle \phi_1 | H | \phi_2 \rangle, \quad \langle \phi_1 | H | \phi_3 \rangle, \dots$$

where ϕ_1 , ϕ_2 , ϕ_3 are Slater determinants

In the second quantization,

$$\phi_1 = a_{\alpha}^{\dagger} a_{\beta}^{\dagger} a_{\gamma}^{\dagger} \dots \boxed{|0\rangle}$$

$$\phi_2 = a_{\alpha'}^{\dagger} a_{\beta'}^{\dagger} a_{\gamma'}^{\dagger} \dots \boxed{|0\rangle}$$

$$\phi_3 = \dots$$

closed shell



$$H = \sum_i \epsilon_i n_i + \sum_{i,j,k,l} v_{ij,kl} a_i^{\dagger} a_j^{\dagger} a_l a_k$$

Step 2 : Obtain the matrix of Hamiltonian, **H**

$$\mathbf{H} = \begin{bmatrix} \langle \phi_1 | \mathbf{H} | \phi_1 \rangle & \langle \phi_1 | \mathbf{H} | \phi_2 \rangle & \langle \phi_1 | \mathbf{H} | \phi_3 \rangle & \dots \\ \langle \phi_2 | \mathbf{H} | \phi_1 \rangle & \langle \phi_2 | \mathbf{H} | \phi_2 \rangle & \langle \phi_2 | \mathbf{H} | \phi_3 \rangle & \dots \\ \langle \phi_3 | \mathbf{H} | \phi_1 \rangle & \langle \phi_3 | \mathbf{H} | \phi_2 \rangle & \langle \phi_3 | \mathbf{H} | \phi_3 \rangle & \dots \\ \langle \phi_4 | \mathbf{H} | \phi_1 \rangle & \cdot & \cdot & \cdot \\ \cdot & \cdot & \cdot & \cdot \end{bmatrix}$$

Step 3: Solve the eigenvalue problem : $\mathbf{H} \Psi = \mathbf{E} \Psi$

$$\begin{pmatrix}
 \langle \phi_1 | \mathbf{H} | \phi_1 \rangle & \langle \phi_1 | \mathbf{H} | \phi_2 \rangle & \cdot & \cdot & \cdot \\
 \langle \phi_2 | \mathbf{H} | \phi_1 \rangle & \langle \phi_2 | \mathbf{H} | \phi_2 \rangle & \cdot & \cdot & \cdot \\
 \langle \phi_3 | \mathbf{H} | \phi_1 \rangle & \cdot & \cdot & \cdot & \cdot \\
 \langle \phi_4 | \mathbf{H} | \phi_1 \rangle & \cdot & \cdot & \cdot & \cdot \\
 \cdot & \cdot & \cdot & \cdot & \cdot
 \end{pmatrix}
 \begin{pmatrix}
 c_1 \\
 c_2 \\
 c_3 \\
 c_4 \\
 \cdot \\
 \cdot
 \end{pmatrix}
 = \mathbf{E}
 \begin{pmatrix}
 c_1 \\
 c_2 \\
 c_3 \\
 c_4 \\
 \cdot \\
 \cdot
 \end{pmatrix}$$

$$\Psi = c_1 \phi_1 + c_2 \phi_2 + c_3 \phi_3 + \dots$$

c_i probability amplitudes

With Slater determinants $\phi_1, \phi_2, \phi_3, \dots$,
the eigen wave function is expanded as

$$\Psi = c_1 \phi_1 + c_2 \phi_2 + c_3 \phi_3 + \dots$$

c_i probability amplitudes

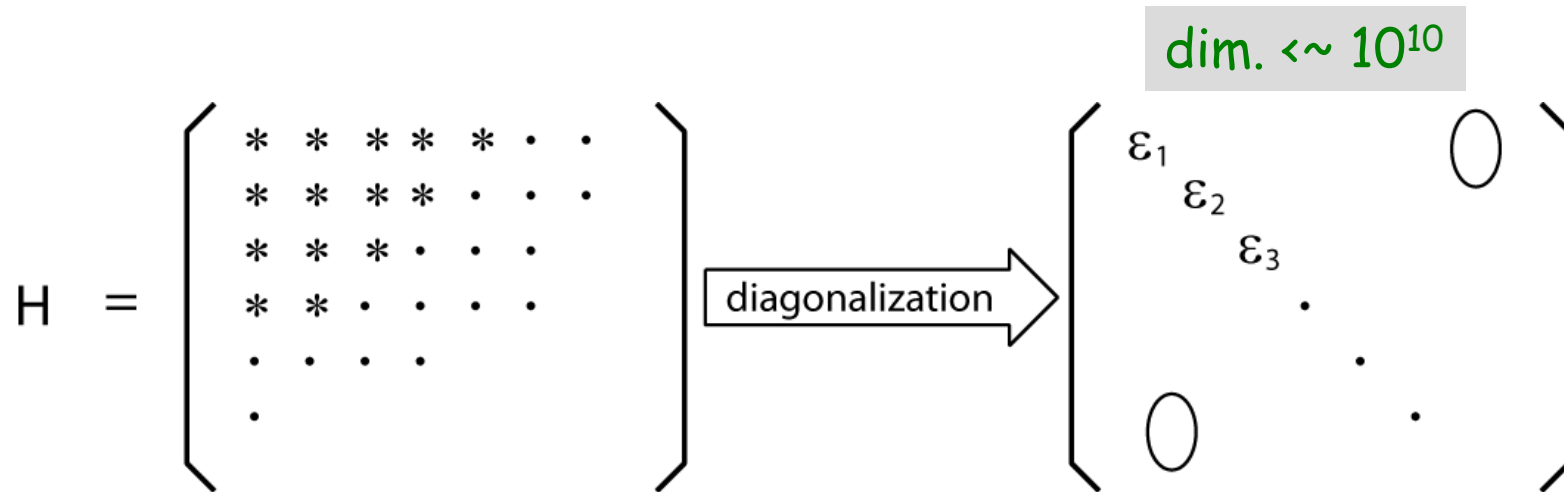
With this, we can calculate various physical quantities by
 $\langle \Psi' | T | \Psi \rangle$.

For instance, $E2, M1, \dots$ matrix elements

(transitions and moments)

spectroscopic factors (SF) and two-nucleon
amplitudes (TNA) for transfer reactions

Two types of shell-model calculations



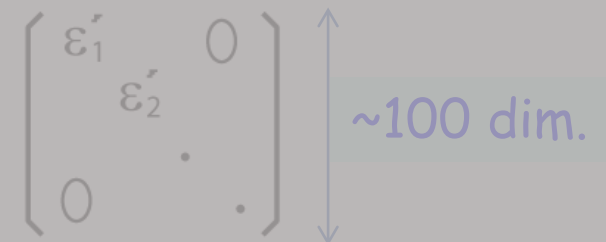
Conventional Shell Model
all Slater determinants

Direct diagonalization



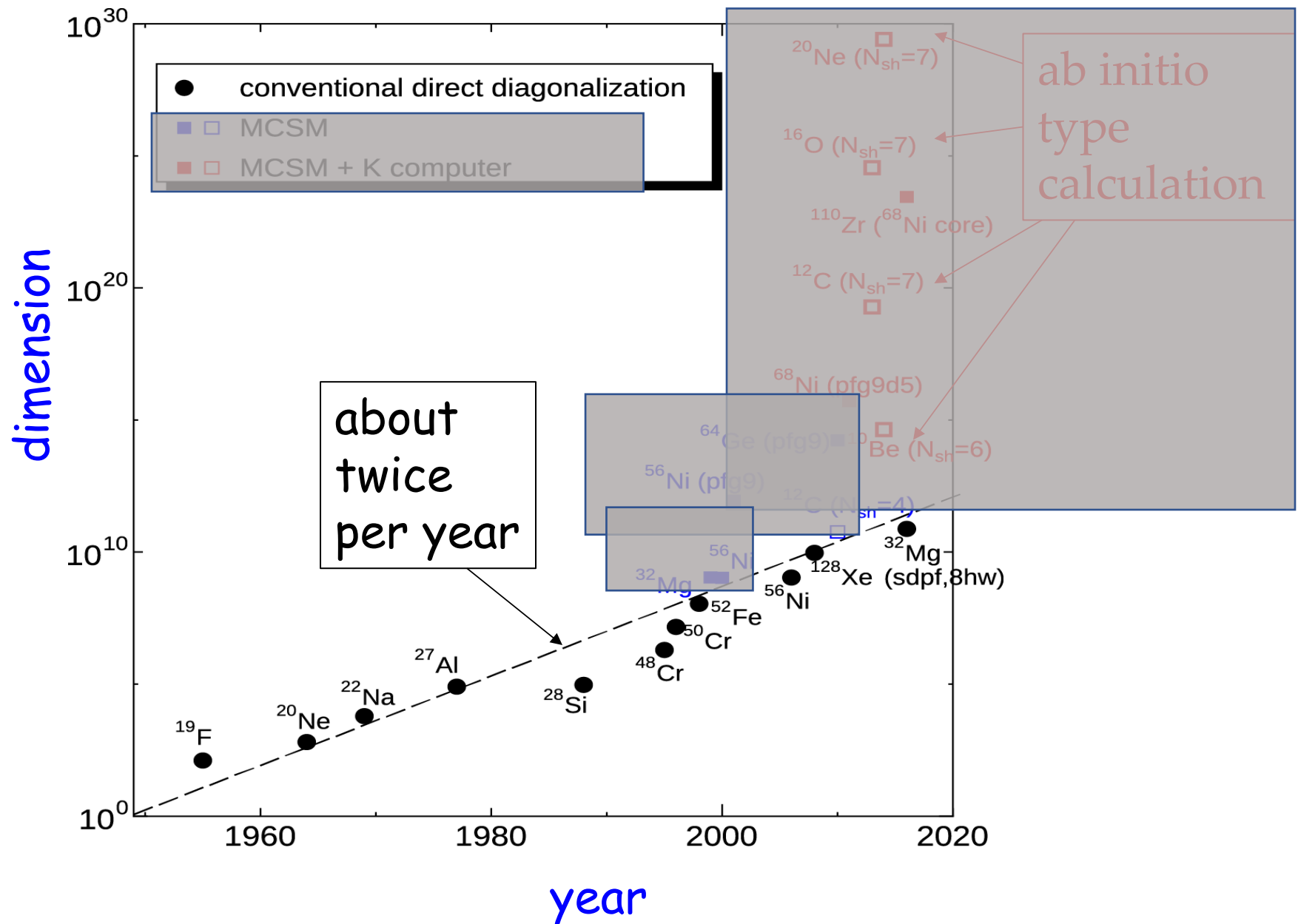
Monte Carlo Shell Model
bases important for a specific eigenstate

For even bigger problem,



Selected
important basis vectors

Dimension of the shell-model calculation



Monte Carlo Shell Model

Auxiliary-Field Monte Carlo (AFMC) method

general method for quantum many-body problems

For nuclear physics, **Shell Model Monte Carlo (SMMC)** calculation has been introduced by **Koonin** et al. Good for finite temperature.

- minus-sign problem
- only ground state, not for excited states in principle.

Quantum Monte Carlo Diagonalization (QMCD) method

No sign problem. Symmetries can be restored.

Excited states can be obtained.

→ **Monte Carlo Shell Model (MCSM)**

Background of Monte Carlo Shell Model (I)

Two-body interaction can be rewritten as $V = (1/2) \sum_{\alpha} v_{\alpha} \mathbf{O}_{\alpha}^2$

α : index

\mathbf{O}_{α} : one-body operators (rearranged by diagonalization)

Hubberd-Stratonovichi transformation

True eigenstate : $\psi = \sum_{\sigma} e^{-\beta h(\sigma)} e^{-\beta h(\sigma')} \dots \psi_0$

imaginary time (β) evolution by one-body field $h(\sigma)$

One-body operator is introduced as $h(\sigma) = \sum_{\alpha} s_{\alpha} \sigma_{\alpha} v_{\alpha} \mathbf{O}_{\alpha}$

σ_{α} : random number (Gaussian) σ : set of σ_{α} 's

s_{α} : phase (1 for $v_{\alpha} < 0$, i for $v_{\alpha} > 0$)

Background of Monte Carlo Shell Model (II)

True eigenstate : $\psi = \sum_{\sigma, \sigma', \dots} e^{-h(\sigma)} e^{-h(\sigma')} \dots \psi_0$

Use $\phi(\sigma, \sigma', \dots) = e^{-\beta h(\sigma)} e^{-\beta h(\sigma')} \dots \psi_0$
as a basis for shell-model diagonalization

$\phi(\sigma, \sigma', \dots)$ are selected and refined :

- (i) Random sampling -> only those lowering energy are kept
- (ii) Polished by varying σ, σ', \dots gradually (random noise reduced)
- (iii) Symmetry restoration (Angular momentum, parity)

Slater determinants (or Cooper-pair type wave functions)
are used

Usually, 20~50 bases are kept (many more thrown away)

Advanced Monte Carlo Shell Model (currently used)

N_B : number of basis vectors (dimension)

N_p : number of (active) particles

N_{sp} : number of single-particle states

$$|\Psi(D)\rangle = \sum_{n=1}^{N_B} c_n P^{J,\Pi} |\phi(D^{(n)})\rangle$$

amplitude

Projection op.

$$|\phi(D^{(n)})\rangle = \prod_{\alpha=1}^{N_p} \left(\sum_{i=1}^{N_{sp}} a_i^\dagger D_{i\alpha}^{(n)} \right) |-\rangle$$

n-th basis vector
(Slater determinant)

$$E(D) = \langle \Psi(D) | H | \Psi(D) \rangle$$

Minimize $E(D)$ as a function of D utilizing qMC and conjugate gradient methods

Stochastically "deformed"
single-particle state

Step 1 : stochastic generation of candidates of the **n-th MCSM** basis vector

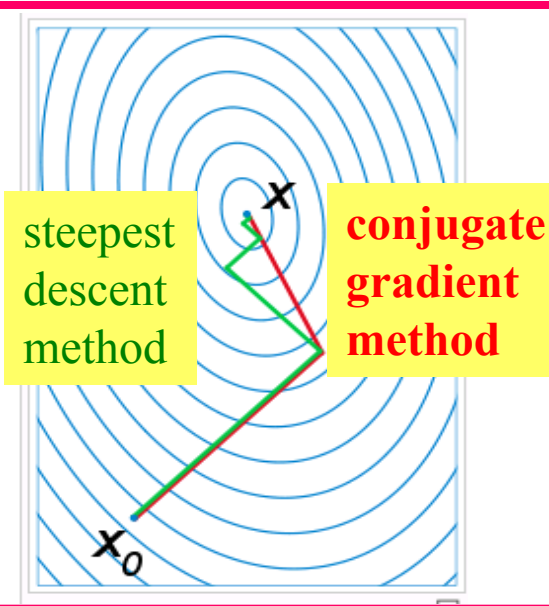
~~$$|\phi(\sigma)\rangle = \prod e^{\Delta\beta h(\sigma)} \cdot |\phi^{(0)}\rangle$$~~ *only theoretical background*

Shift randomly matrix elements of the matrix D .

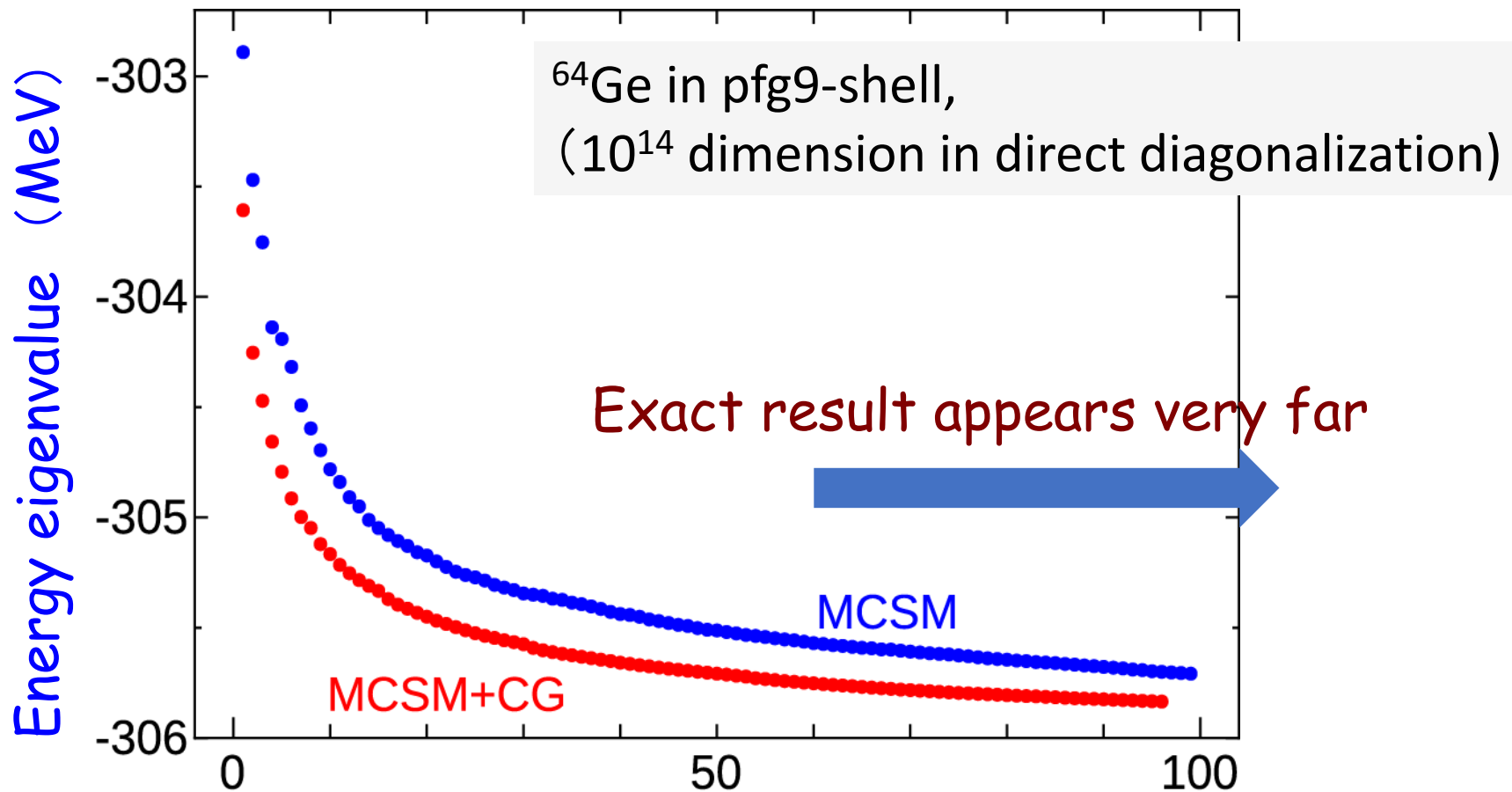
(The very initial one can be a Hartree-Fock state.)

Select the one producing the lowest $E(D)$ (rate < 0.1 %)

Step 2 : polish D by means of the **conjugate gradient (CG)** method "variationally".



Example of MCSM calculation



N_B : number of adapted basis vectors (Slater determinants)

Numerous MC trials and CG optimization for each basis vector

Advanced Monte Carlo Shell Model (currently used)

N_B : number of basis vectors (dimension)
 N_p : number of (active) particles
 N_{sp} : number of single-particle states

$$|\Psi(D)\rangle = \sum_{n=1}^{N_B} c_i P^{J,\Pi} |\phi(D^{(n)})\rangle$$

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$$|\phi(D^{(n)})\rangle = \prod_{\alpha=1}^{N_p} \left(\sum_{i=1}^{N_{sp}} a_i^\dagger D_{i\alpha}^{(n)} \right) |-\rangle$$

Deformed single-particle state n-th basis vector (Slater determinant)

$$E(D) = \langle \Psi(D) | H | \Psi(D) \rangle$$

Minimize $E(D)$ as a function of D utilizing qMC and conjugate gradient methods

Step 1 : stochastic generation of candidates of the n-th MCSM basis vector

~~$$|\phi(\sigma)\rangle = \prod e^{\Delta\beta h(\sigma)} \cdot |\phi^{(0)}\rangle$$~~

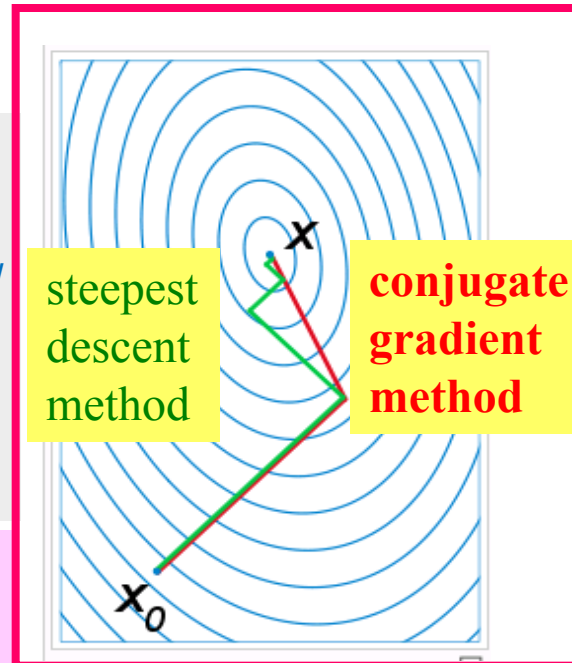
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Shift randomly matrix elements of the matrix D.

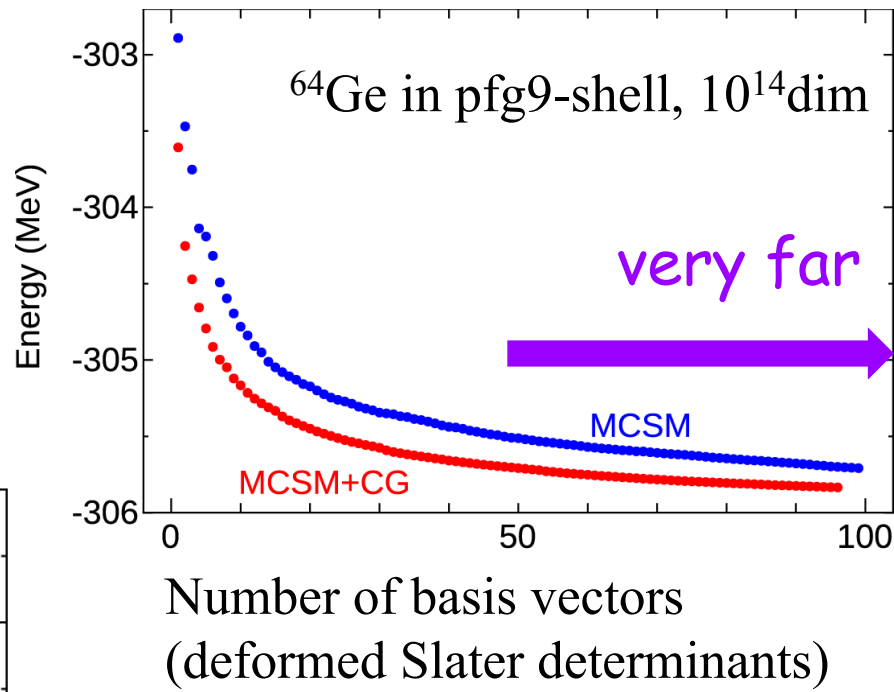
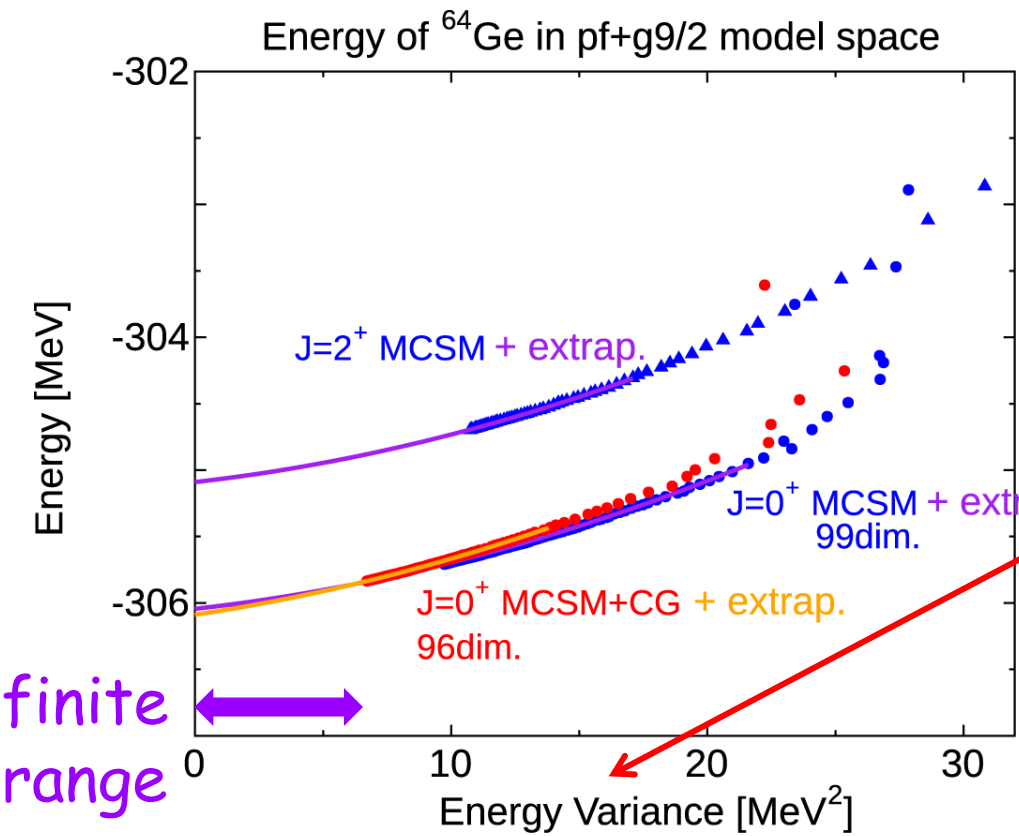
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Extrapolation by Energy Variance



Variance : $\langle \Delta H^2 \rangle = \langle H^2 \rangle - \langle H \rangle^2$

$$\langle H \rangle = E_0 + a \langle \Delta H^2 \rangle + b \langle \Delta H^2 \rangle^2 + \dots$$

N. Shimizu, et al.,
Phys. Rev. C **82**, 061305(R) (2010).

MCSM (Monte Carlo Shell Model -Advanced version-)

1. Selection of important many-body basis vectors
by **quantum Monte-Carlo** + diagonalization methods
*basis vectors : about 100 selected Slater determinants
composed of "deformed" single-particle states*
 2. **Variational** refinement of basis vectors
conjugate gradient method
 3. Variance **extrapolation** method -> **exact** eigenvalues
- + innovations in algorithm and code (=> now moving to GPU)



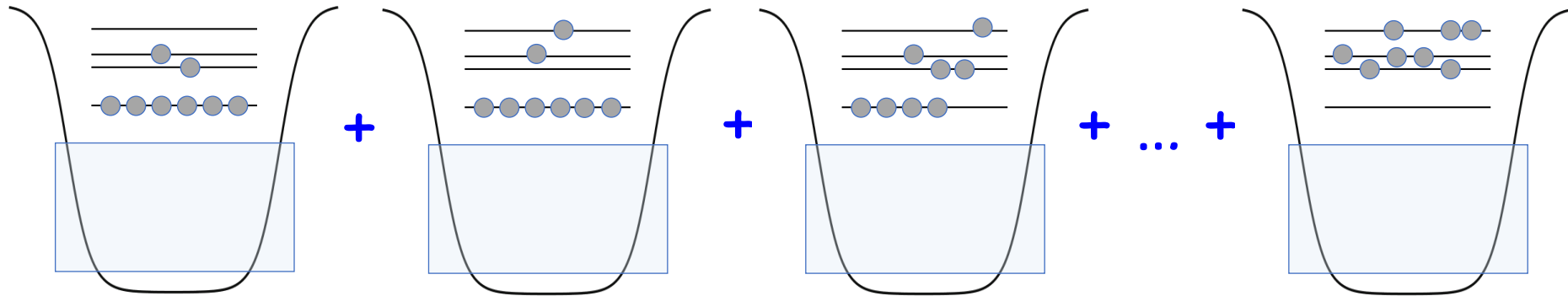
K computer (in Kobe) 10 peta flops machine

⇒ *Projection of basis vectors*

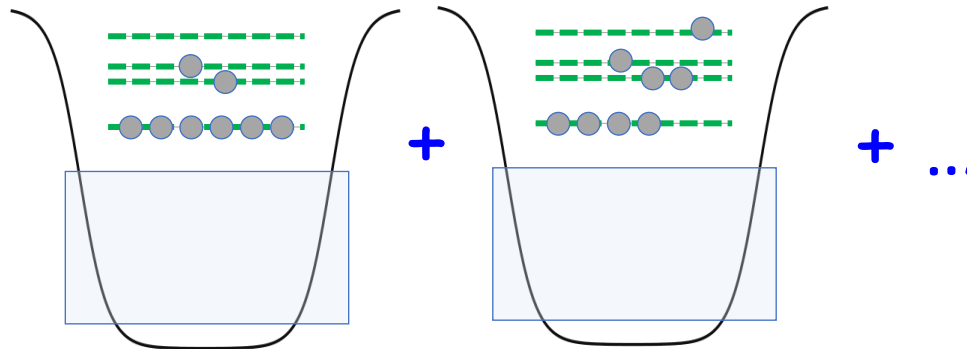
*Rotation with three Euler angles
with about 50,000 mesh points*

Example : $8^+ \text{ } ^{68}\text{Ni}$ 7680 core x 14 h

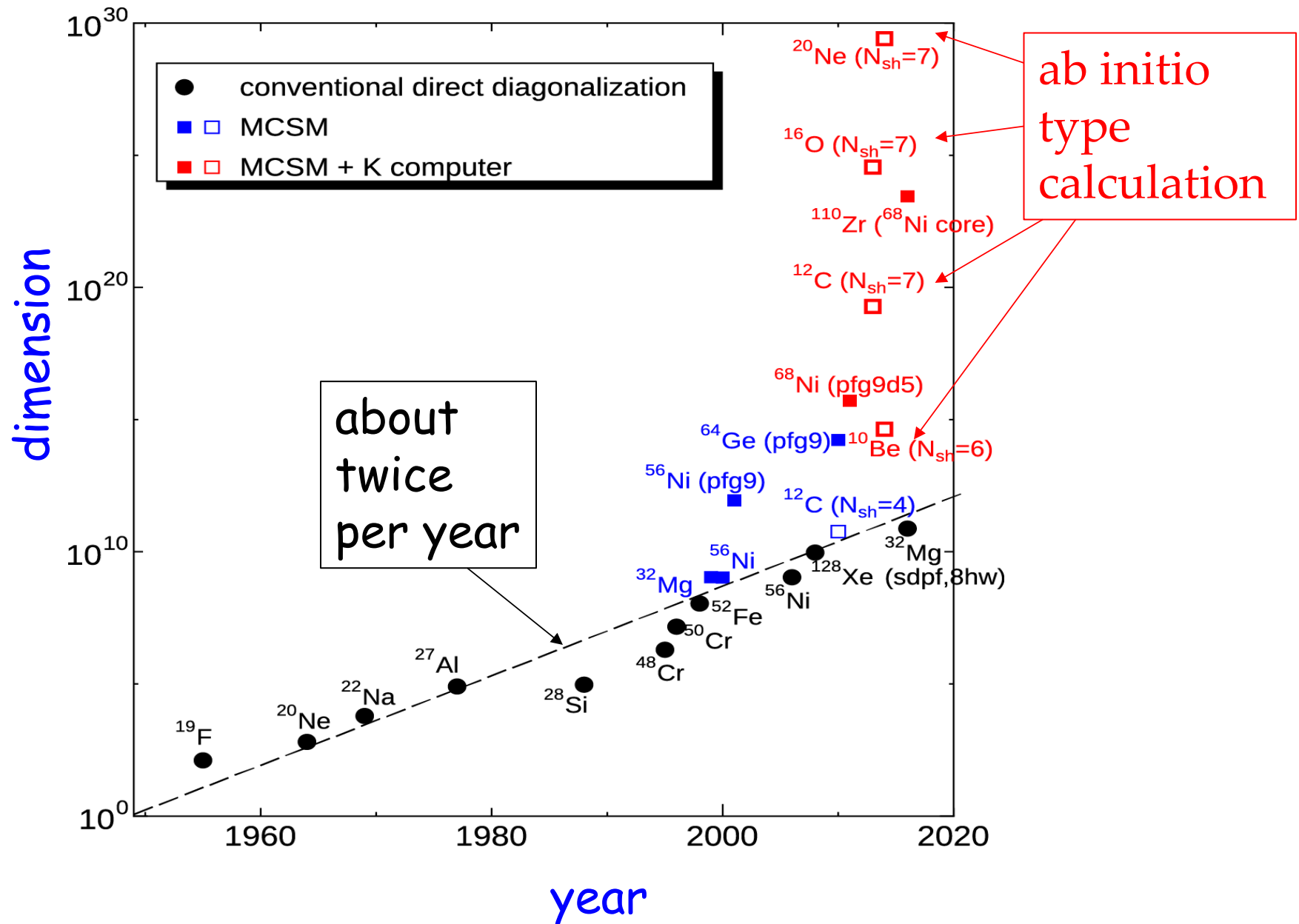
Possible configurations: 10^{23} ways at maximum for Zr isotopes to be discussed



Superposition of original orbits \Rightarrow Select most important ~ 100 ones



Dimension of the shell-model many-body Hilbert space



MCSM basis vectors on Potential Energy Surface

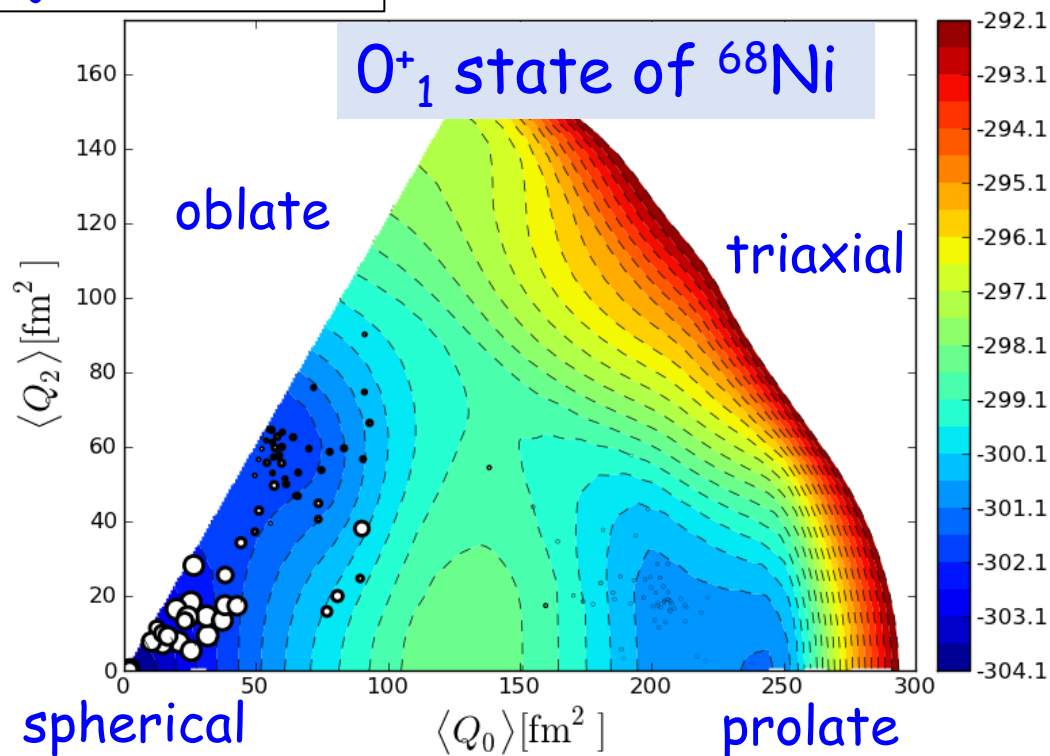
eigenstate $\Psi = \sum_i c_i P[J^\pi](\Phi_i)$

amplitude

projection onto J^π

Slater determinant of deformed s. p. states \rightarrow intrinsic shape

- **PES** is calculated by **CHF** for the shell-model Hamiltonian
- **Location of circle** : quadrupole deformation of unprojected MCSM basis vectors
- **Area of circle** : overlap probability between each projected basis and eigen wave function



Called ***T-plot*** in reference to
Y. Tsunoda, *et al.*
PRC 89, 031301 (R) (2014)



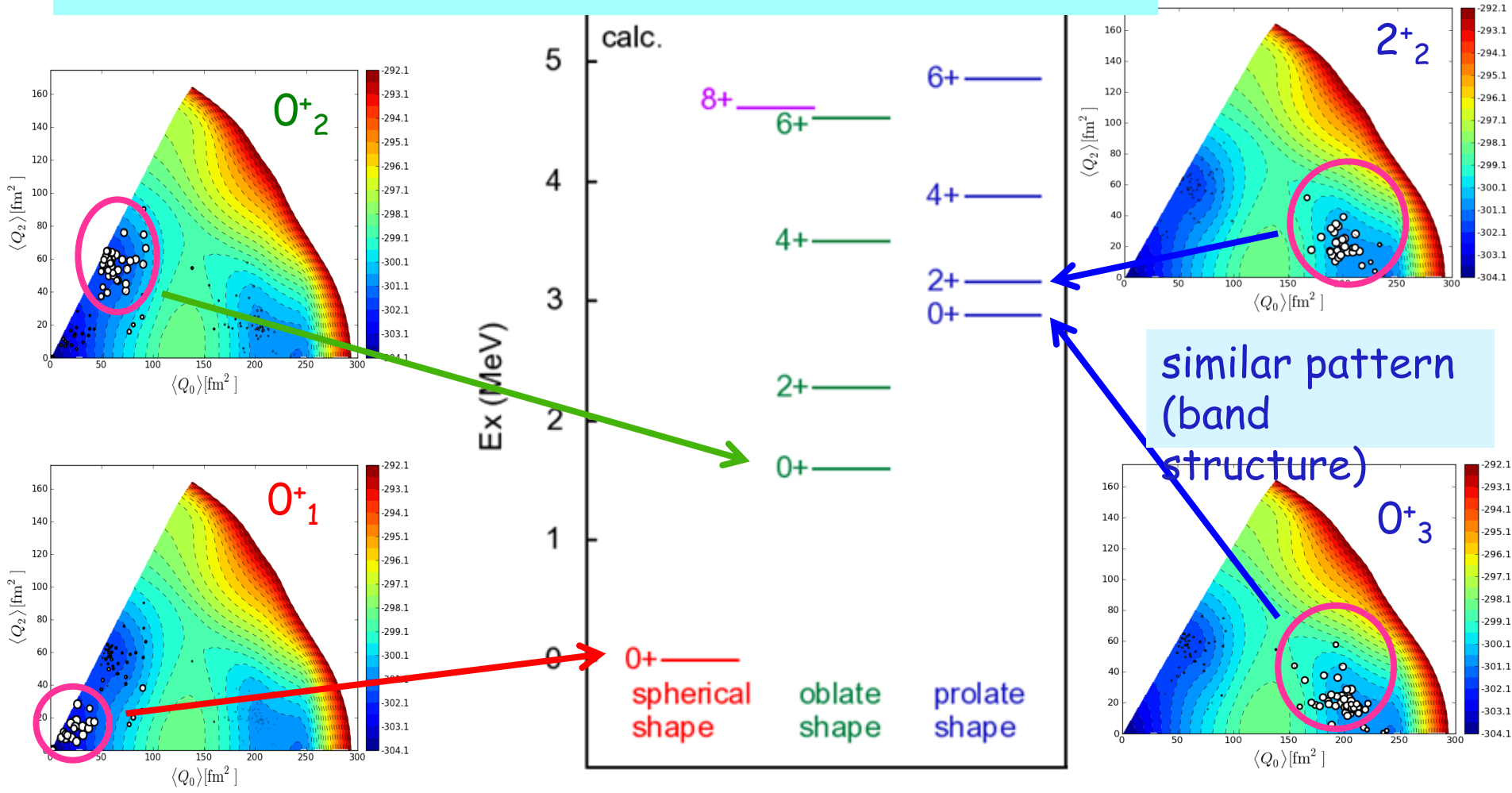
General properties of T-plot :

Certain number of large circles in a small region of PES

⇔ pairing correlations

Spreading beyond this can be due to shape fluctuation

Example : shape assignment to various 0^+ states of ^{68}Ni



Three pillars combined for future

computation

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Shell Model
(MCSM)

(almost)
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massive
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Hamiltonian

pf
pfg9d5 (A3DA) (Ni)
8+8 on ^{56}Ni core (Zr)
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Shell evolution
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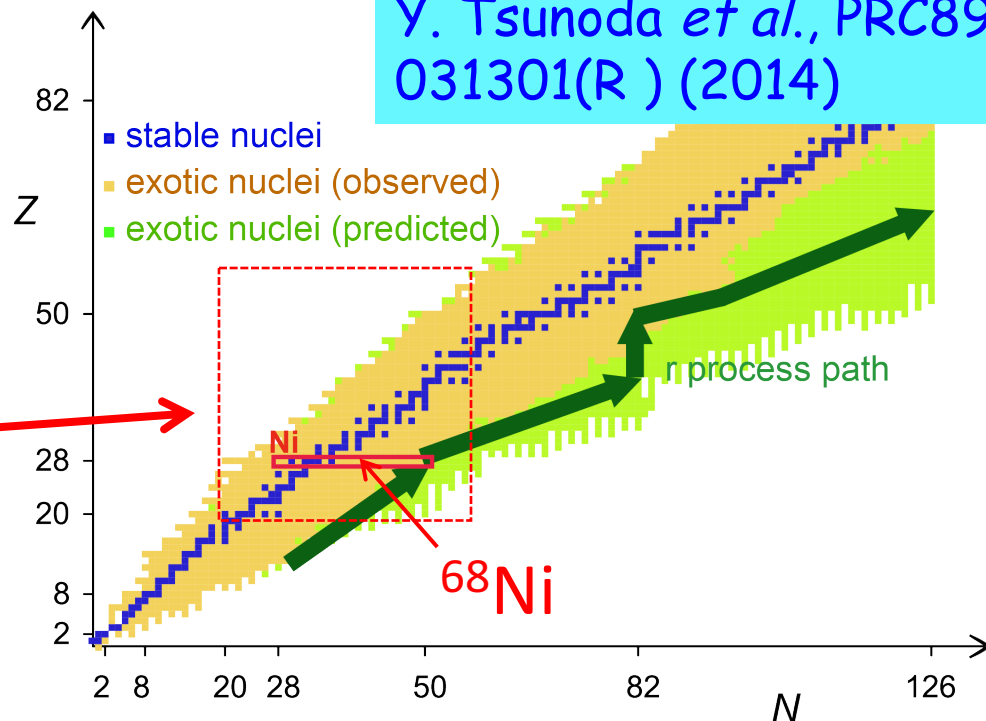
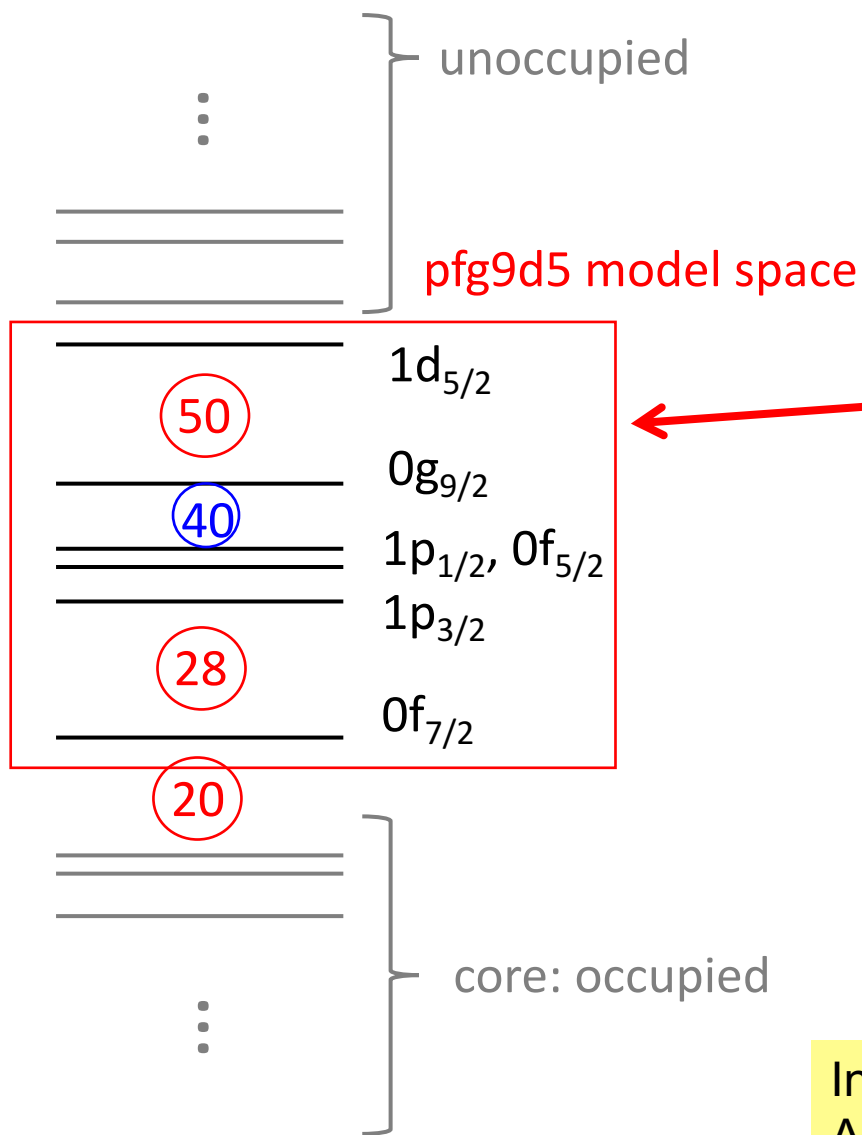
Quantum Phase
Transition

Shape coexistence

Quantum
Self-organization

Monte Carlo Shell Model (MCSM) calculation on Ni isotopes

Y. Tsunoda *et al.*, PRC89, 031301(R) (2014)



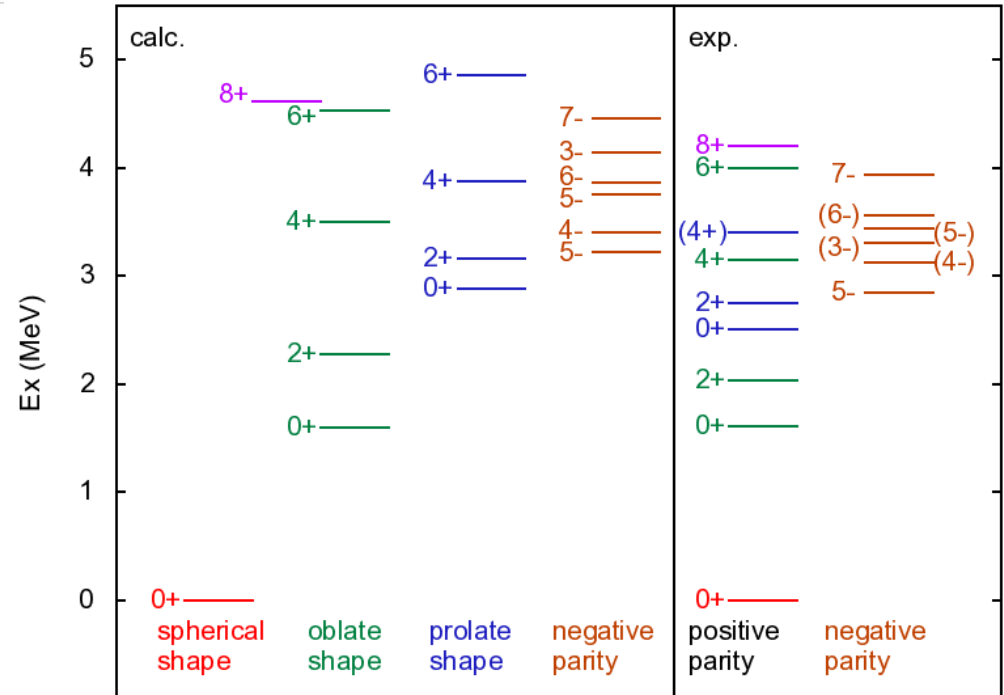
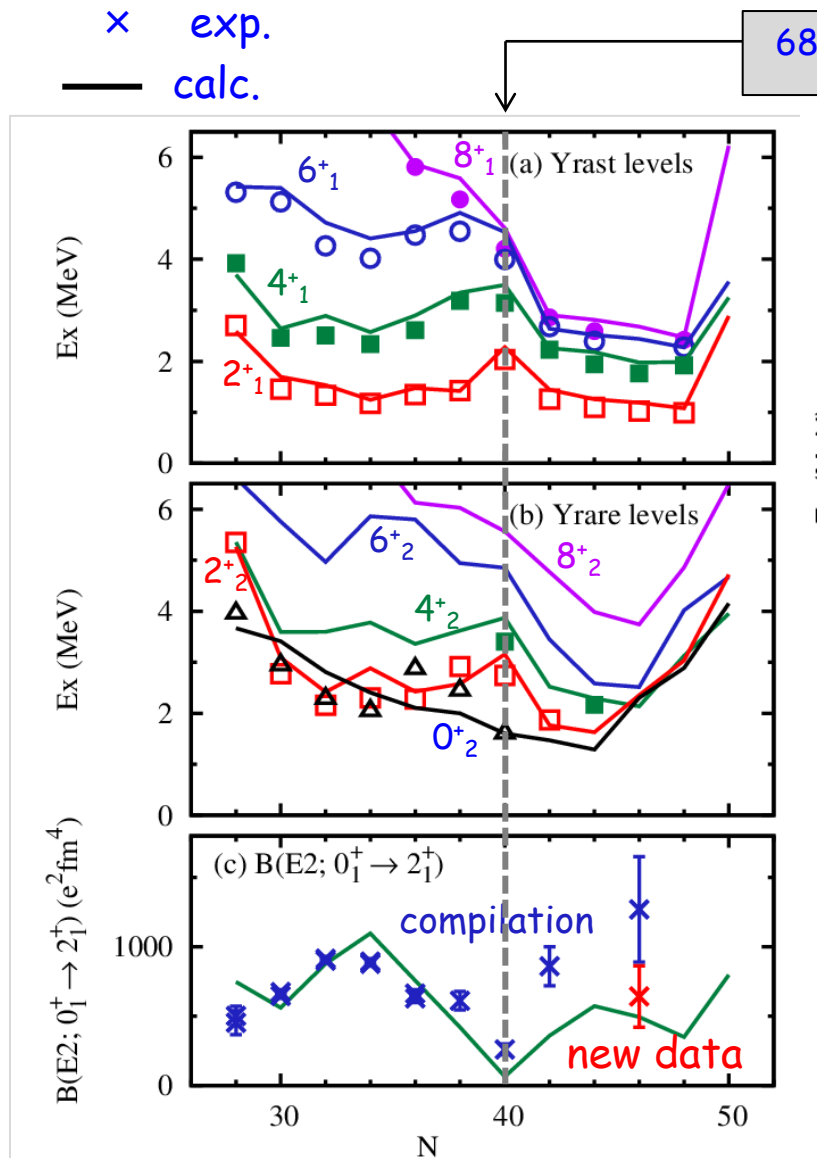
This model space is wide enough to discuss how **magic numbers 28, 50** and **semi-magic number 40** are visible or smeared out.

Interaction:
A3DA interaction is used with minor corrections

Energy levels and B(E2) values of Ni isotopes

Description by the same Hamiltonian

Shape coexistence in ^{68}Ni



Y. Tsunoda, TO, Shimizu, Honma and Utsuno, PRC 89, 031301 (R) (2014)



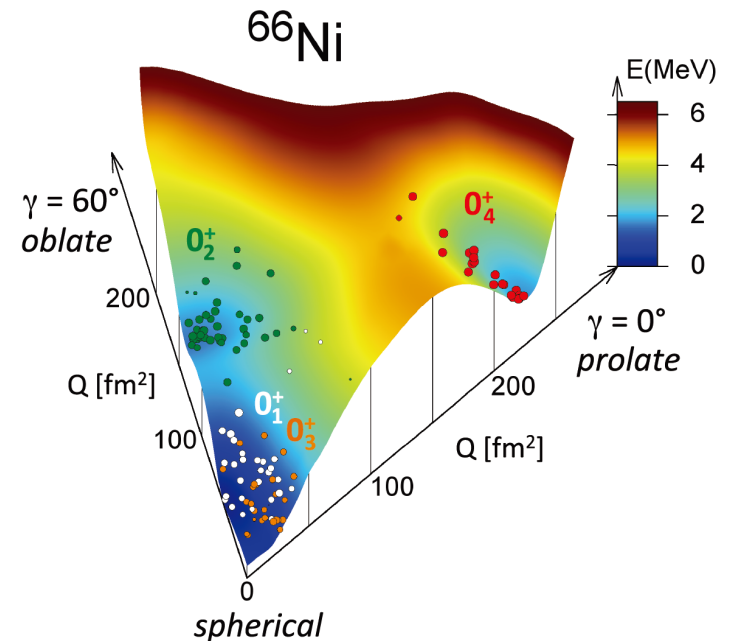
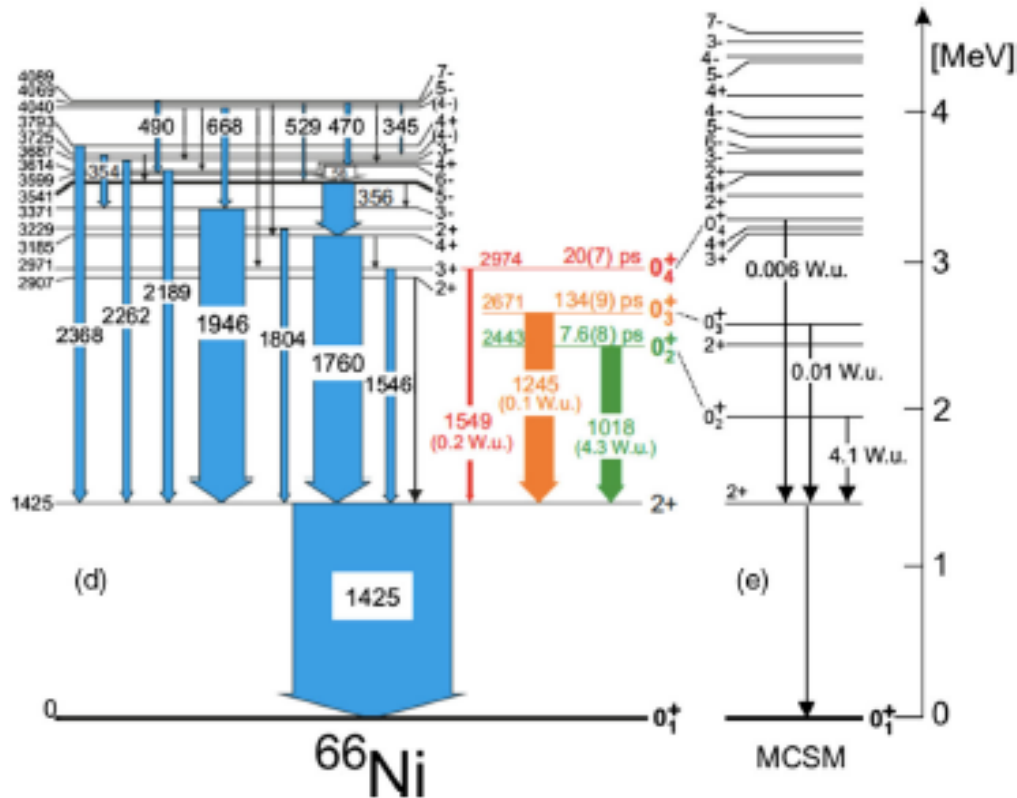
Multifaceted Quadruplet of Low-Lying Spin-Zero States in ^{66}Ni : Emergence of Shape Isomerism in Light Nuclei

S. Leoni,^{1,2,*} B. Fornal,³ N. Mărginean,⁴ M. Sferrazza,⁵ Y. Tsunoda,⁶ T. Otsuka,^{6,7,8,9} G. Bocchi,^{1,2} F.C. L. Crespi,^{1,2}
A. Bracco,^{1,2} S. Aydin,¹⁰ M. Boromiza,^{4,11} D. Bucurescu,⁴ N. Cieplicka-Oryńczak,^{2,3} C. Costache,⁴ S. Călinescu,⁴
N. Florea,⁴ D.G. Ghiță,⁴ T. Glodariu,⁴ A. Ionescu,^{4,11} Ł. W. Iskra,³ M. Krzysiek,³ R. Mărginean,⁴ C. Mihai,⁴ R. E. Mihai,⁴
A. Mitu,⁴ A. Negreț,⁴ C. R. Niță,⁴ A. Olăcel,⁴ A. Oprea,⁴ S. Pascu,⁴ P. Petkov,⁴ C. Petrone,⁴ G. Porzio,^{1,2} A. Șerban,^{4,11}
C. Sotty,⁴ L. Stan,⁴ I. Știru,⁴ L. Stroe,⁴ R. Șuvăilă,⁴ S. Toma,⁴ A. Turturică,⁴ S. Ujeniuc,⁴ and C. A. Ur^{1,2}

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²INFN sezione di Milano via Celoria 16, 20133, Milano, Italy

³Institute of Nuclear Physics, PAN, 31-342 Kraków, Poland

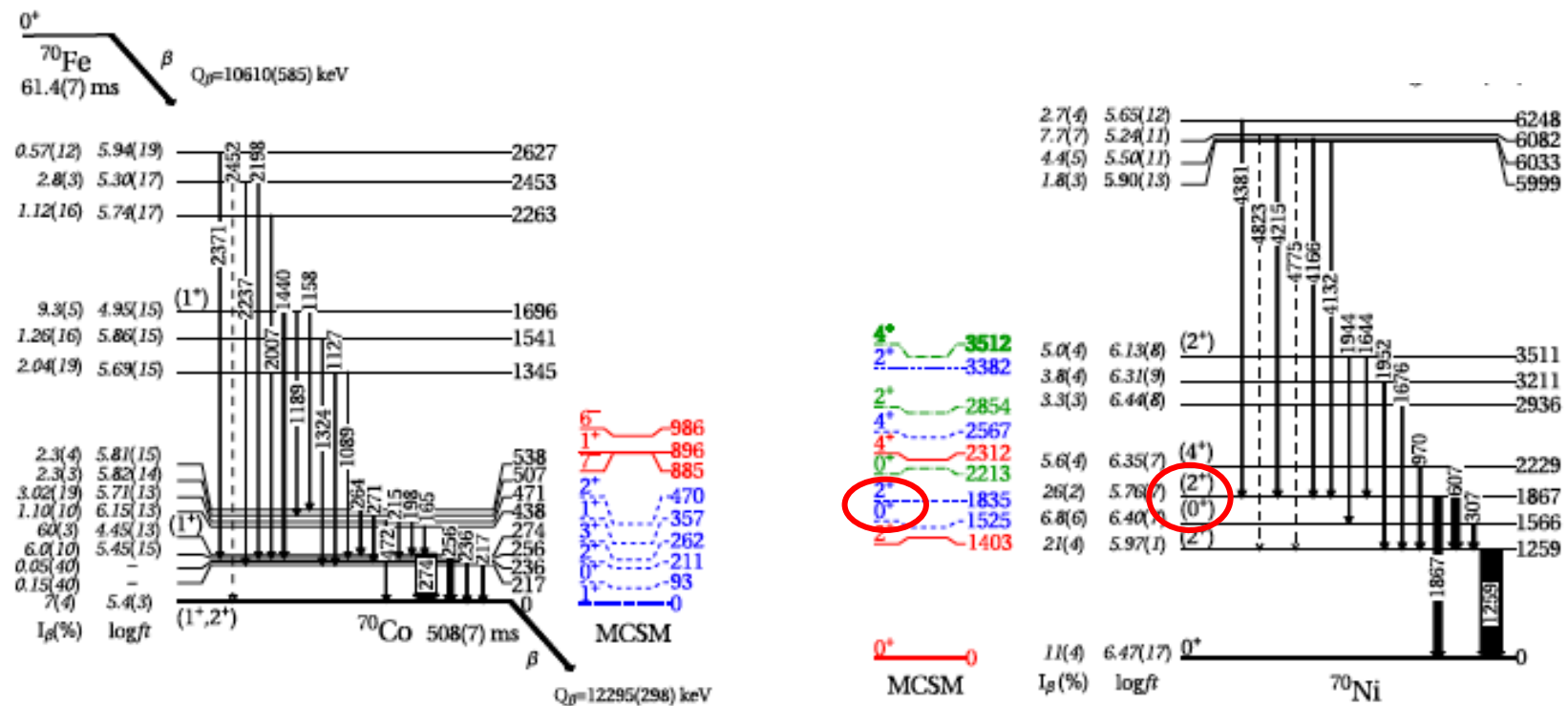


Type II shell evolution in $A = 70$ isobars from the $N \geq 40$ island of inversion

A.I. Morales^{a,b,*}, G. Benzoni^a, H. Watanabe^{c,d}, Y. Tsunoda^e, T. Otsuka^{f,g,h}, S. Nishimura^d, F. Browne^{i,d}, R. Daido^j, P. Doornenbal^d, Y. Fang^j, G. Lorusso^d, Z. Patel^{k,d}, S. Rice^{k,d}, L. Sinclair^{l,d}, P.-A. Söderström^d, T. Sumikama^m, J. Wu^d, Z.Y. Xu^{f,d}, A. Yagi^j, R. Yokoyama^f, H. Baba^d, R. Avigo^{a,b}, F.L. Bello Garroteⁿ, N. Blasi^a, A. Bracco^{a,b}, F. Camera^{a,b}, S. Ceruti^{a,b}, F.C.L. Crespi^{a,b}, G. de Angelis^o, M.-C. Delattre^p, Zs. Dombradi^q, A. Gottardo^o, T. Isobe^d, I. Kojouharov^r, N. Kurz^r, I. Kuti^q, K. Matsui^f, B. Melon^s, D. Mengoni^{t,u}, T. Miyazaki^f, V. Modamio-Hoybjør^o, S. Momiyama^f, D.R. Napoli^o, M. Niikura^f, R. Orlandi^{h,v}, H. Sakurai^{d,f}, E. Sahinⁿ, D. Sohler^q, H. Schaffner^r, R. Taniuchi^f, J. Taprogge^{w,x}, Zs. Vajta^q, J.J. Valiente-Dobón^o, O. Wieland^a, M. Yalcinkaya^y

^a Istituto Nazionale di Fisica Nucleare, Sezione di Milano, Via Celoria 16, 20133 Milano, Italy

^b Dipartimento di Fisica, Università degli Studi di Milano, Via Celoria 16, 20133 Milano, Italy



End of the 1st lecture

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*Recent developments
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2nd lecture



This work has been supported by MEXT and JICFuS as a priority issue (Elucidation of the fundamental laws and evolution of the universe) to be tackled by using Post ‘K’ Computer

Three pillars combined for future

In order to understand the physics behind these calculations, let us jump to

computation

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Shell Model
(MCSM)

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pf
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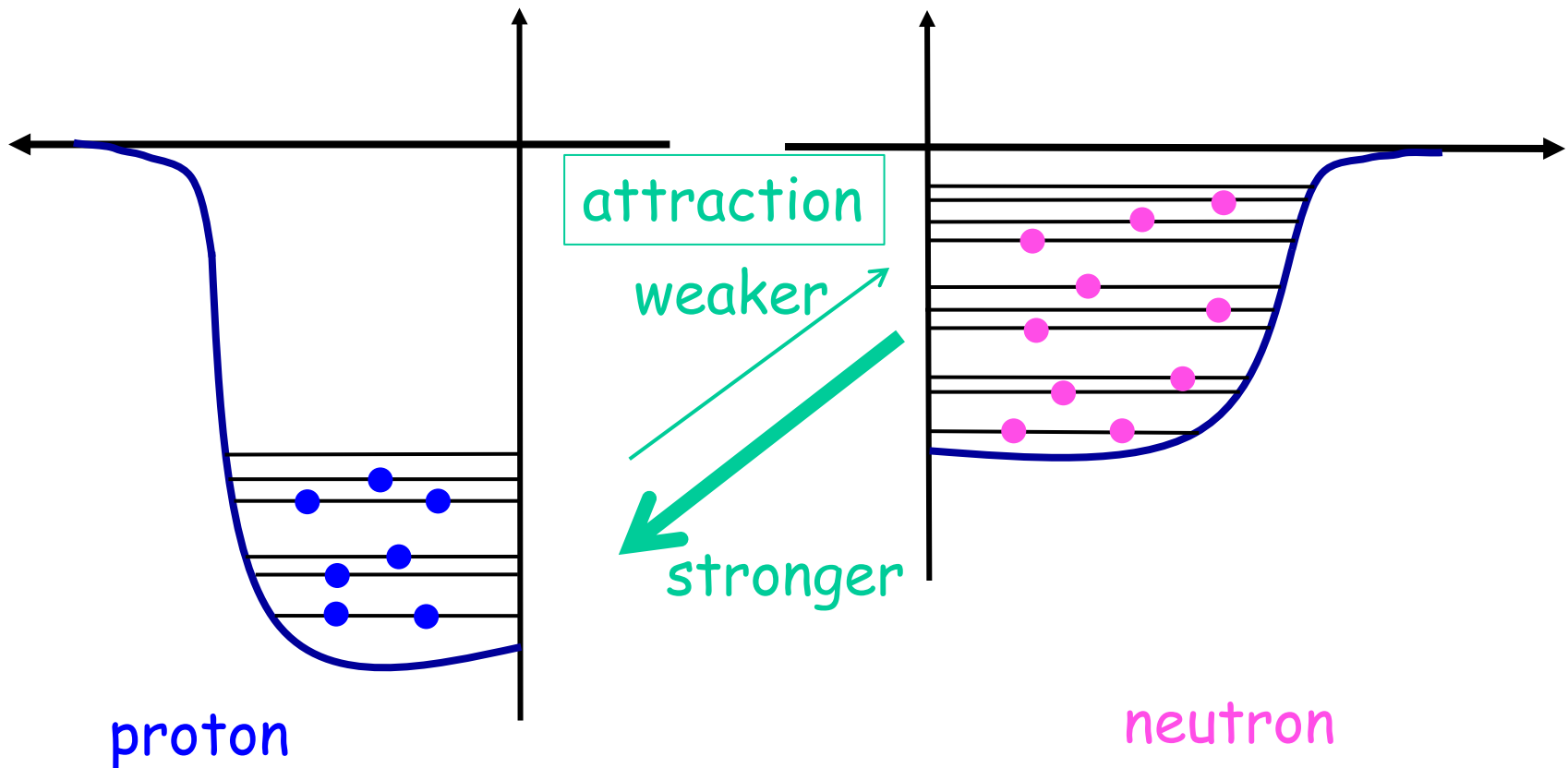
Quantum
Self-organization

proton-neutron interaction

>> proton-proton or neutron-neutron

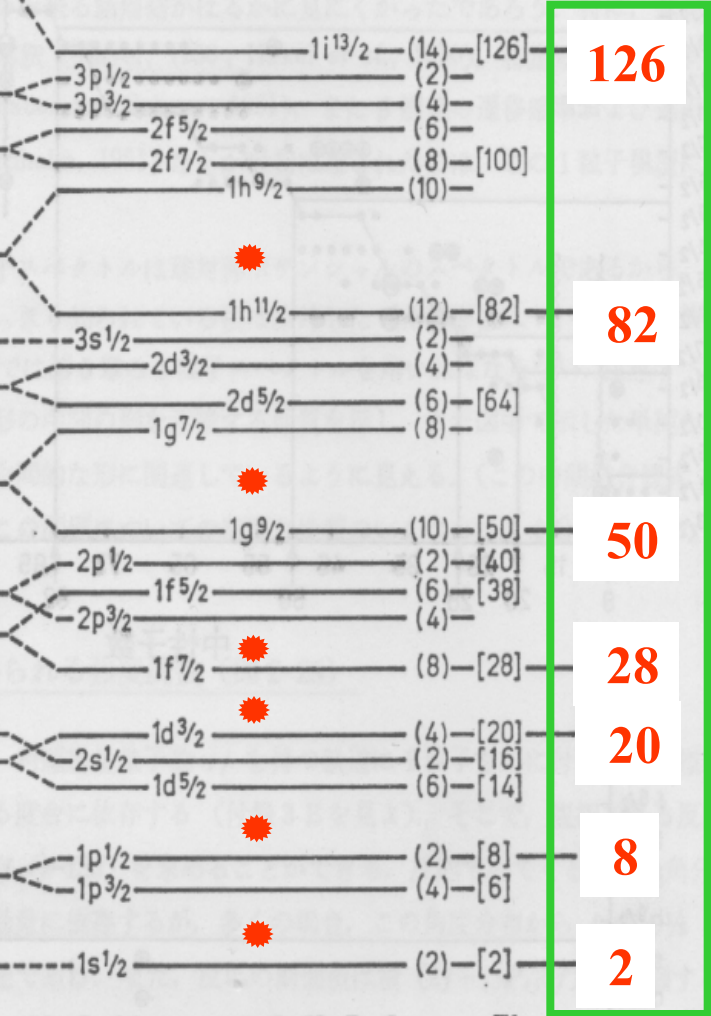
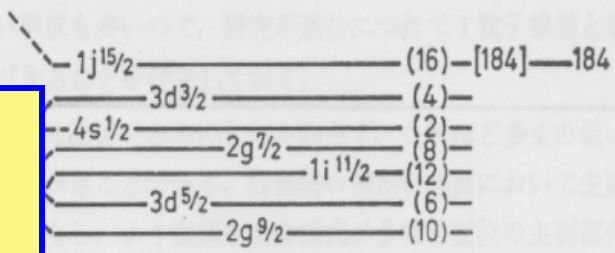
If $Z \ll N$, protons are more bound.

Relative relations among orbits are preserved, because basically only the depths change. Can we assume this?



Eigenvalues of
HO potential

5ħω
4ħω
3ħω
2ħω
1ħω
0



Magic numbers
by
Mayer and
Jensen (1949)



R SHELL MODEL

図 2-23 1 粒子軌道の順序。図は M. G. Mayer and J. H. D. Jensen, *Elementary Theory of Nuclear Shell Structure*, p. 58, Wiley, New York, 1955 からとった。

From simple but general properties as,

density saturation
+ short-range NN interaction
+ spin-orbit splitting

→ Mayer-Jensen's magic number
with rather **constant gaps**
(except for gradual A dependence)

robust feature → no way out ???

This question leads us to one of the major developments of recent nuclear structure studies.

Let's see what occurs
in the shell structure of exotic nuclei.

In other words, let's discuss whether
the **shell evolution**,
the change of the shell structure,
occurs or not.

If it occurs, how ?

The key tool is
the **monopole interaction**.

Monopole matrix element between orbits j and j'

$$V_{nn}^m(j, j') = \frac{\sum_{(m, m')} \langle j, m; j', m' | \hat{v}_{nn} | j, m; j', m' \rangle}{\sum_{(m, m')} 1},$$

v_{nn} is interaction; m, m' are magnetic substates

Monopole matrix element between orbits j and j'

$$\begin{aligned} & \langle \text{red up} \text{ blue up} | v | \text{red up} \text{ blue up} \rangle + \langle \text{red up} \text{ blue up} | v | \text{red up} \text{ blue up} \rangle + \langle \text{red up} \text{ blue up} | v | \text{red up} \text{ blue up} \rangle + \dots \\ & + \langle \text{red up} \text{ blue up} | v | \text{red up} \text{ blue up} \rangle + \dots \dots \dots \dots + \langle \text{red up} \text{ blue up} | v | \text{red up} \text{ blue up} \rangle \\ & = \frac{\dots}{\text{number of matrix elements in the summation}} \end{aligned}$$

   : magnetic substates of the orbit j
   : magnetic substates of the orbit j'

Monopole matrix elements can be written equivalently by usual TBMEs as

$$V_T^m(j, j') = \frac{\sum_J (2J + 1) \langle j, j'; J, T | \hat{v} | j, j'; J, T \rangle}{\sum_J (2J + 1)}.$$

for $T = 0$ and 1 ,

Monopole interaction for n-n (or p-p) interaction is defined as

$$\hat{v}_{nn}^m(j, j') = \begin{cases} V_{nn}^m(j, j) \frac{1}{2} \hat{n}_j (\hat{n}_j - 1) & \text{for } j = j' \\ V_{nn}^m(j, j') \hat{n}_j \hat{n}_{j'} & \text{for } j \neq j' \end{cases}$$

T=1 part of the p-n monopole interaction is given as

$$\begin{aligned}\hat{v}_{pn,mono,T=1} = & \sum_{j,j'} V_{T=1}^m(j, j') \frac{1}{2} \hat{n}_j^p \hat{n}_{j'}^n \\ & + \sum_{j < j'} V_{T=1}^m(j, j') \frac{1}{2} \left\{ \hat{\tau}_j^+ \hat{\tau}_{j'}^- + \hat{\tau}_j^- \hat{\tau}_{j'}^+ \right\} \\ & + \sum_j V_{T=1}^m(j, j) \frac{1}{2} : \hat{\tau}_j^+ \hat{\tau}_j^- : .\end{aligned}$$

Finally, we get the full expression for the p-n interaction

$$\begin{aligned}\hat{v}_{pn,mono} = & \sum_{j,j'} \frac{1}{2} \left\{ V_{T=0}^m(j, j') + V_{T=1}^m(j, j') \right\} \boxed{\hat{n}_j^p \hat{n}_{j'}^n} \\ & - \sum_{j < j'} \frac{1}{2} \left\{ V_{T=0}^m(j, j') - V_{T=1}^m(j, j') \right\} \\ & \quad \left\{ \hat{\tau}_j^+ \hat{\tau}_{j'}^- + \hat{\tau}_j^- \hat{\tau}_{j'}^+ \right\} \\ & - \sum_j \frac{1}{2} \left\{ V_{T=0}^m(j, j) - V_{T=1}^m(j, j) \right\} : \hat{\tau}_j^+ \hat{\tau}_j^- : .\end{aligned}$$

This is consistent with the final form of the monopole interaction of Poves and Zuker (Phys. Rep. 70, 235 (1981)) besides different formulation.

$$\begin{aligned}
\hat{v}_{pn,mono} = & \sum_{j,j'} \frac{1}{2} \left\{ V_{T=0}^m(j, j') + V_{T=1}^m(j, j') \right\} \hat{n}_j^p \hat{n}_{j'}^n \\
& - \sum_{j < j'} \frac{1}{2} \left\{ V_{T=0}^m(j, j') - V_{T=1}^m(j, j') \right\} \\
& \quad \left\{ \hat{\tau}_j^+ \hat{\tau}_{j'}^- + \hat{\tau}_j^- \hat{\tau}_{j'}^+ \right\} \\
& - \sum_j \frac{1}{2} \left\{ V_{T=0}^m(j, j) - V_{T=1}^m(j, j) \right\} : \hat{\tau}_j^+ \hat{\tau}_j^- :
\end{aligned}$$

The second and third terms can be visualized as shown in the figure.

They are still monopole, but proton and neutron must exchange their states !

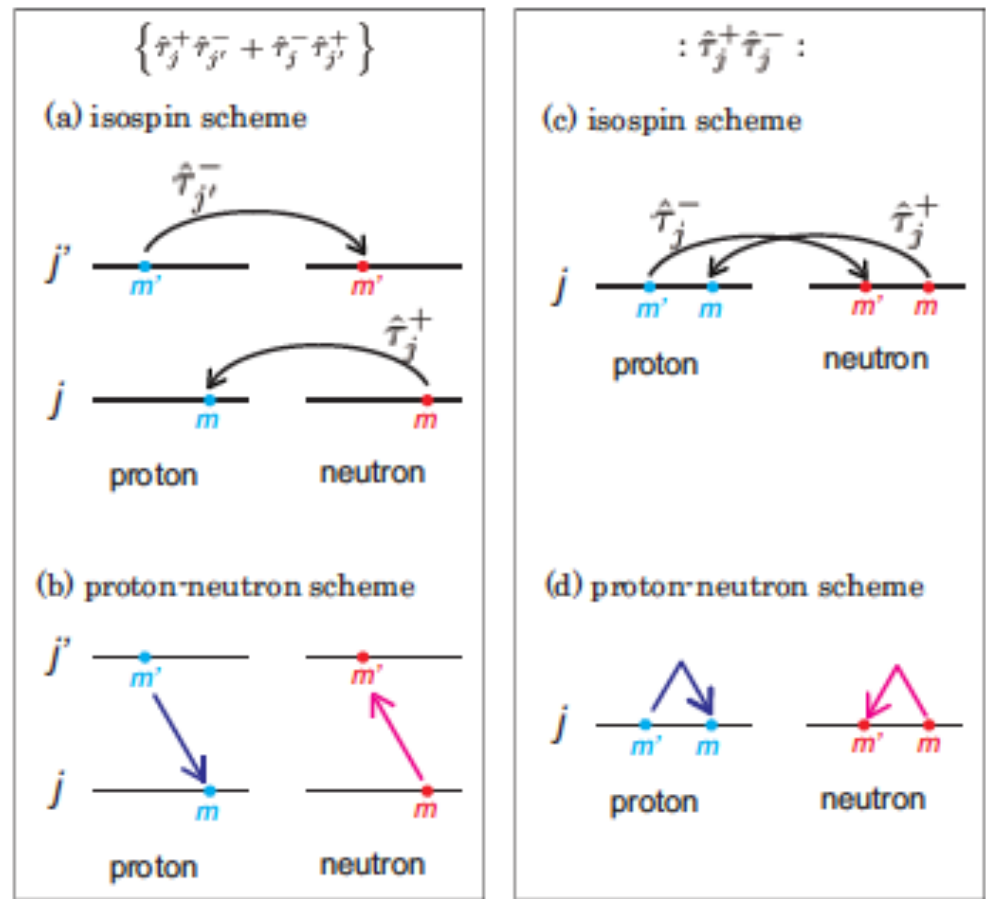


Figure 16 Implication of $\hat{\tau}_j^+ \hat{\tau}_{j'}^-$ terms. Panels (a) and (c) are for the $\{\hat{\tau}_j^+ \hat{\tau}_{j'}^- + \hat{\tau}_j^- \hat{\tau}_{j'}^+\}$ and $:\hat{\tau}_j^+ \hat{\tau}_j^-:$ cases in the isospin scheme, respectively. Panels (b) and (d) are similar to (a) and (c), respectively, in the proton-neutron scheme. The magnetic substates are indicated by m and m' .

Effective single-particle energy (ESPE)

$$\epsilon_j = \epsilon_j^0 + \hat{\epsilon}_j .$$

Contribution from **valence nucleons** through the monopole interaction

contribution from the **inert core** (closed shell) ; constant within a given nucleus

Changes of ESPEs between nuclei

for protons

$$\Delta \hat{\epsilon}_j^p = \sum_{j'} V_{T=1}^m(j, j') \Delta \hat{n}_{j'}^p + \sum_{j'} V_{pn}^m(j, j') \Delta \hat{n}_{j'}^n$$

for neutrons

$$\Delta \hat{\epsilon}_j^n = \sum_{j'} V_{T=1}^m(j, j') \Delta \hat{n}_{j'}^n + \sum_{j'} V_{pn}^m(j', j) \Delta \hat{n}_{j'}^p$$

For a chain of isotopes, $\Delta \hat{n}_{j'}^n$ denotes the change of the neutron number in the orbit j' .

If monopole matrix elements, $V_{T=1}^m(j, j')$ and $V_{pn}^m(j, j')$, are uniform, nothing happens.

However, the **shell evolution** occurs, if $V_{T=1}^m(j, j')$ or $V_{pn}^m(j, j')$ change significantly depending on j and j' .

Monopole matrix element of the **central force** with a Gaussian dependence on the distance.

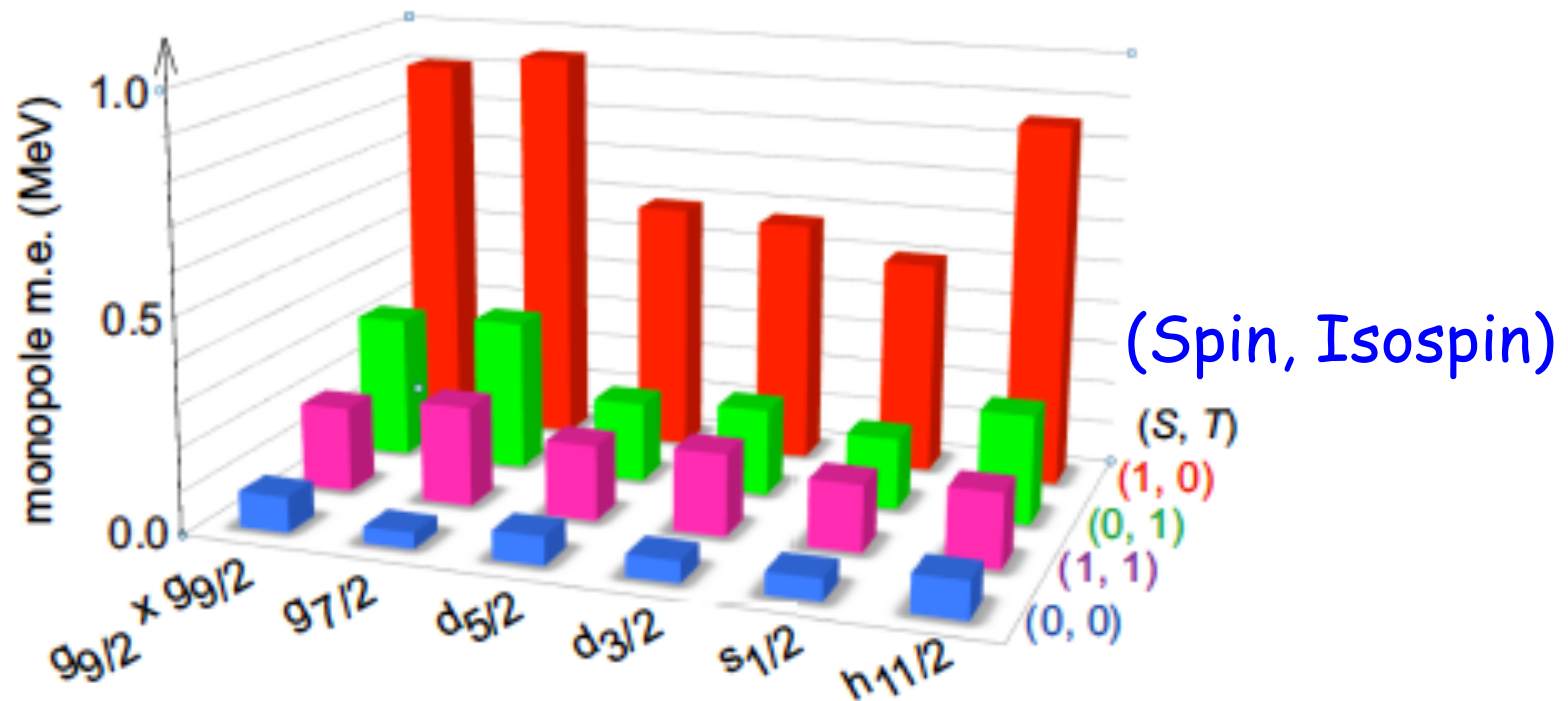
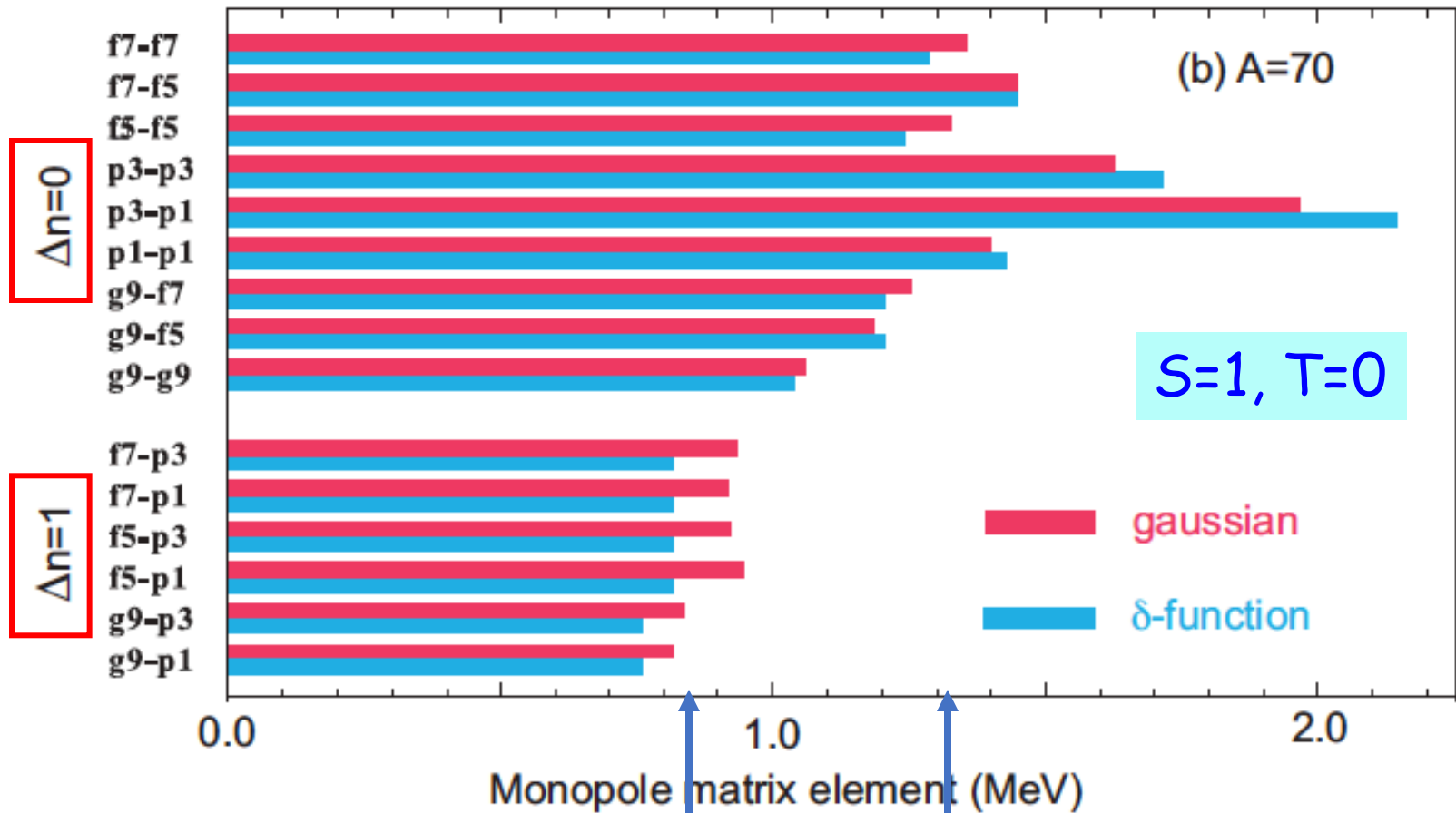


Figure 24 Monopole matrix elements of central gaussian interactions for (S, T) channels with an equal strength parameter (see the text). One of the orbit is $1g_{9/2}$, and the other is shown.

Monopole matrix element from a central force : $A=70$



mean values ~ 0.8 MeV ~ 1.3 MeV

variations ~ 0.1 MeV ~ 0.3 MeV

difference
 $\sim < 0.5$ MeV

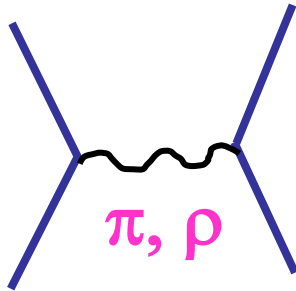
Figure 26 Monopole matrix elements of central gaussian and delta interactions for ($S = 1, T = 0$) channel. The orbit labeling is abbreviated like g9 for $1g_{9/2}$, etc. The orbits are from valence shell for (a) $A = 100$ and (b) $A = 70$

Besides central force,
another important contribution
comes from the tensor force

Tensor Force

π meson : primary source

ρ meson ($\sim \pi+\pi$) : minor ($\sim 1/4$) cancellation



Ref: Osterfeld, Rev. Mod. Phys. 64, 491 (92)

Multiple pion exchanges

→ strong effective central forces in NN interaction
(as represented by σ meson, etc.)

→ nuclear binding

Presently : First-order tensor-force effect
(at medium and long ranges)

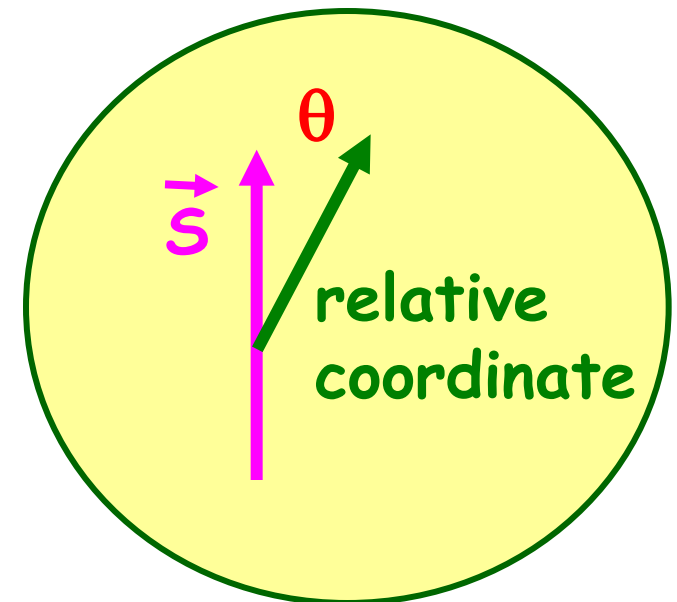
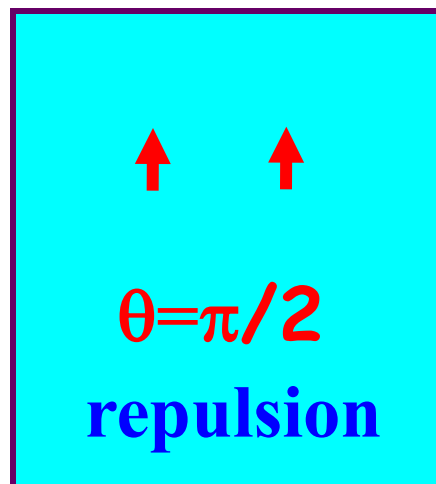
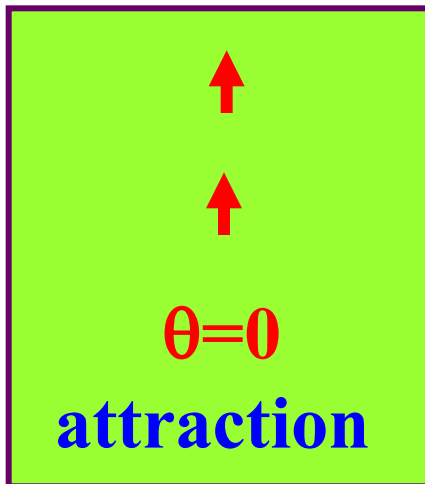
One pion exchange → Tensor force

How does the tensor force work ?

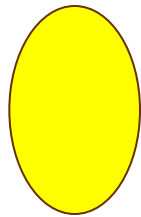
Spin of each nucleon \uparrow is parallel, because the total spin must be $S=1$

The potential has the following dependence on the angle θ with respect to the total spin \vec{S} .

$$V \sim Y_{2,0} \sim 1 - 3 \cos^2 \theta$$



Monopole effects due to the tensor force



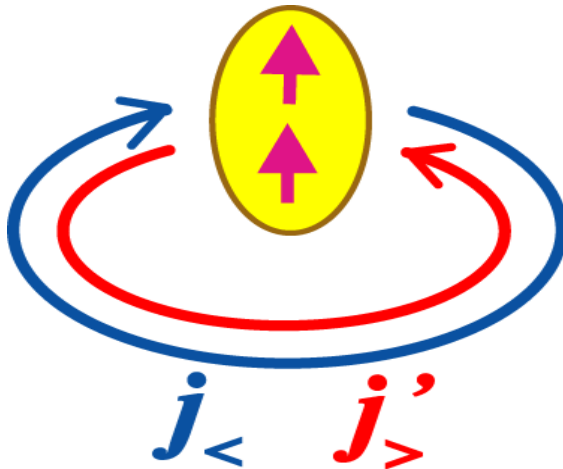
wave function of **relative motion**



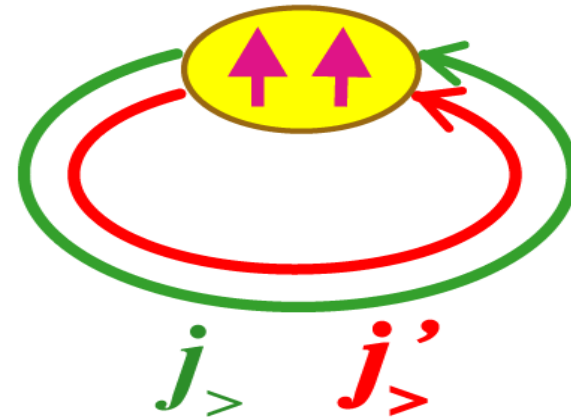
spin of nucleon

large relative momentum

small relative momentum



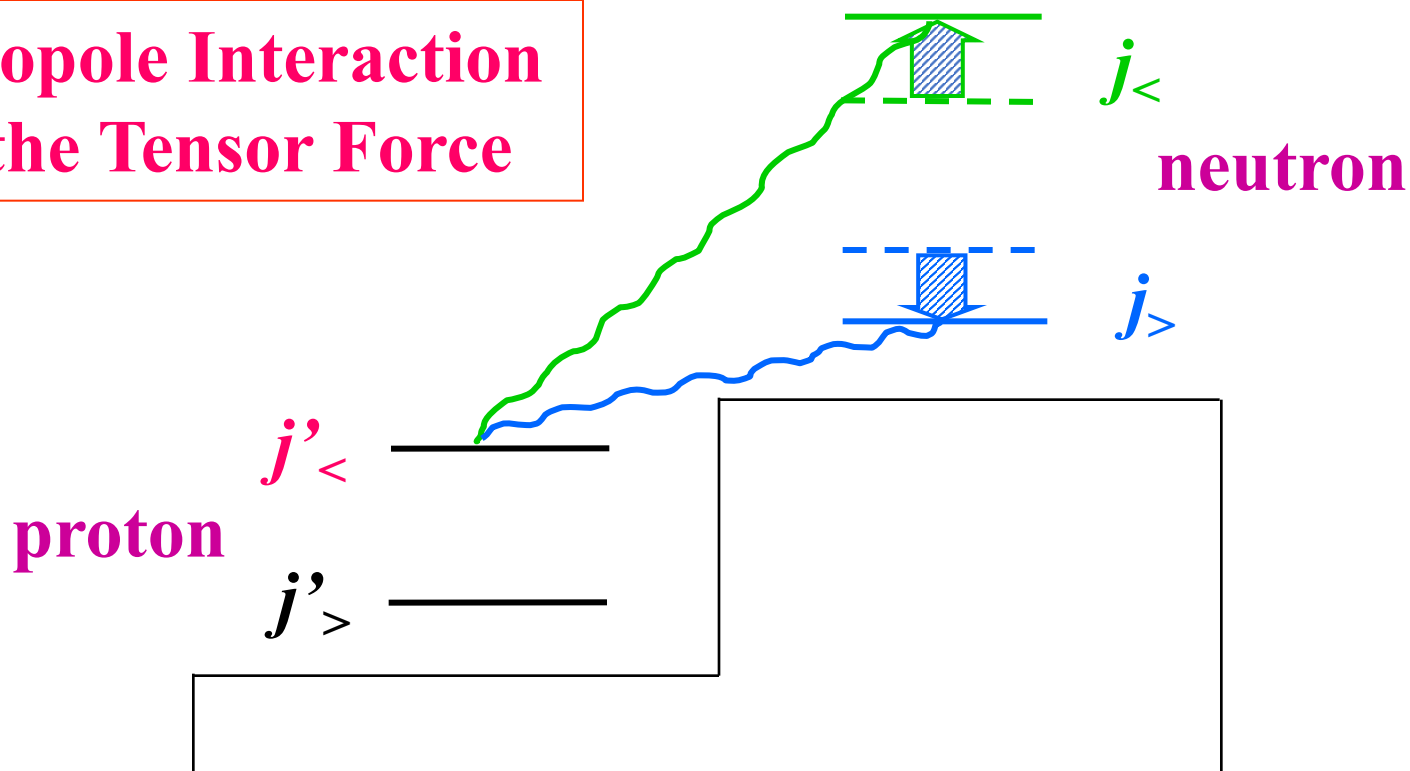
attractive



repulsive

$$j_{>} = l + \frac{1}{2}, \quad j_{<} = l - \frac{1}{2}$$

Monopole Interaction of the Tensor Force



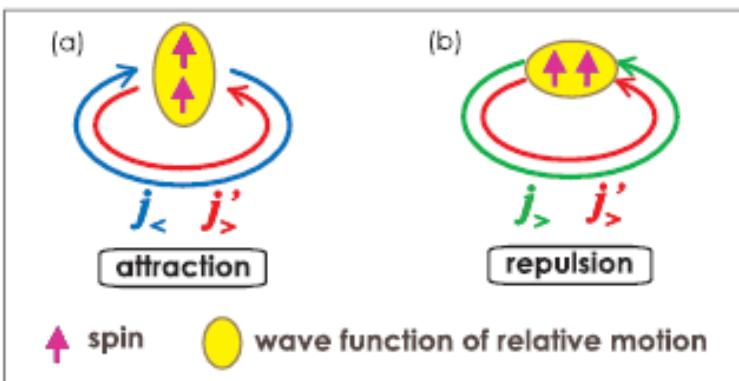
Identity for tensor monopole interaction

$$(2j_> + 1) v_{m,T}^{(j' j_>)} + (2j_< + 1) v_{m,T}^{(j' j_<)} = 0$$

$v_{m,T}$: monopole strength for isospin T

Variation of monopole matrix element from tensor force : $A=70$

variations 0.5 ~ 1 MeV



$\Delta n=0$

$\Delta n=1$

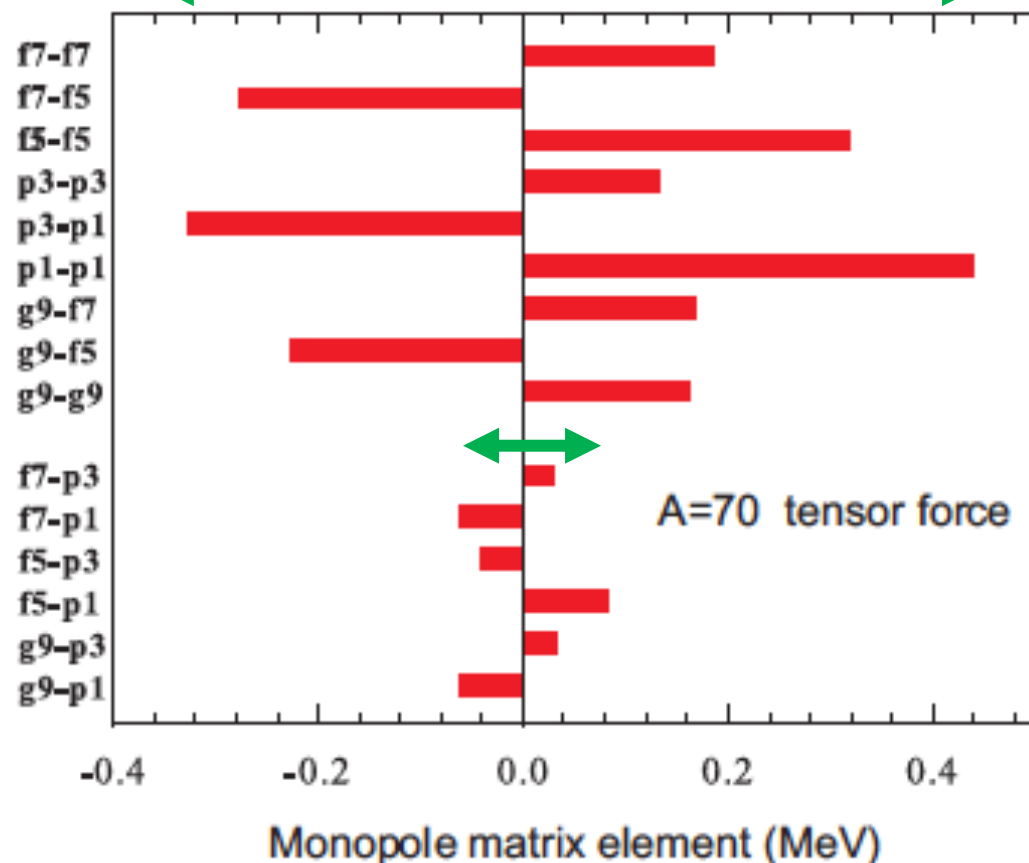


Figure 34 Monopole matrix elements of the tensor force in the $T=0$ channel. The orbit labeling is abbreviated like f7 for $1f_{7/2}$, etc. The orbits are from valence shell for $A = 70$.

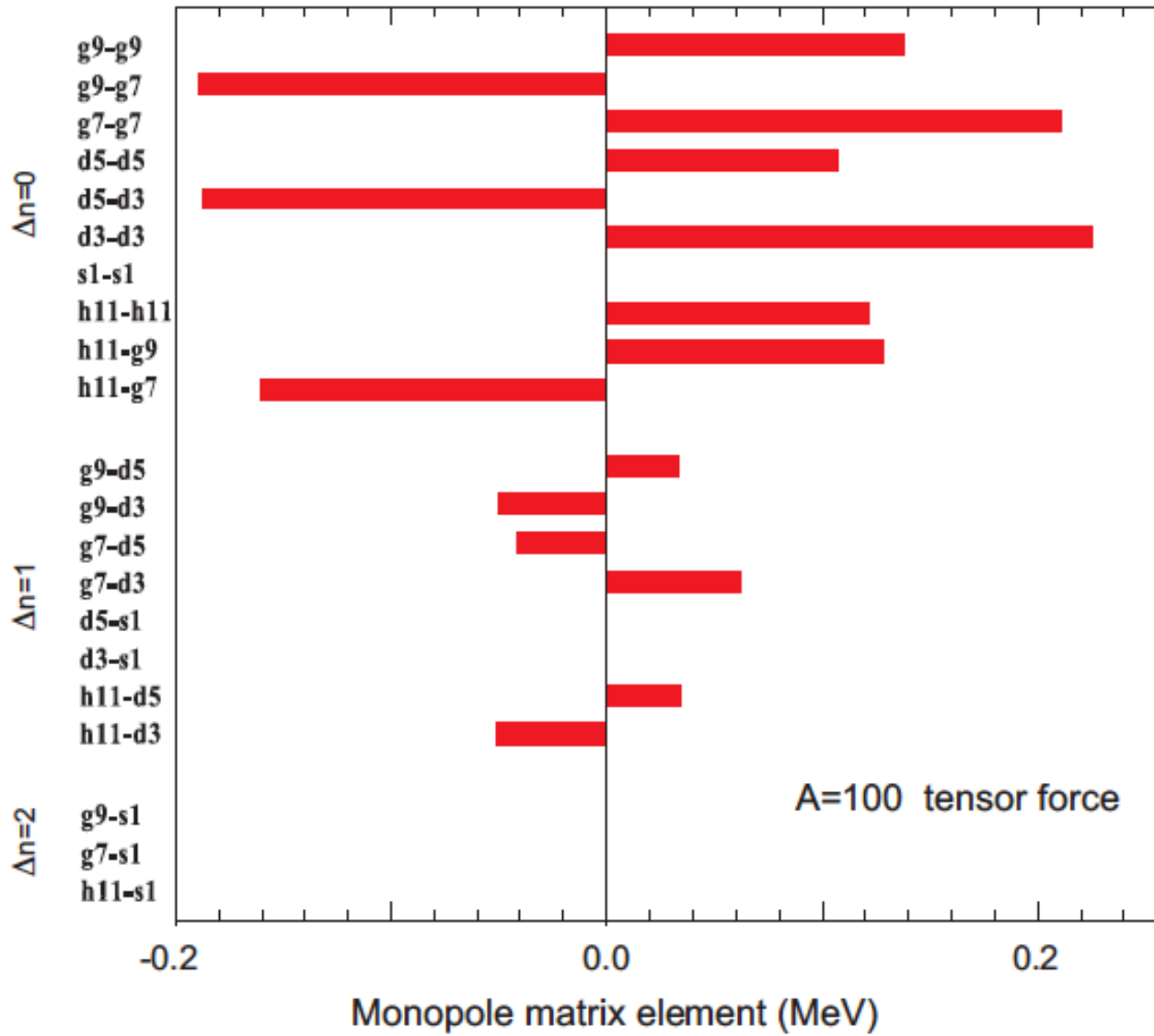
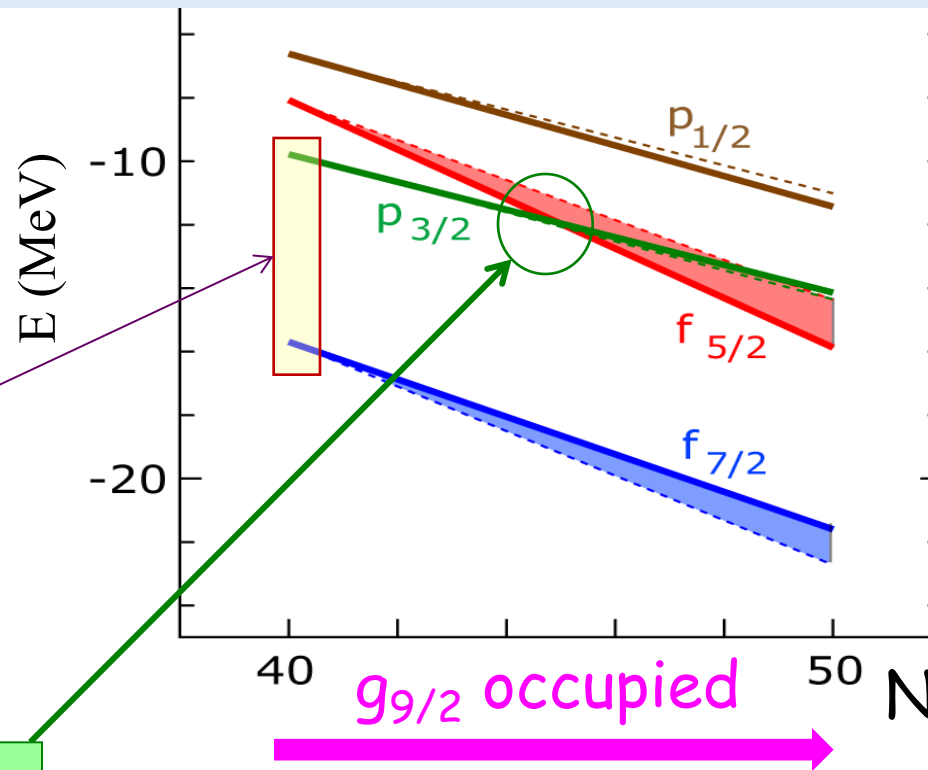


Figure 33 Monopole matrix elements of the tensor force in the $T=0$ channel. The orbit labeling is abbreviated like g9 for $1g_{9/2}$, etc. The orbits are from valence shell for $A = 100$.

Proton single-particles levels of Ni isotopes



Central Gaussian
+ Tensor

solid line:
full effect

dotted line:
central only

shaded area :
effect of
tensor force

PRL 103, 142501 (2009)

PHYSICAL REVIEW LETTERS

week ending
2 OCTOBER 2009

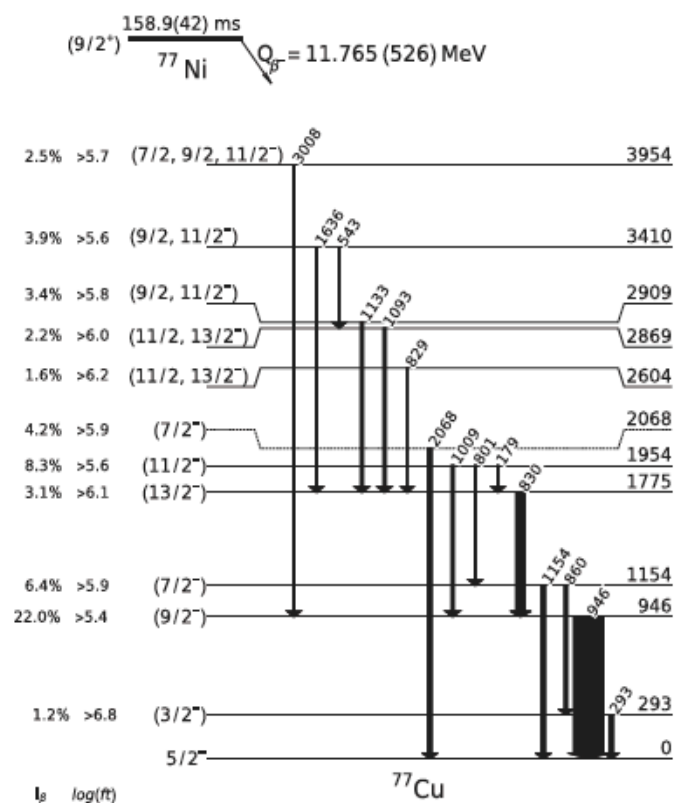
Nuclear Spins and Magnetic Moments of $^{71,73,75}\text{Cu}$: Inversion of $\pi 2p_{3/2}$ and $\pi 1f_{5/2}$ Levels in ^{75}Cu

K. T. Flanagan,^{1,2} P. Vingerhoets,¹ M. Avgoulea,¹ J. Billowes,³ M. L. Bissell,¹ K. Blaum,⁴ B. Cheal,³ M. De Rydt,¹ V. N. Fedosseev,⁵ D. H. Forest,⁶ Ch. Geppert,^{7,8} U. Köster,¹⁰ M. Kowalska,¹¹ J. Krämer,⁹ K. L. Kratz,⁹ A. Krieger,⁹ E. Mané,³ B. A. Marsh,⁵ T. Matema,¹⁰ L. Mathieu,¹² P. L. Molkanov,¹³ R. Neugart,⁹ G. Neyens,¹ W. Nörtershäuser,^{9,7} M. D. Seliverstov,^{13,16} O. Serot,¹² M. Schug,⁴ M. A. Sjoedin,¹⁷ J. R. Stone,^{14,15} N. J. Stone,^{14,15} H. H. Stroke,¹⁸ G. Tungate,⁶ D. T. Yordanov,⁴ and Yu. M. Volkov¹³

¹Instituut voor Kern- en Stralingsfysica, Katholieke Universiteit Leuven, B-3001 Leuven, Belgium

Shell Evolution towards ^{78}Ni : Low-Lying States in ^{77}Cu

E. Sahin,^{1,*} F. L. Bello Garrote,¹ Y. Tsunoda,² T. Otsuka,^{2,3,4,5} G. de Angelis,⁶ A. Görgen,¹ M. Niikura,³ S. Nishimura,⁷



above ~ 1 MeV. A rather large $Z = 28$ gap is consistent with this, while the gap decreases modestly by ~ 2 MeV from $N = 40$ to 50.

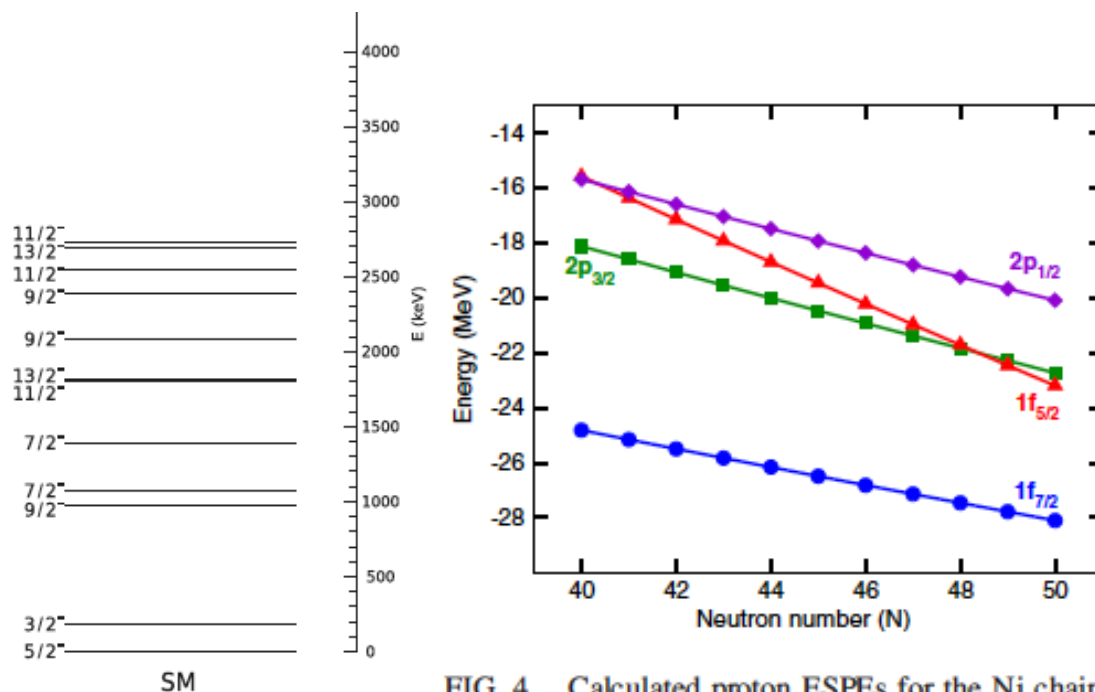
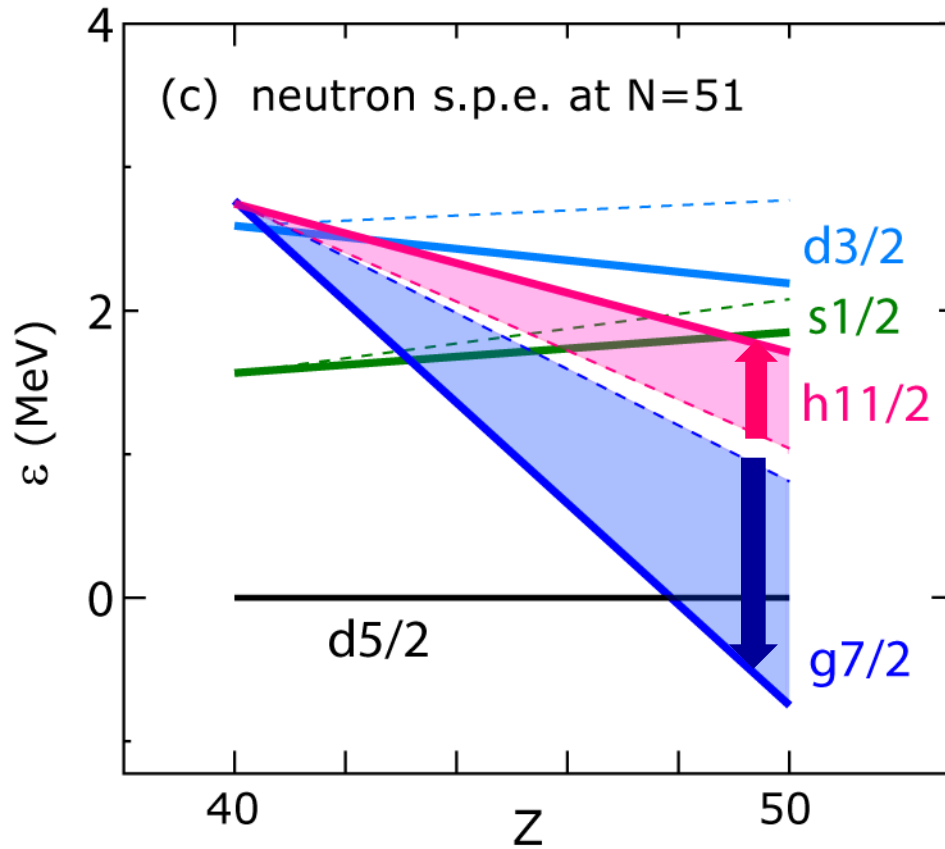


FIG. 4. Calculated proton ESPEs for the Ni chain.

Shell structure of a key nucleus ^{100}Sn



solid line : full
(central + tensor)

dashed line : central only
Fedderman-Pittel (1977)

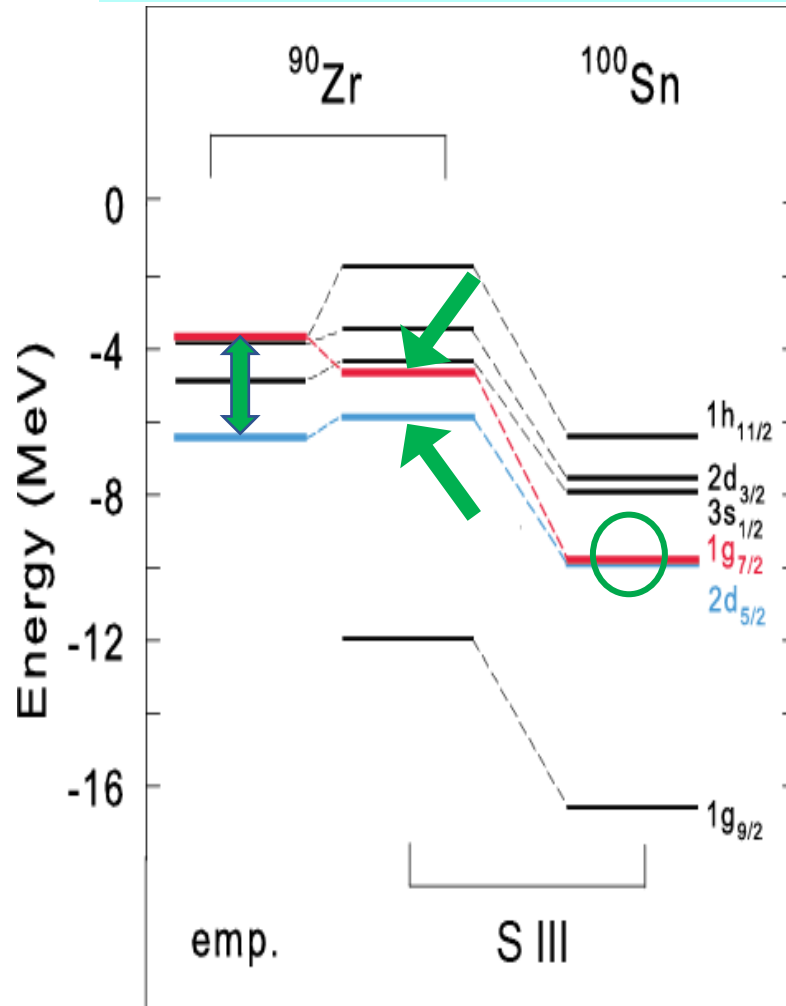
shaded area :
effect of tensor force

Exp. $d5/2$ and $g7/2$ should be close
Seweryniak et al.

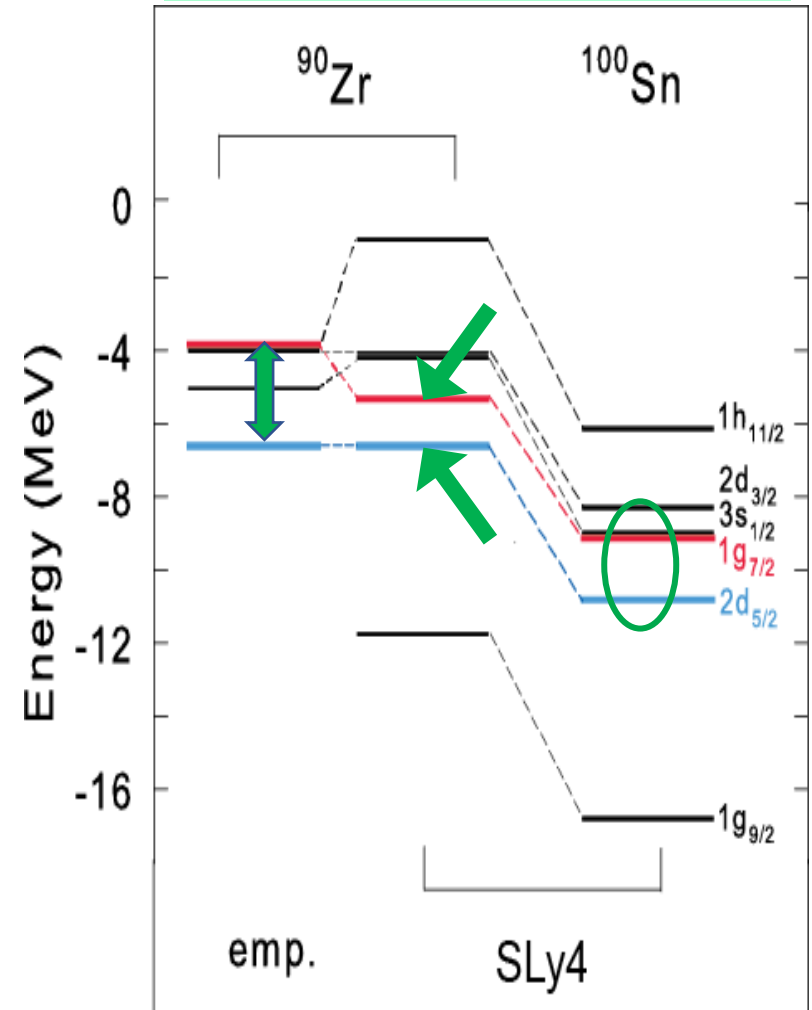
Phys. Rev. Lett. 99, 022504 (2007)
Gryzywacz et al.

Predictions by Skyrme model (HF calculation)

$g_{7/2}$ lowered relative to $d_{5/2}$
by ~ 2 MeV (\leftrightarrow previous page)



$g_{7/2}$ raised relative to $d_{5/2}$
by ~ 2 MeV (opposite)



An example with $_{51}\text{Sb}$ isotopes

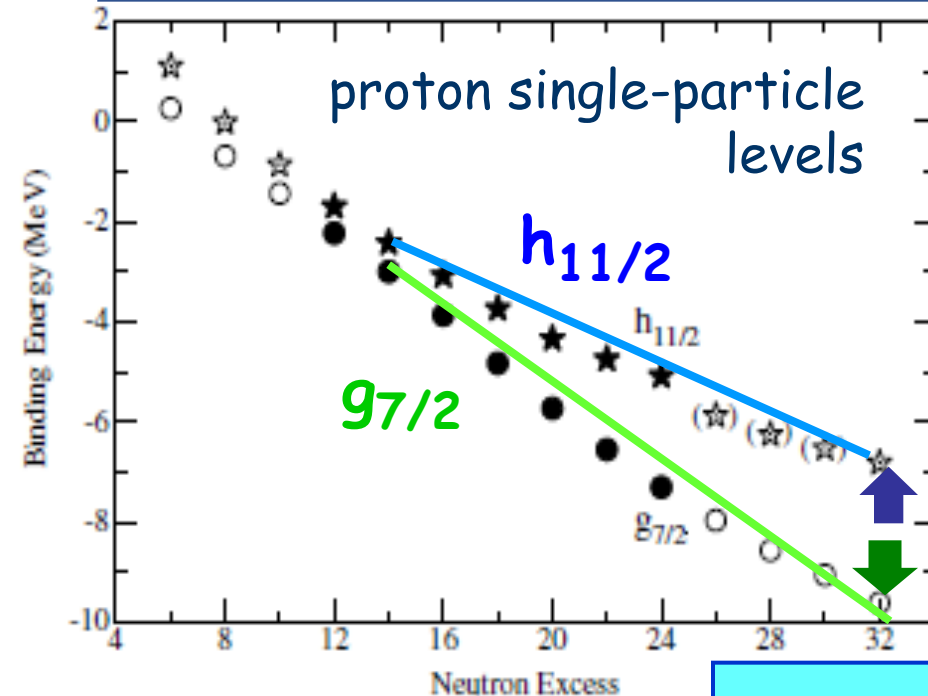
VOLUME 92, NUMBER 16

PHYSICAL REVIEW LETTERS

week ending
23 APRIL 2004

Is the Nuclear Spin-Orbit Interaction Changing with Neutron Excess?

J. P. Schiffer,¹ S. J. Freeman,^{1,2} J. A. Caggiano,³ C. Deibel,³ A. Heinz,³ C.-L. Jiang,¹ R. Lewis,³ A. Parikh,³ P. D. Parker,³
K. E. Rehm,¹ S. Sinha,¹ and J. S. Thomas⁴



$Z=51$ isotopes

change driven
by neutrons in $1h_{11/2}$

$h_{11/2} - h_{11/2}$ repulsive \uparrow

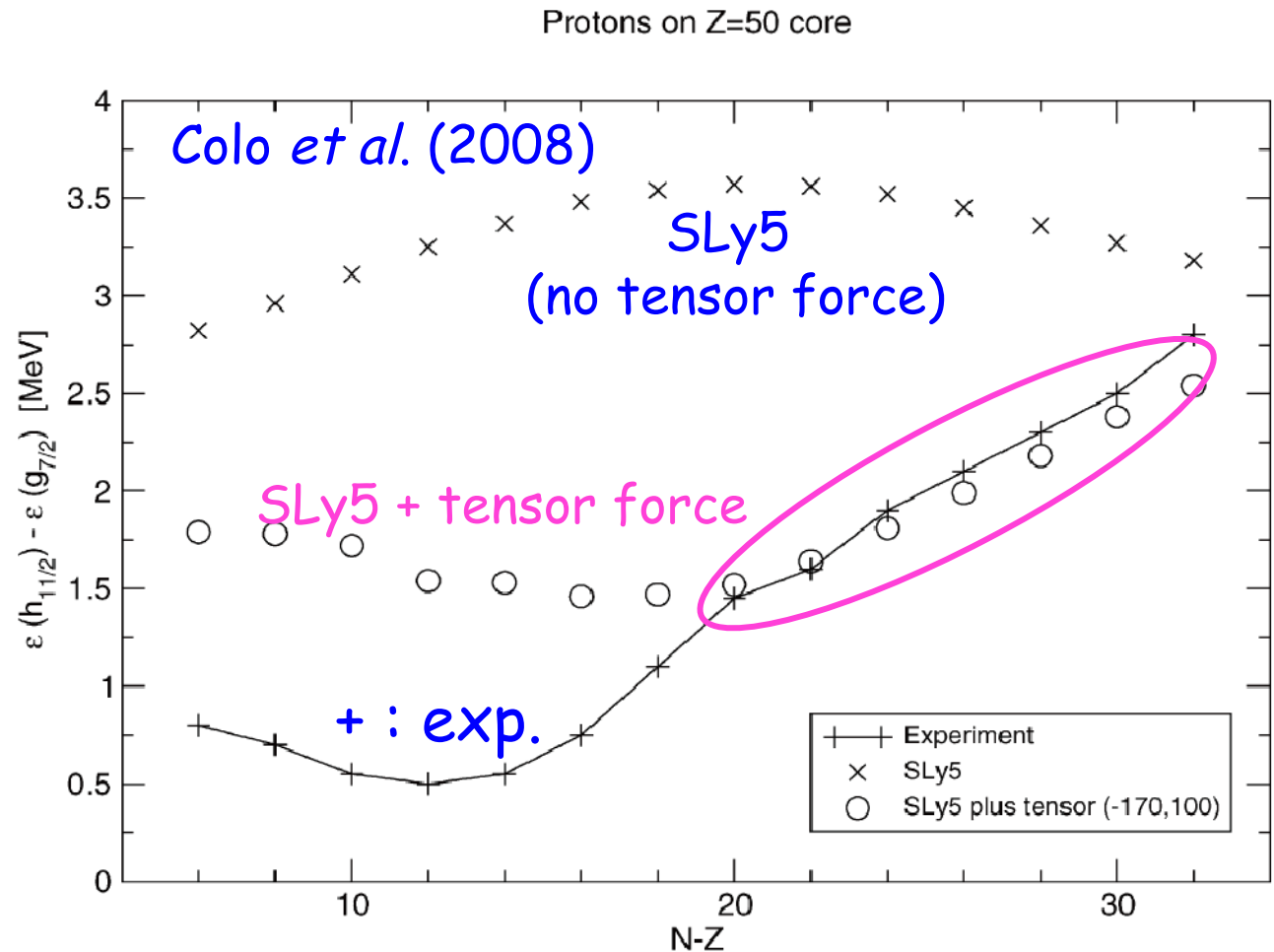
$h_{11/2} - g_{7/2}$ attractive \downarrow

$\pi + \rho$ meson exchange tensor force
(splitting increased by ~ 2 MeV)

No mean field theory,
(Skyrme, Gogny, RMF)
explained this before.

TO *et al.*, PRL 95, 232502 (2005)

Tensor force effect added to the Skyrme (mean-field) model :
Proton $g_{7/2}$ and $h_{11/2}$ single-particle orbits
on the $Z=50$ core with $N=64-82$



New magic number ?

Eigenvalues of
HO potential

Magic numbers
Mayer and
Jensen (1949)

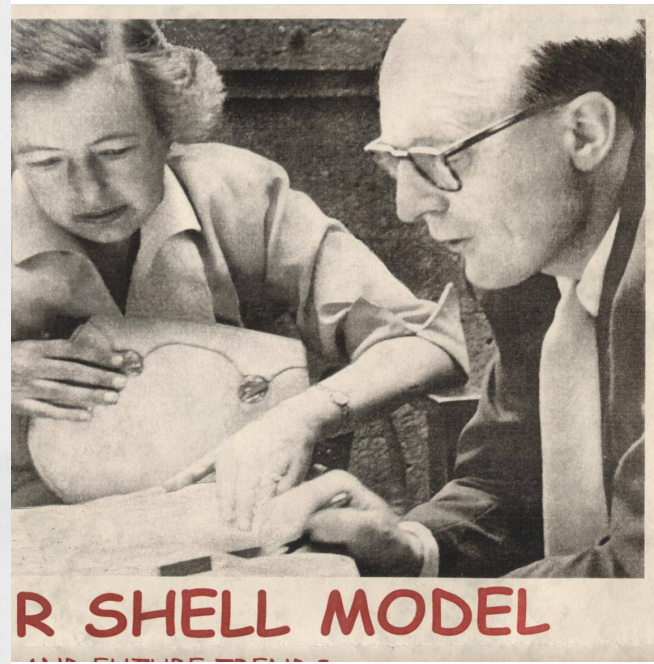
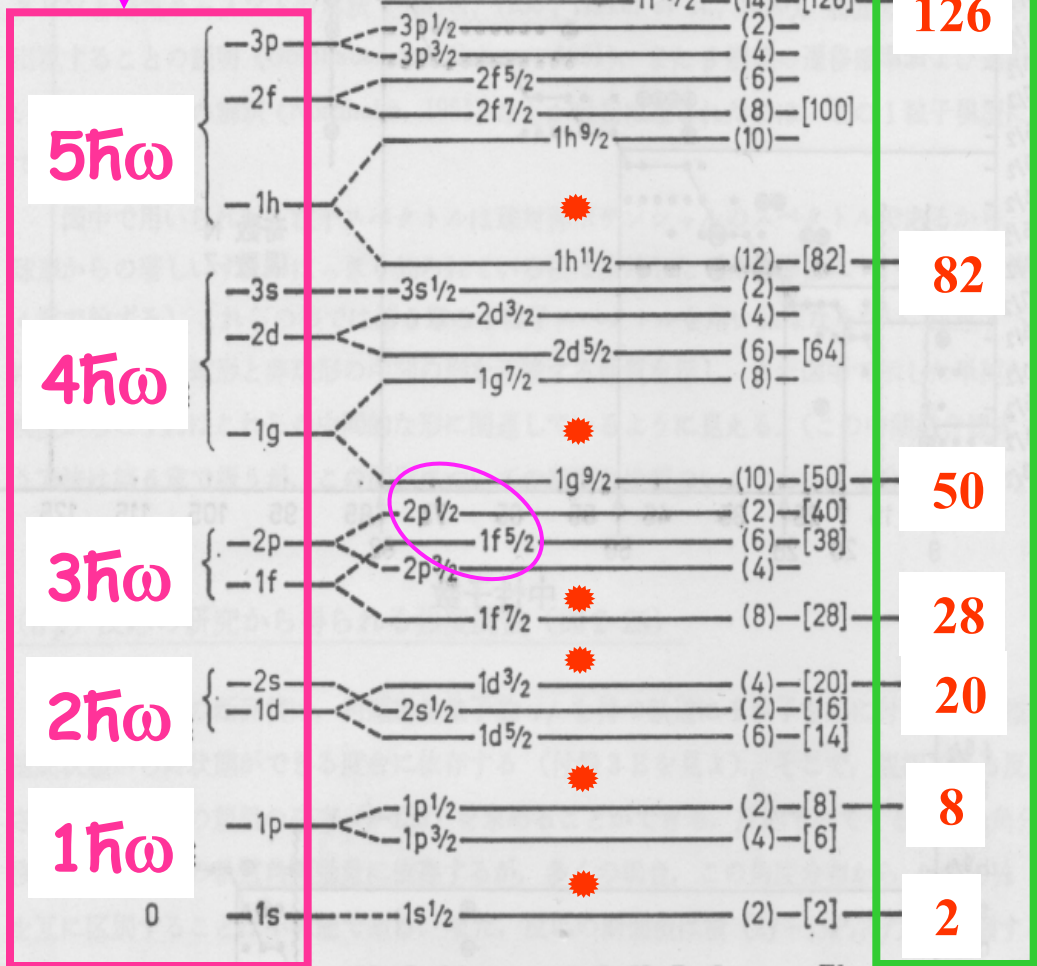
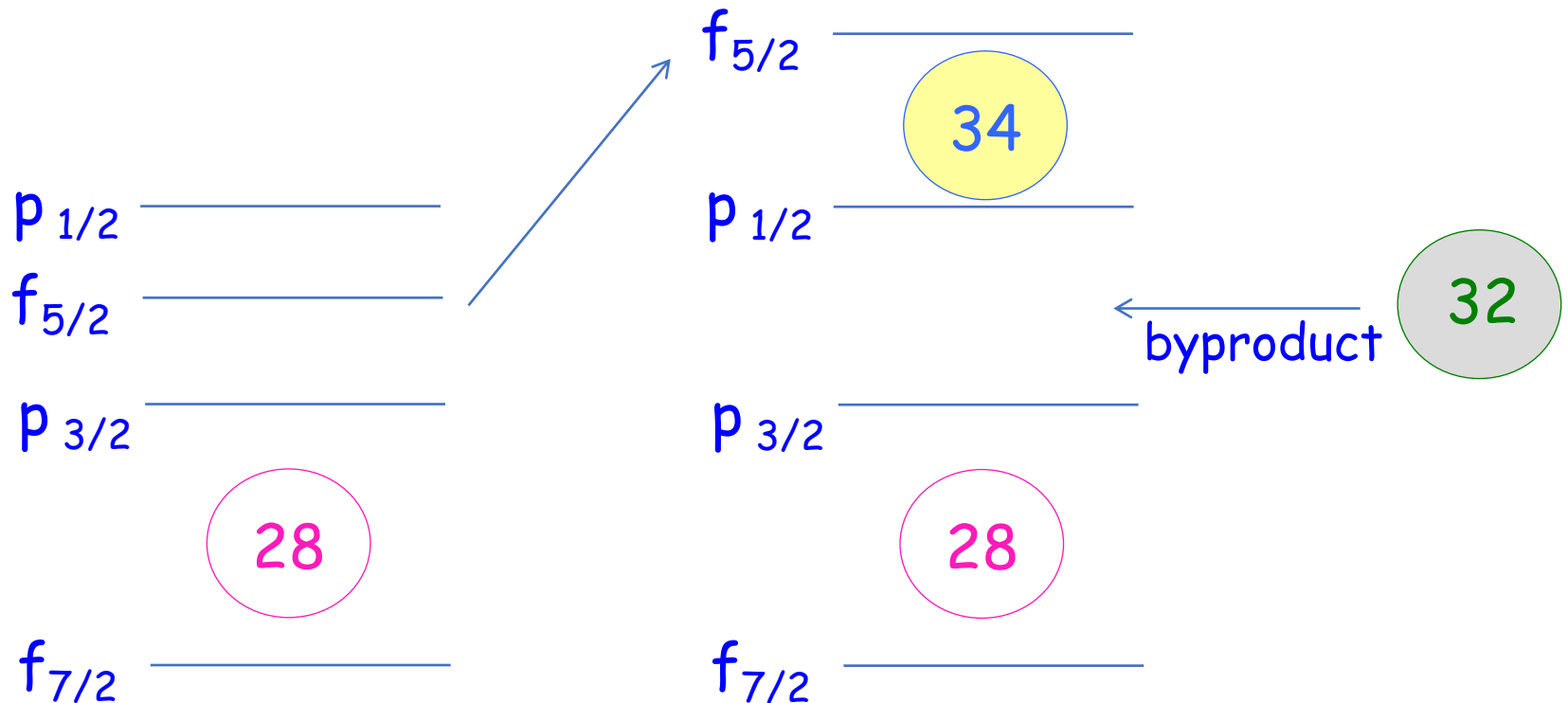


図 2-23 1 粒子軌道の順序. 図は M. G. Mayer and J. H. D. Jensen, *Elementary Theory of Nuclear Shell Structure*, p. 58, Wiley, New York, 1955 からとった.

Basic picture

shell structure
for **neutrons**
in **Ni** isotopes
($f_{7/2}$ fully occupied)

N=34 magic number may appear
if proton $f_{7/2}$ becomes vacant (**Ca**)
($f_{5/2}$ becomes less bound)



Appearance of new magic number N=34

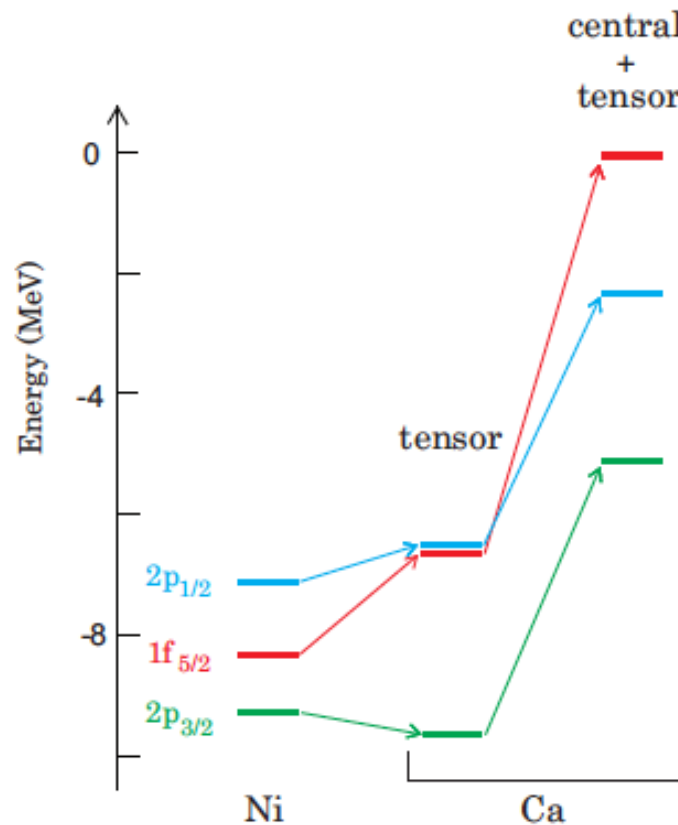


Figure 41 Change of single-particle energies from ^{56}Ni to ^{48}Ca .

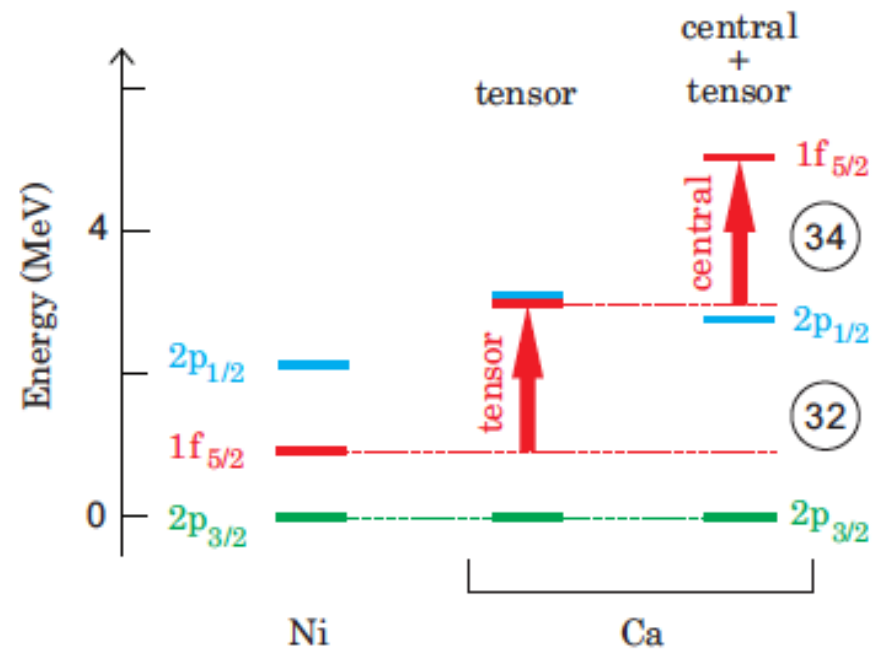


Figure 42 Change of single-particle energies from ^{56}Ni to ^{48}Ca relative to the $2p_{3/2}$ orbit. The red arrows indicate the change of the $1f_{5/2}$ ESPE. The arising magic numbers, $N=32$ and 34 , are shown in black circles.

Is there $N=34$ magic number ?

In comparison to $N=32$ magic number known experimentally for nearly 30 years.

Moving back to heavier nuclei, from the strong interaction in Fig. 1(c), we can predict other magic numbers, for instance, $N = 34$ associated with the $0f_{7/2}-0f_{5/2}$ interaction. In heavier nuclei, $0g_{7/2}$, $0h_{9/2}$, etc. are shifted upward in neutron-rich exotic nuclei, disturbing the magic numbers $N = 82, 126$, etc. It is of interest how the r process of nucleosynthesis is affected by it.

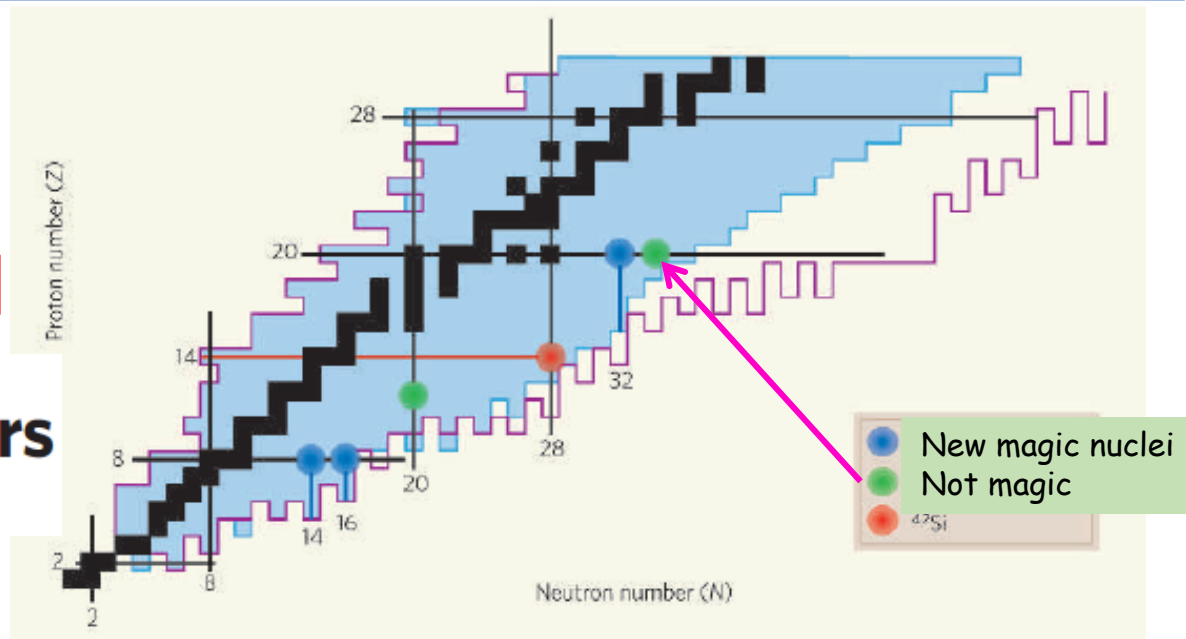
TO et al.
PRL 87 (2001)

NATURE|Vol 435|16 June 2005

NUCLEAR PHYSICS

Elusive magic numbers

Robert V.F. Janssens



Experiment @ RIBF → Finally confirmed

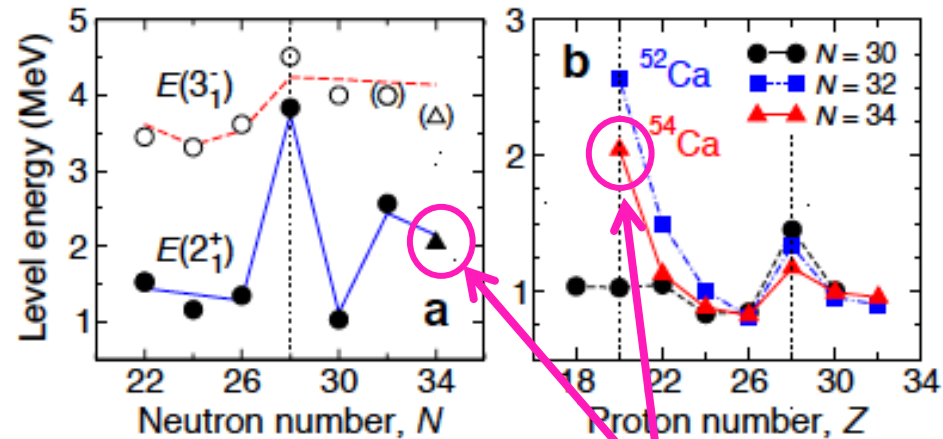
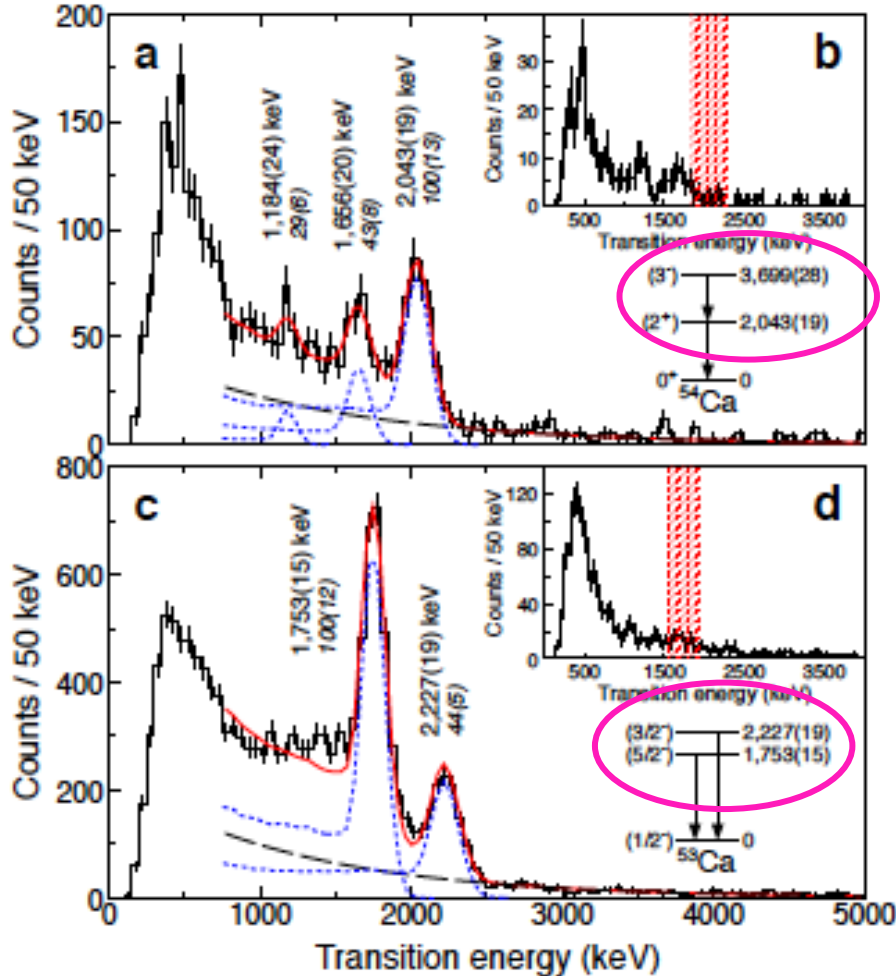


FIG. 4: Systematics of excited-state energies for even-even Ca isotopes and neighbouring nuclei. **a**, Level energies of first 2^+ (closed symbols) and 3^- (open symbols) states for even-even $^{42-54}\text{Ca}$ isotopes [28]. The results of the present study are indicated by triangular markers. Solid and dashed lines are shell-model predictions of $E(2_1^+)$ and $E(3_1^-)$, respectively (see text for details). Tentative spin-parity assignments are enclosed by parentheses. **b**, $E(2_1^+)$ along the $N = 30, 32$ and 34 isotonic chains. The solid and dashed lines are intended to guide the eye. Vertical dotted lines represent the traditional magic numbers in both plots.

new
RIBF
data

er-corrected γ -ray energy spectra. De-excitation γ rays measured in coinci-

^{54}Ca and **c**, ^{53}Ca reaction products. Peaks a

ive intensities are indicated by italic fonts. The short-blue and long-black dashed

Steppenbeck *et al.* Nature, 502, 207 (2013)

Shell evolution
with the modern nuclear forces

shell model powered by modern nuclear forces

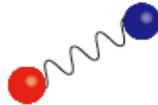
QCD



Lattice QCD

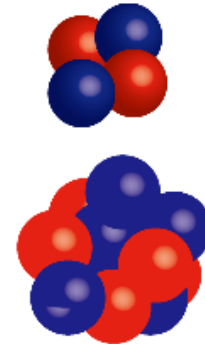
Effective Field Theory

Nuclear force

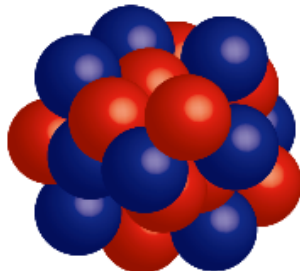


Few body techniques
No core shell model
and many others...

Light nuclei $\sim A \approx 10-20$



Medium mass nuclei $\sim A \approx 20-100$



shell model with core
via the **effective interaction**
derived from **nuclear force**

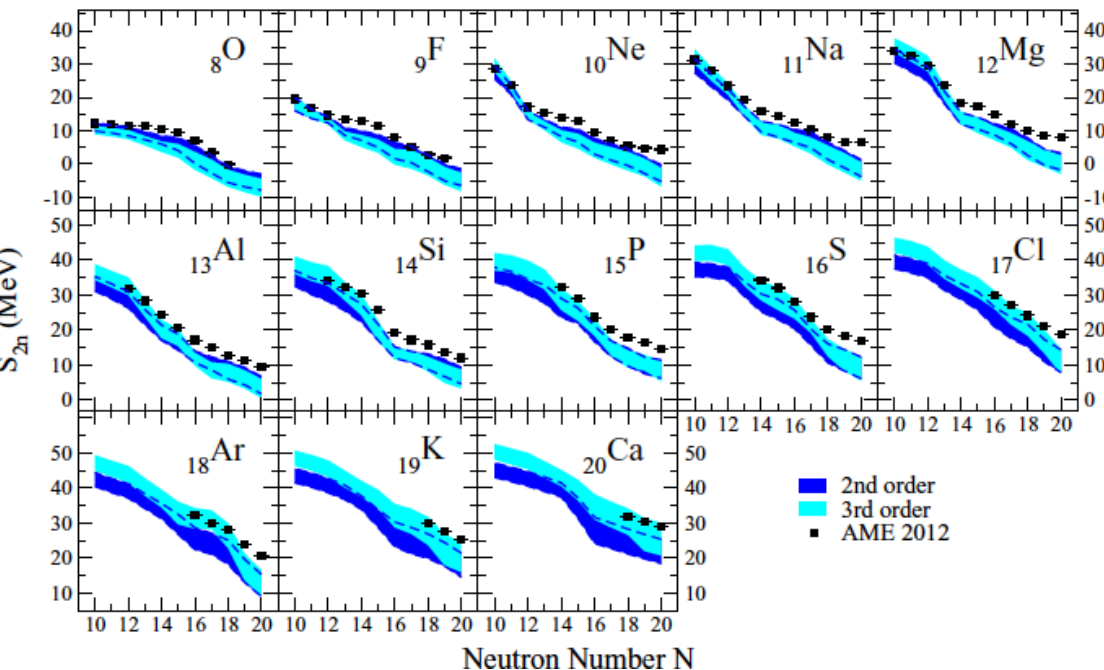
Courtesy from N. Tsunoda

Examples of structure calculations starting from chiral EFT forces

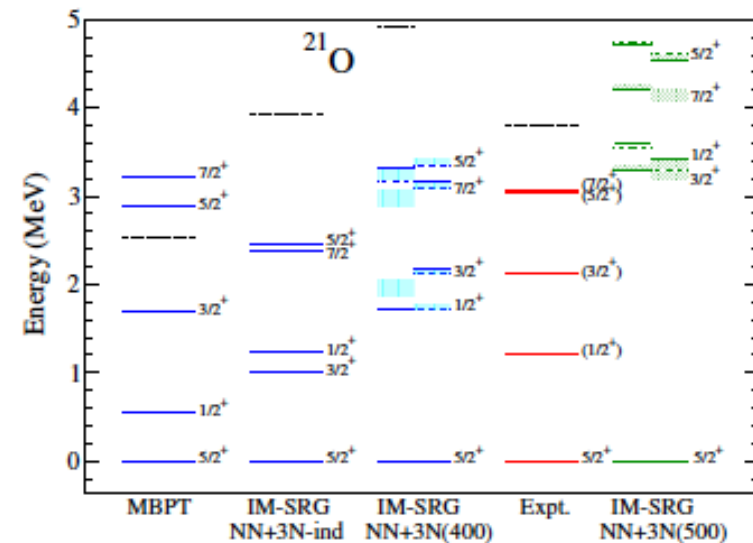
NN force: N3LO + 3N force: N2LO
 -> valence shell interaction

MBPT
 (Many-body Perturbation Theory)
 applicable to one major shell

IM-SRG
 (In-Medium Similarity
 Renormalization Group)



Simonis et al. PRC 93, 011302(R) (2016)



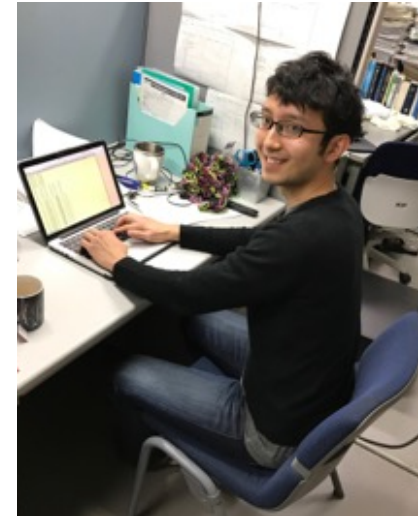
Bogner et al. PRL 113, 142501 (2014)

+ Coupled Cluster calculation + N2LOsat potential +

A recent development starting from *chiral EFT* + *3NF*

EKK method to handle consistently
two (or more) major shells

-> Effective shell-model interaction
(i) without fit of two-body m. e.,
(ii) applicable to broken magicity,
or fusion of two shells,
both are crucial for exotic nuclei.



PHYSICAL REVIEW C 95, 021304(R) (2017)

Exotic neutron-rich medium-mass nuclei with realistic nuclear forces

Naofumi Tsunoda,¹ Takaharu Otsuka,^{1,2,3,4} Noritaka Shimizu,¹ Morten Hjorth-Jensen,^{5,6}
Kazuo Takayanagi,⁷ and Toshio Suzuki⁸

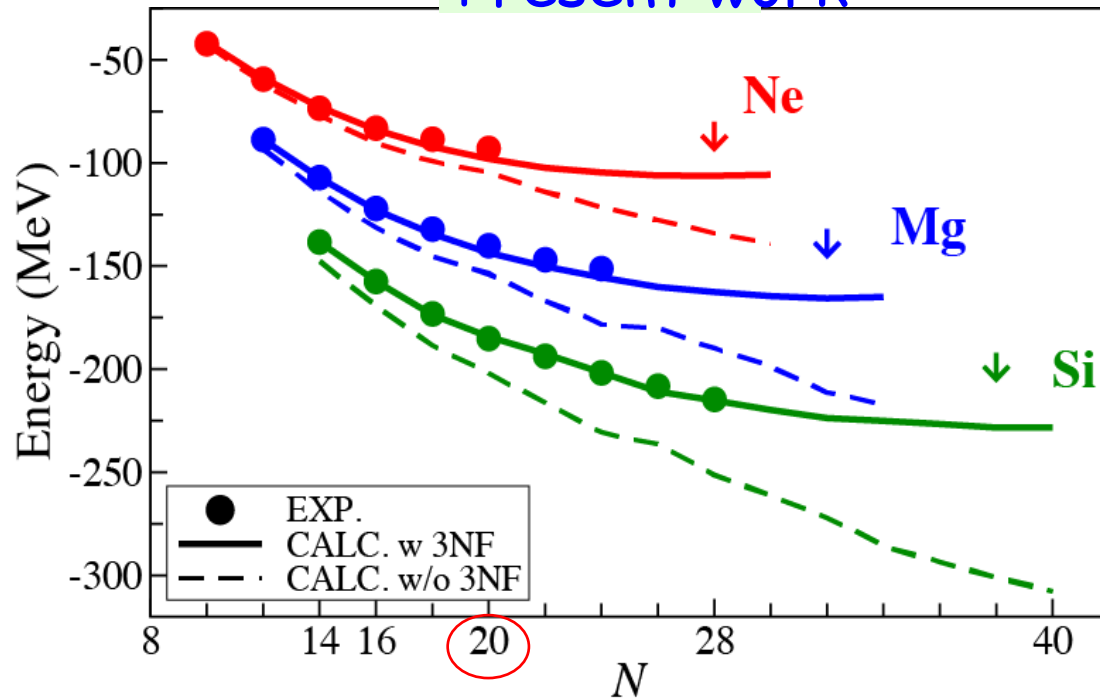
* E. M. Kren ciglowa and T. T. S. Kuo, Nucl. Phys. A 235, 171 (1974).

Re-visit to the "Island of Inversion" with ab initio TBMEs

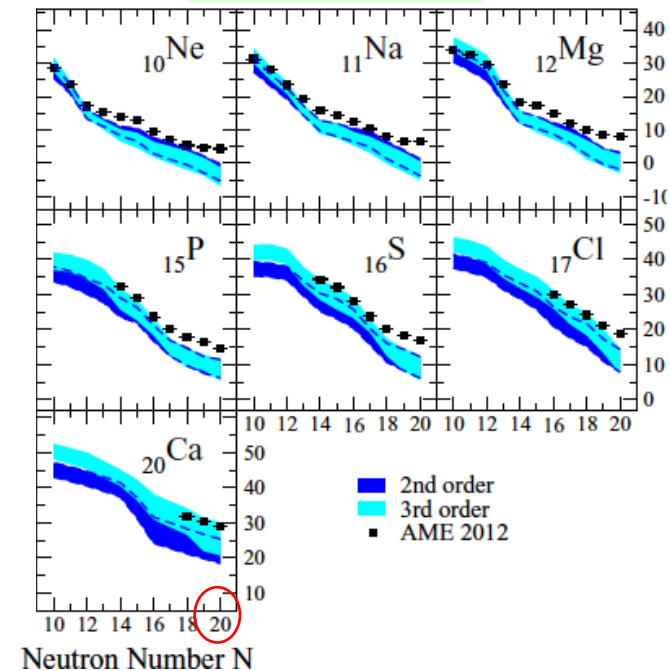
Calculations *with full $sd + pf$ shell*

ground-state energies

Present work



Earlier work



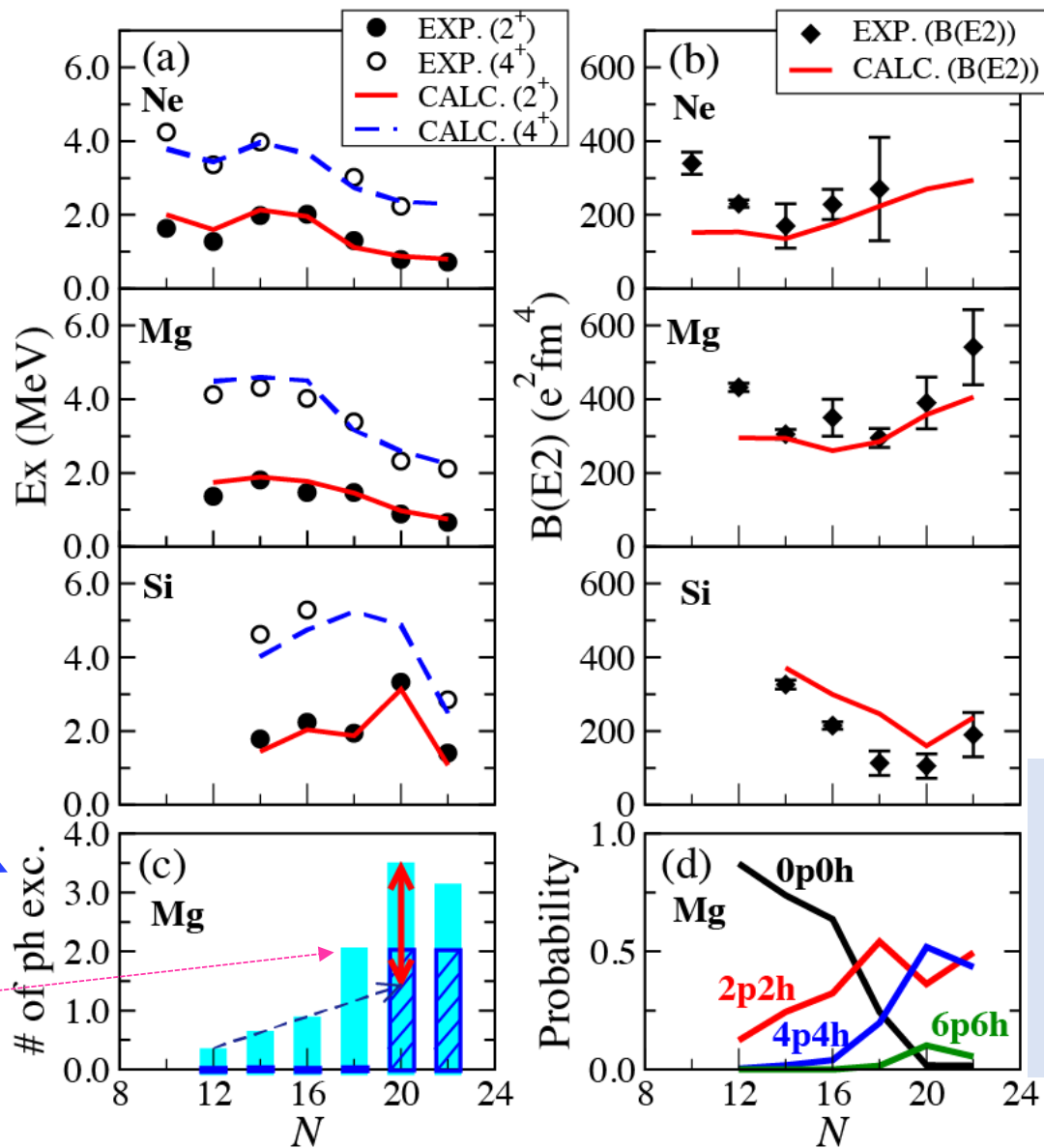
Simonis et al. PRC 93, 011302(R) (2016)

Ne-Mg-Si

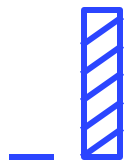
2⁺ & 4⁺ levels
and B(E2)
values

of particle-hole
excitations across
N=20 gap :
(modest) steady
increase

+
abrupt increase
after N=18

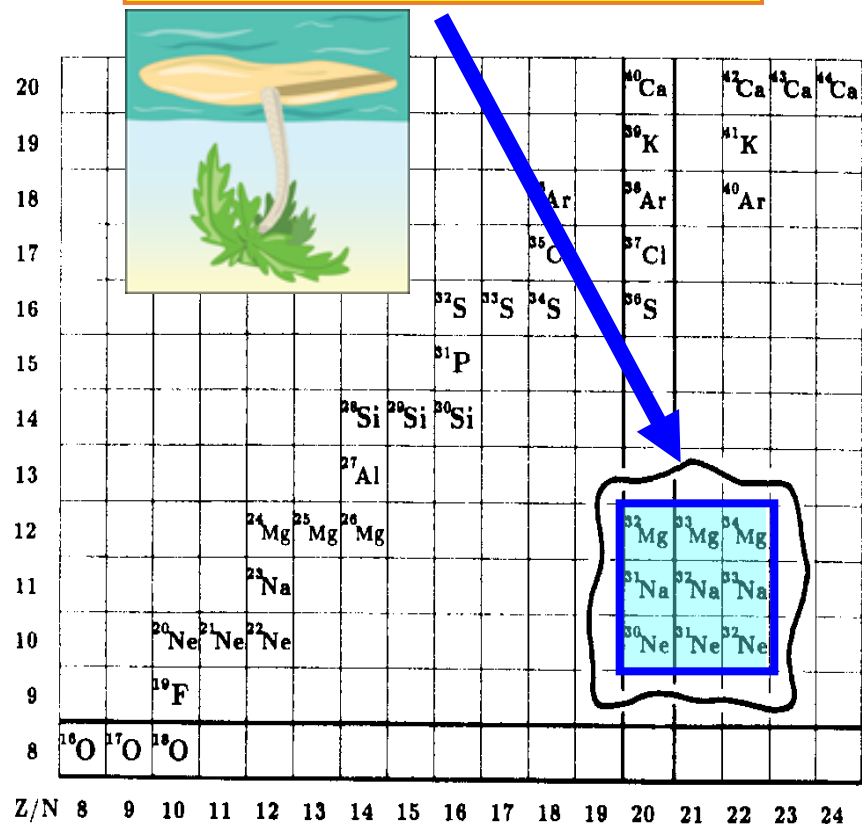


ground
states
of Mg
isotopes



Early idea of the Island of Inversion (WBB)
0p-0h or 2p-2h (discrete)

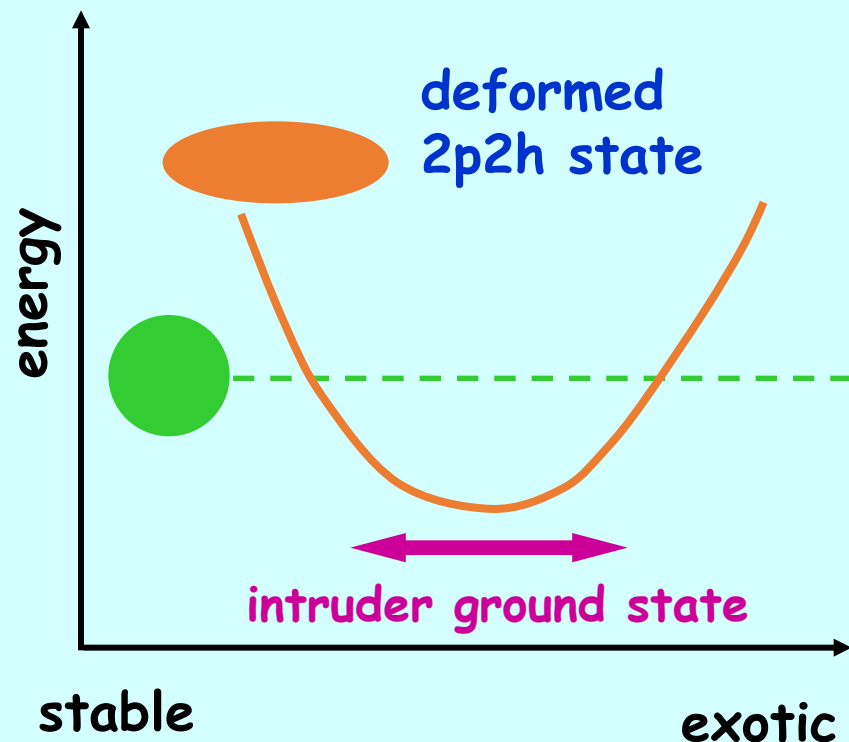
Island of Inversion



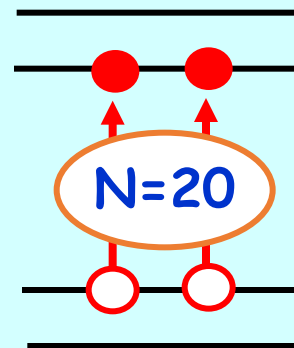
9 nuclei:
Ne, Na, Mg with N=20-22

Phys. Rev. C 41, 1147 (1990),
Warburton, Becker and
Brown

Basic picture was



pf shell



$^{31}_{12}\text{Mg}_{19}$: *very difficult to fit by the shell model*

$1/2^+$: not a simple $s_{1/2}$ single-particle state

mag. moment
 -0.88 exp
 -0.50 th
 -1.91 Schmidt

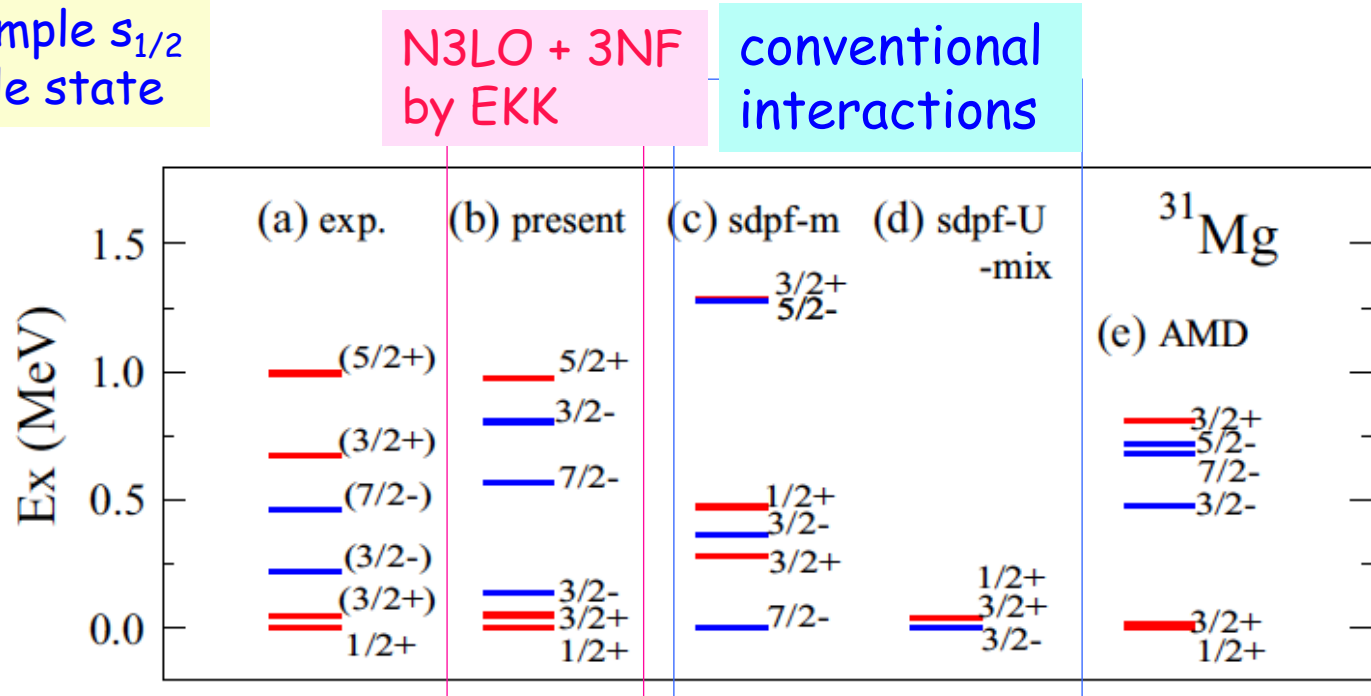


FIG. 4. Energy levels of ^{31}Mg . (a) experimental values, (b) present work, (c) sd-pf-m [16], (d) sd-pf-U-mix [17], and (e) AMD+ calculation [52].

exp. by laser spectroscopy : PRL 94, 022501 (2005), *G. Neyens, et al.*

Mixing between sd and pf shells is crucial

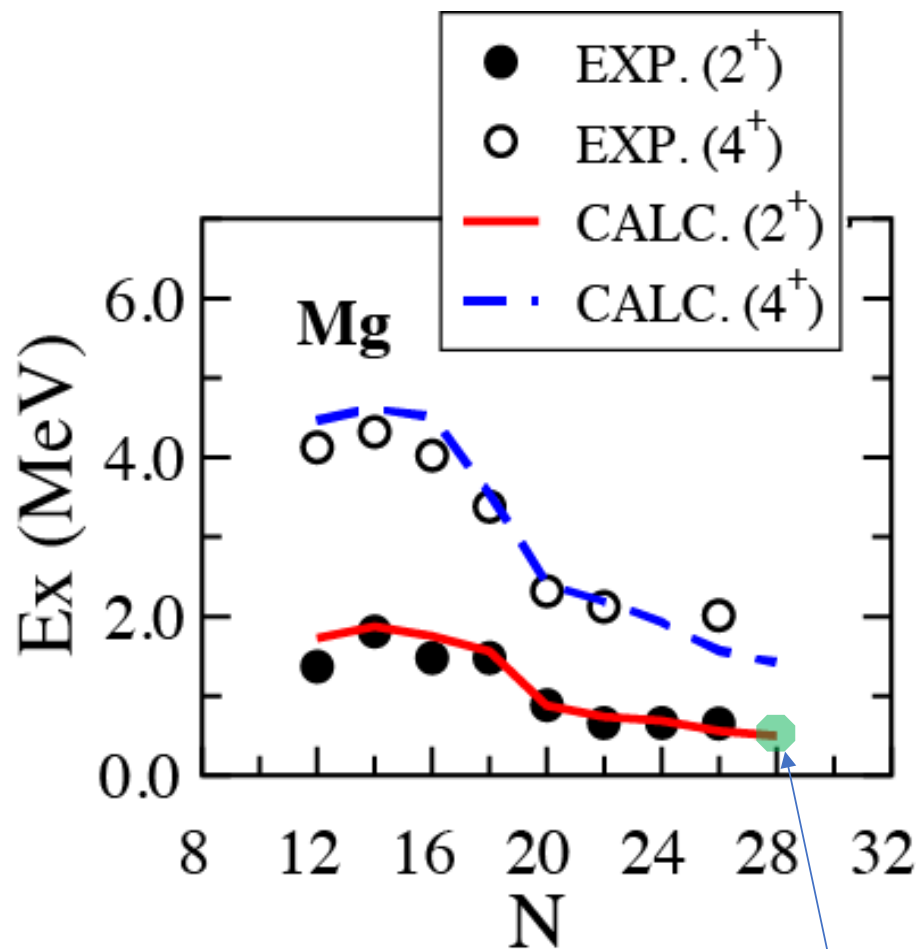
Fit of relevant TBME's is infeasible : too many TBME's but too few data

[16] Y. Utsuno, T. Otsuka, T. Mizusaki, and M. Honma, *Phys. Rev. C* **60**, 054315 (1999).

[17] E. Caurier, F. Nowacki, and A. Poves, *Phys. Rev. C* **90**, 014302 (2014).

[52] M. Kimura, *Phys. Rev. C* **75**, 041302 (2007).

Mg up to N=28 2⁺ & 4⁺ levels

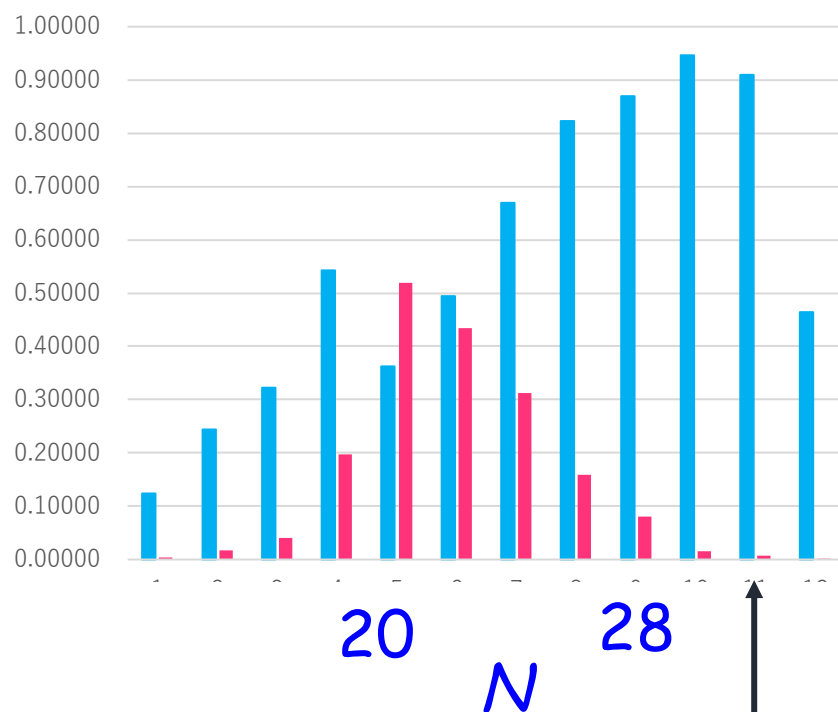


Recent data
Crawford *et al.*
(ARIS 2017 + priv. com.)

Probability of ph configurations over Z=N=20 (ground state)

2 ph excitations over N=20

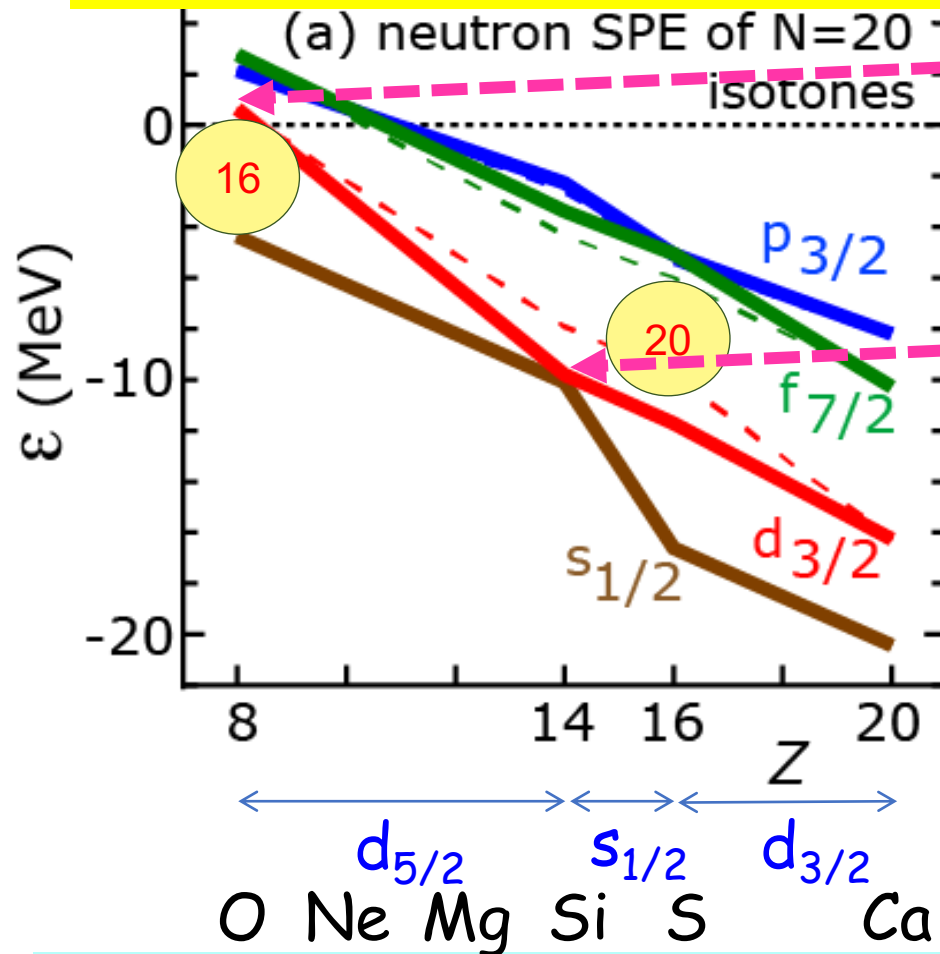
4 ph excitations over N=20



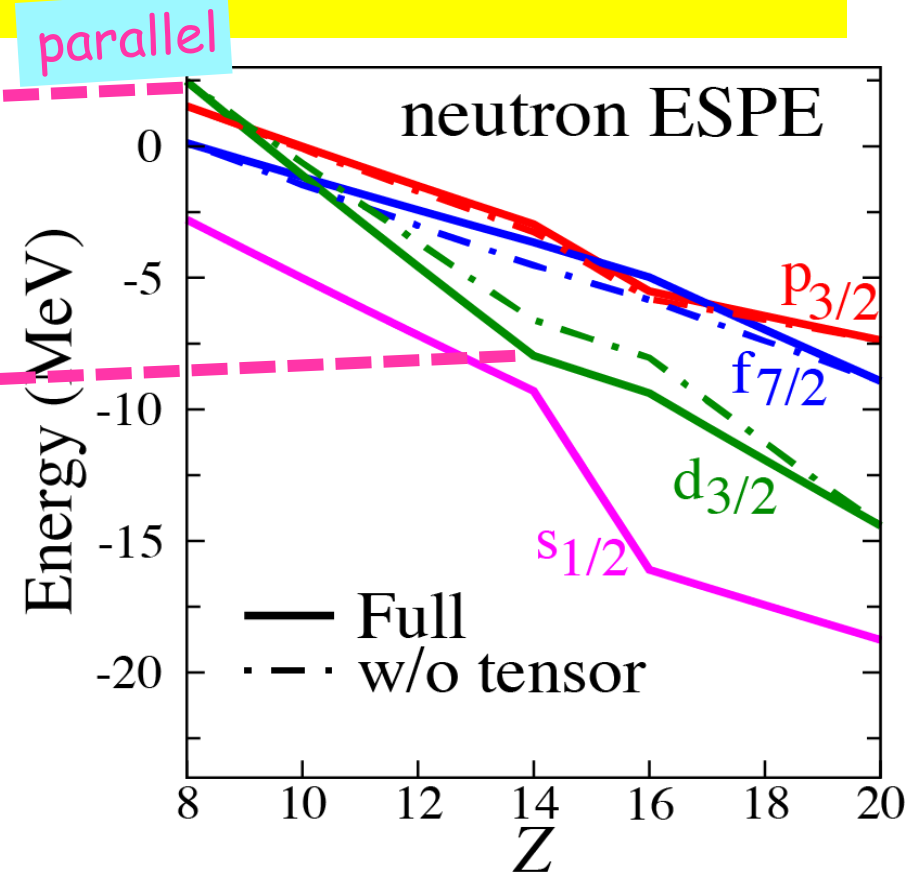
our
drip line

Neutron single-particle energy (SPE) at N=20

Tensor and central force effects are almost identical between the two schemes.



Obtained from VMU interaction :
phenomenological central (Gaussian)
and $\pi + \rho$ meson-exchange tensor interaction



Obtained by EKK method
from chiral EFT NN
interaction

The traditional picture such as the Island of Inversion is being re-visited and re-examined, leading to a renewed picture !

The shell structure can be indeed changed in many cases !

The same physics drives the shape coexistence and in the quantum phase transition. Tomorrow ...

July 14-19, 2017

*Recent developments
in
shell model studies of atomic nuclei*

Takaharu Otsuka

3rd lecture



This work has been supported by MEXT and JICFuS as a priority issue (Elucidation of the fundamental laws and evolution of the universe) to be tackled by using Post ‘K’ Computer

Summary of shell evolution

p-n monopole interaction between orbits j and j'

particularly strong, if $\Delta n = 0$ (n : # of nodes of radial wave func.)

central : attractive

tensor : attractive ($j_{>} - j'_{<}$) or repulsive ($j_{>} - j'_{>}$)

example : $g_{9/2} - f_{5/2}$: coherent $g_{9/2} - f_{7/2}$: cancellation

monopole interaction changes ESPE (effective single-particle energy) with linear dependence on the occupation number $n_{j'}$

→ effect can be magnified → shell evolution

change of spin-orbit splitting

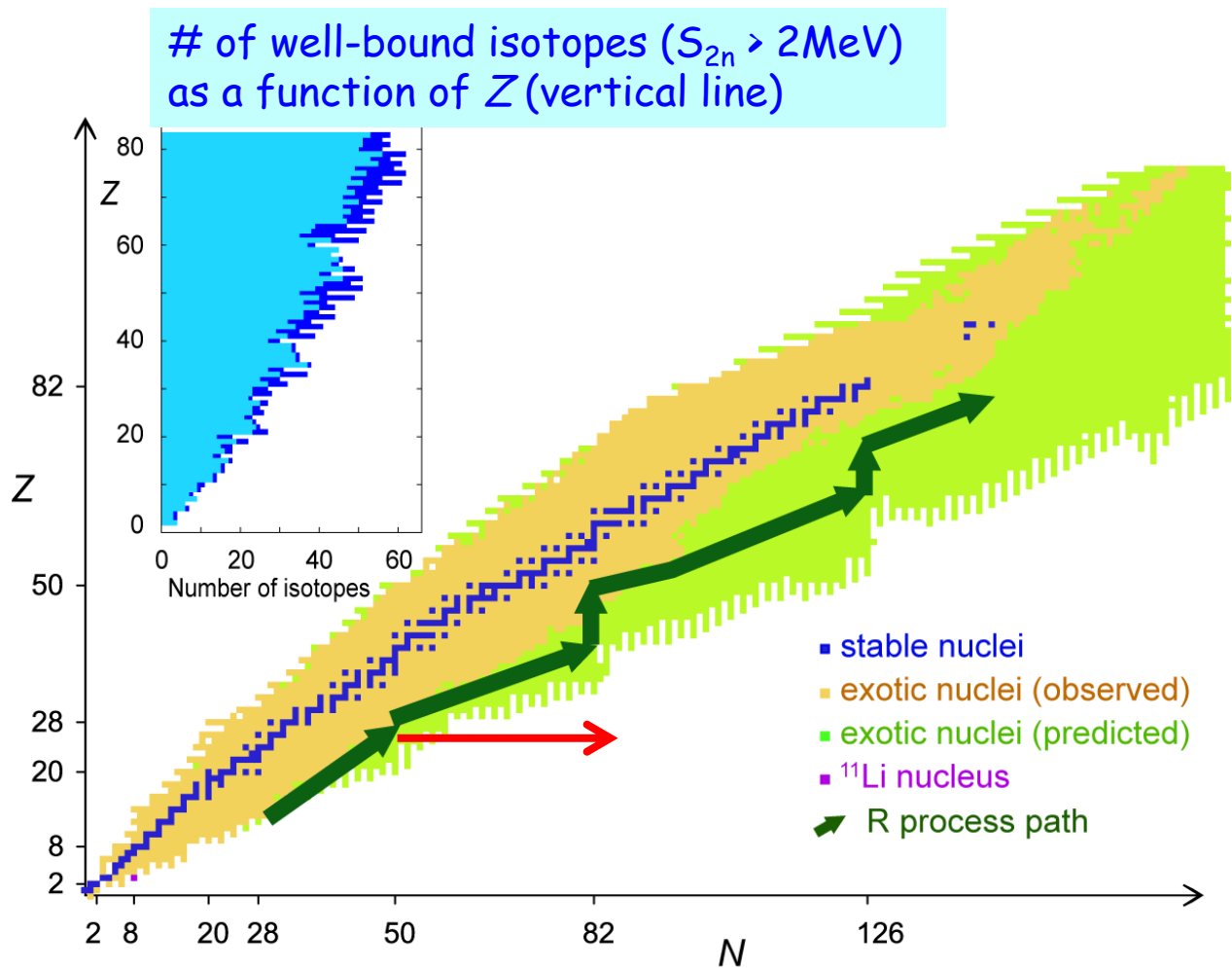
change of shell gap

change of the ordering of orbits

new magic numbers, disappearance of usual magic numbers

➡ Neutron drip line is quite far, and there are many well-bound exotic nuclei.

➡ Nuclear forces can play leading roles !



Effective single-particle energy

This is not uniform as a function of j and j' .

$$\Delta \hat{\epsilon}_j^p = \sum_{j'} V_{T=1}^m(j, j') \Delta \hat{n}_{j'}^p + \sum_{j'} \boxed{V_{pn}^m(j, j')} \Delta \hat{n}_{j'}^n, \quad \text{This can be large.}$$

Three pillars combined for future

computation

Monte Carlo
Shell Model
(MCSM)

(almost)
unlimited
dimensionality

massive
parallel
computers

Hamiltonian

pf
pfg9d5 (A3DA) (Ni)
8+8 on ^{56}Ni core (Zr)
8+8 on ^{80}Zr core (Sn)
8+10 on ^{132}Sn core
(Sm)

...
island of stability
+
 χ EFT based
multi-shell int.

many-body dynamics

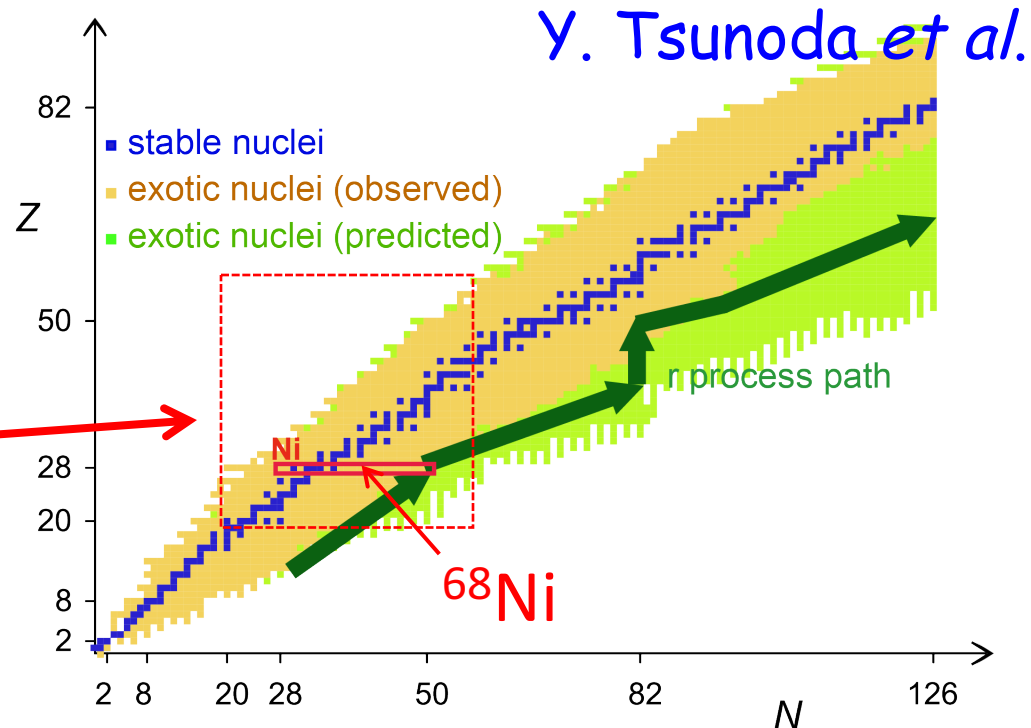
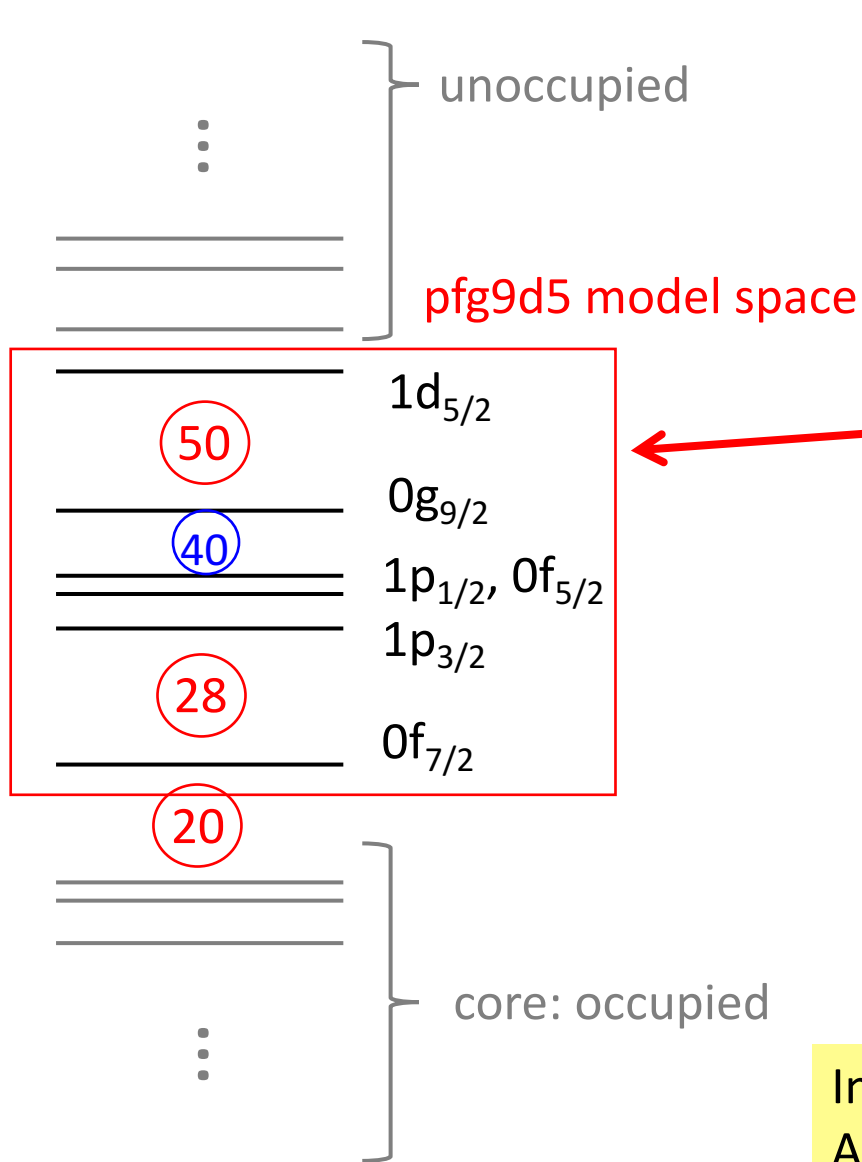
Shell evolution
(Type I & II)

Quantum Phase
Transition

Shape coexistence

Quantum
Self-organization

Monte Carlo Shell Model (MCSM) calculation on Ni isotopes



This model space is wide enough to discuss how **magic numbers 28, 50** and **semi-magic number 40** are visible or smeared out.

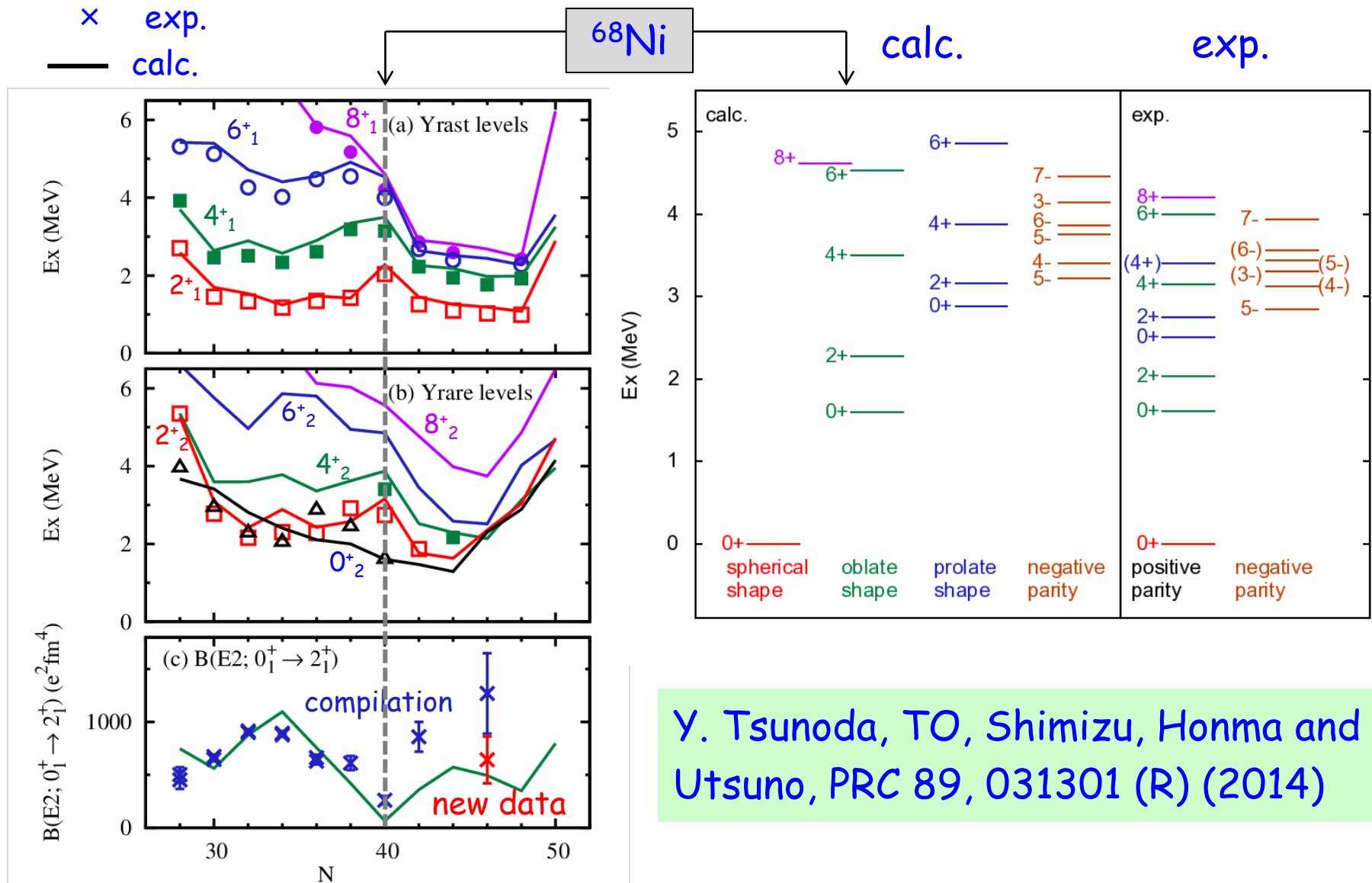
Interaction:

A3DA interaction is used with minor corrections

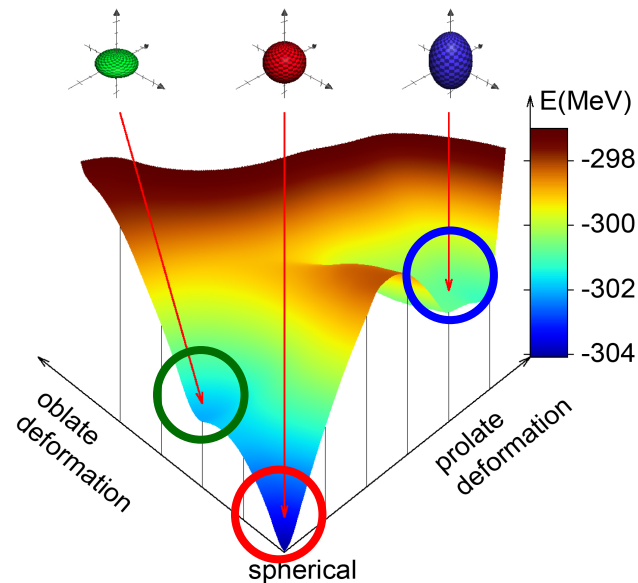
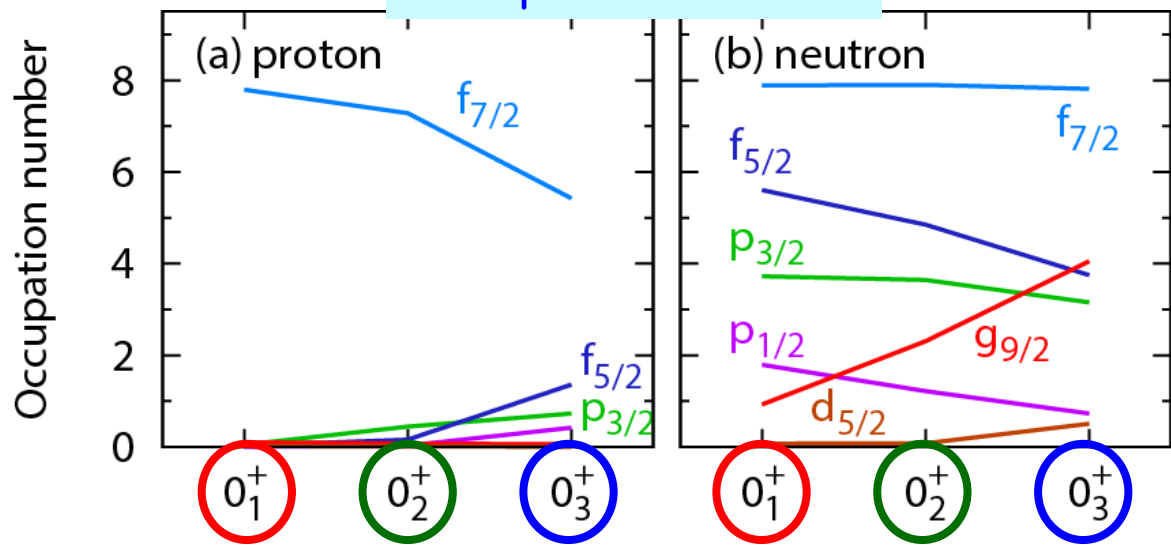
Energy levels and B(E2) values of Ni isotopes

Description by the same Hamiltonian

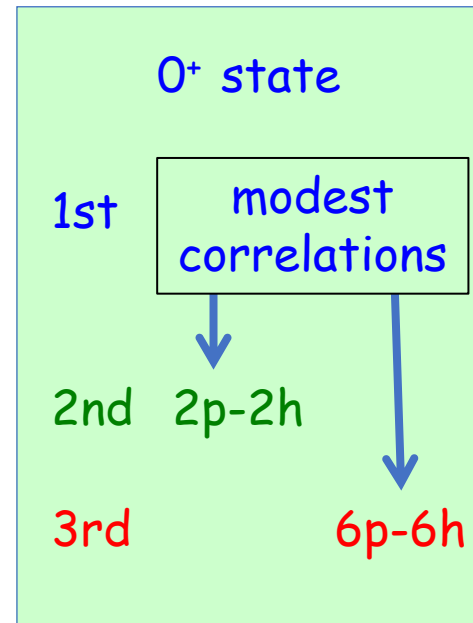
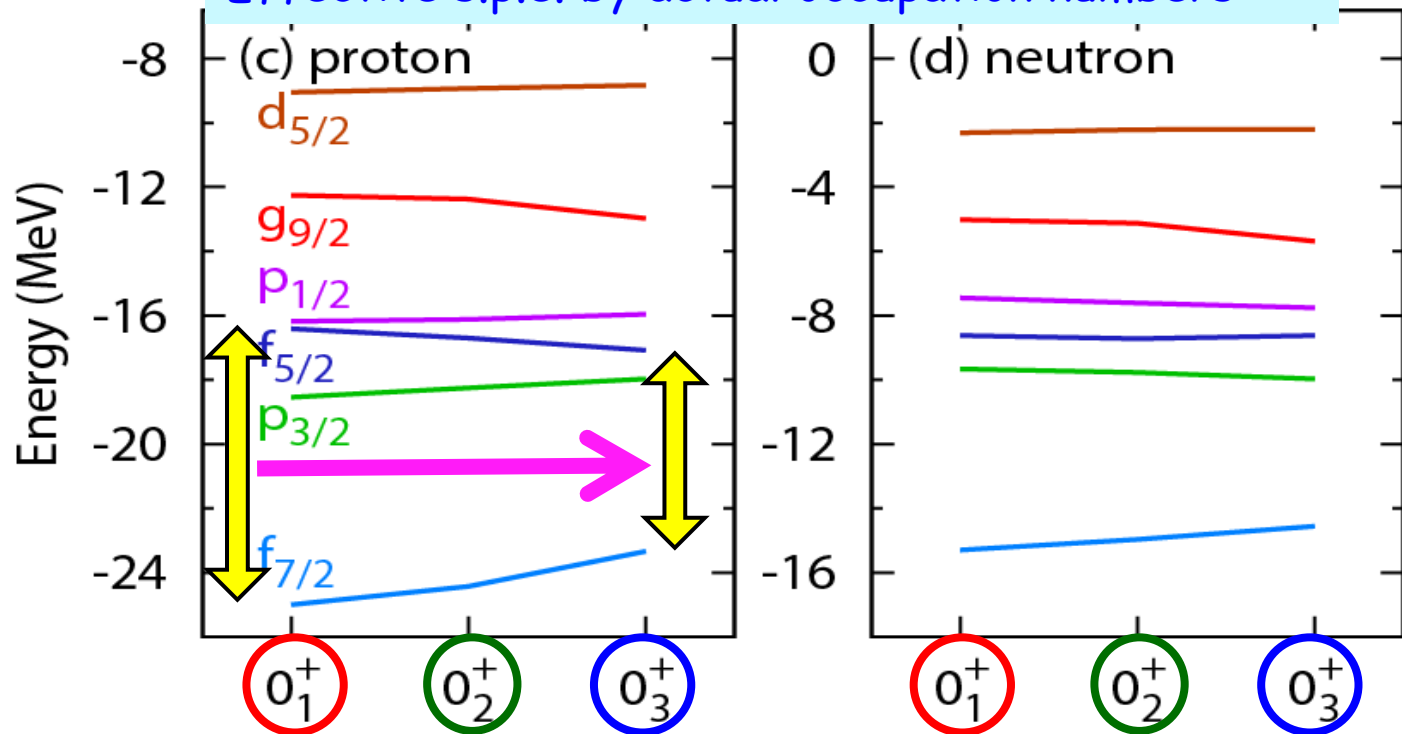
Shape coexistence in ^{68}Ni



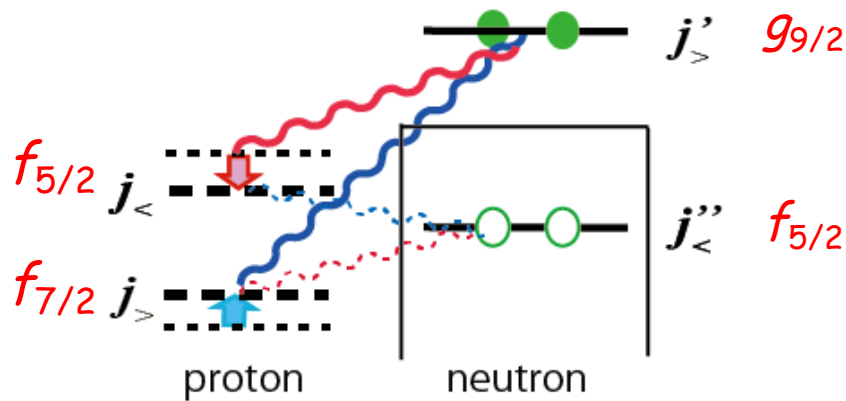
Occupation numbers



Effective s.p.e. by actual occupation numbers



Type II Shell Evolution in ^{68}Ni ($Z=28$, $N=40$)



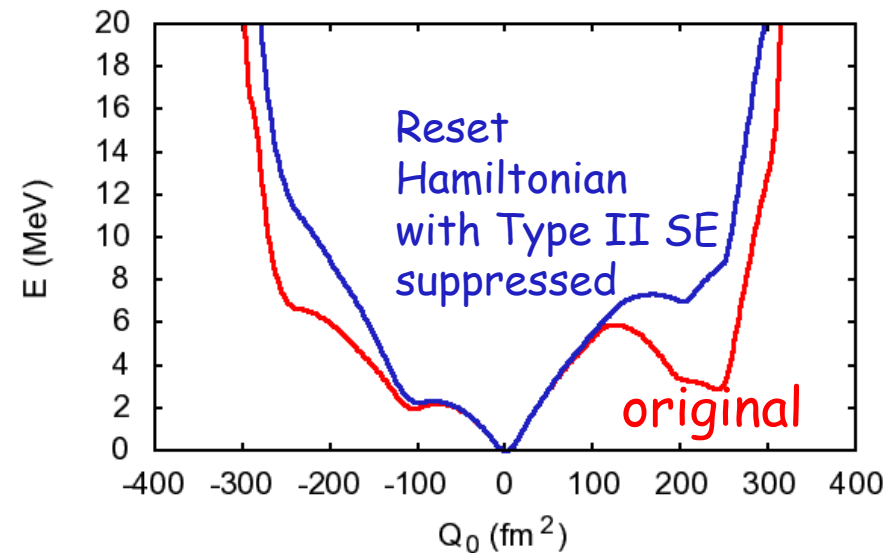
Spin-orbit splitting works against quadrupole deformation (*cf. Elliott's $SU(3)$*).

weakening of spin-orbit splitting

Type II shell evolution

stronger deformation of protons
→ more neutron p-h excitation

PES along axially symmetric shape



Type II shell evolution is suppressed by **resetting monopole interactions** as

$$\pi f_{7/2} - \nu g_{9/2} = \pi f_{5/2} - \nu g_{9/2}$$

$$\pi f_{7/2} - \nu f_{5/2} = \pi f_{5/2} - \nu f_{5/2}$$

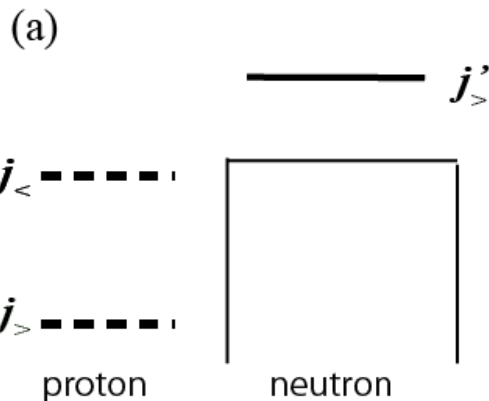
The local minima become much less pronounced.

Shape coexistence is enhanced by **type II shell evolution** because the same quadrupole interaction can work more efficiently.

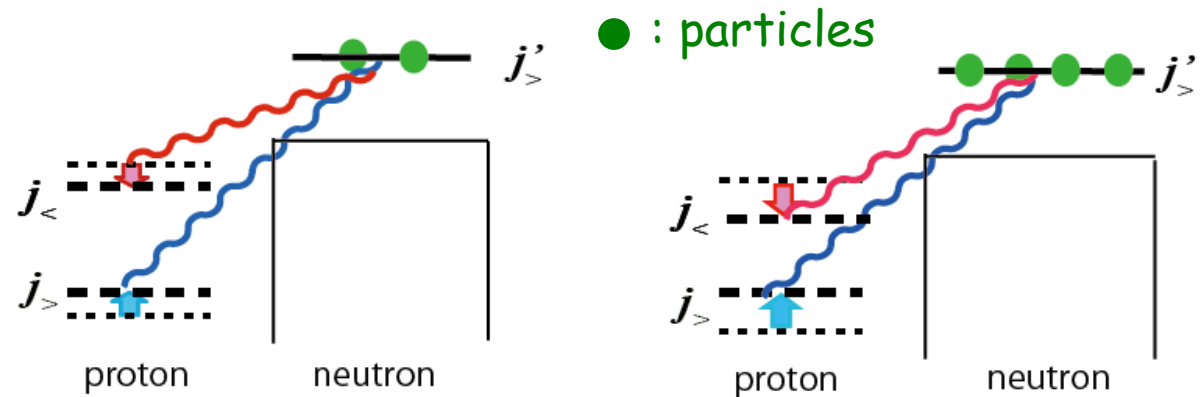
Underlying mechanism of the appearance of low-lying deformed states : Type II Shell Evolution

TO and Y. Tsunoda, J. Phys. G: Nucl. Part. Phys. 43 (2016) 024009

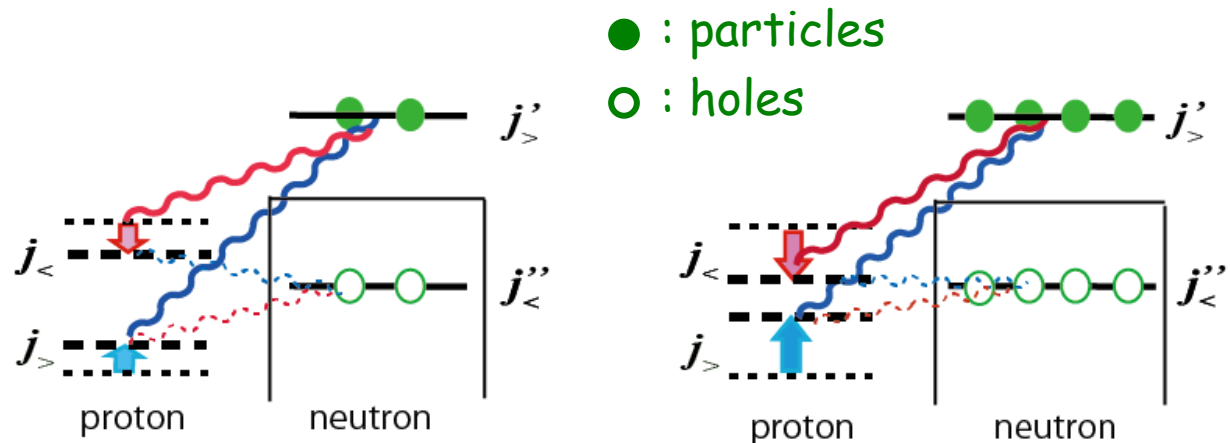
Monopole effects on the shell structure from the tensor interaction



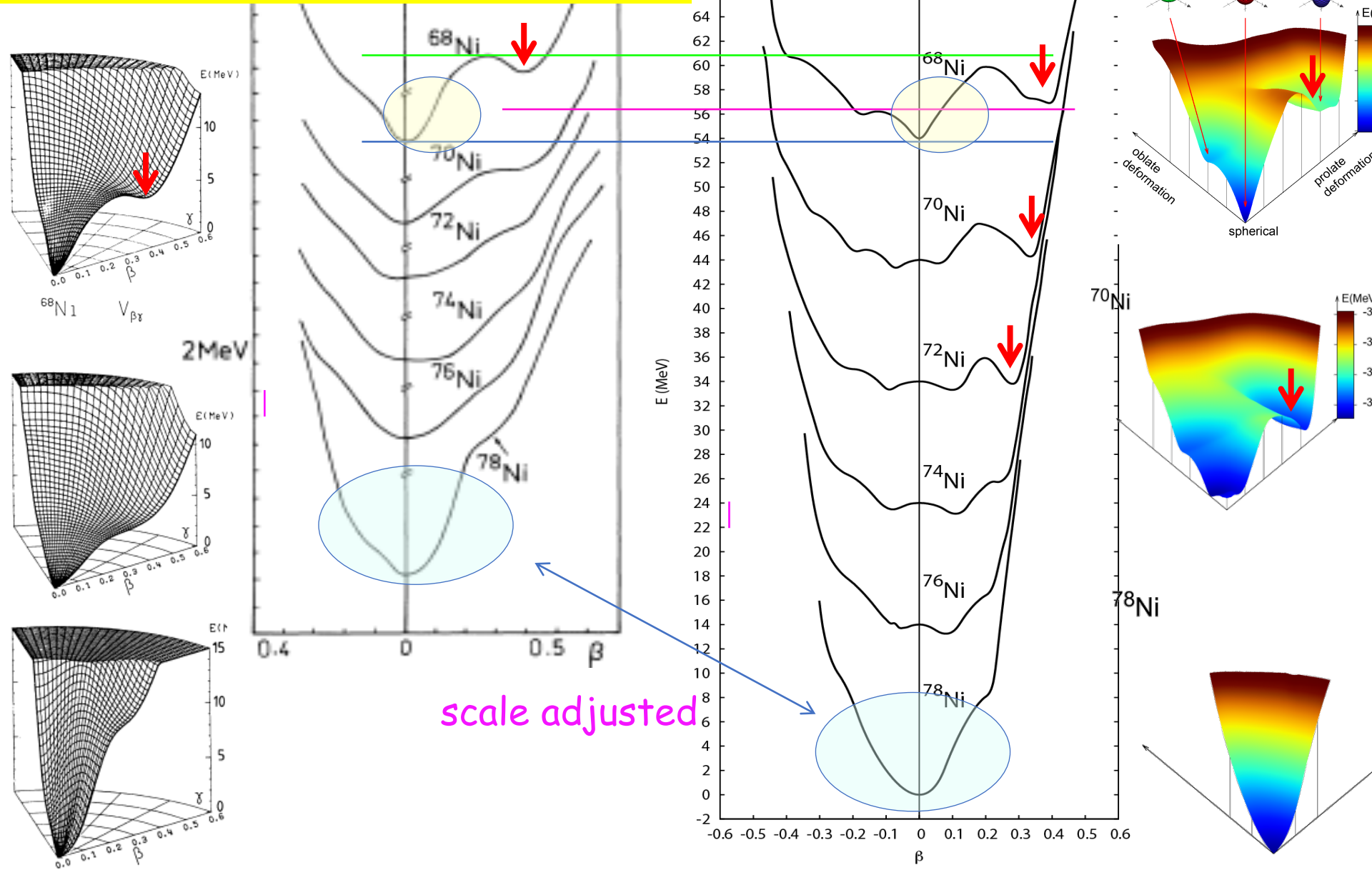
Type I Shell Evolution : different isotopes



Type II Shell Evolution : within the same nucleus

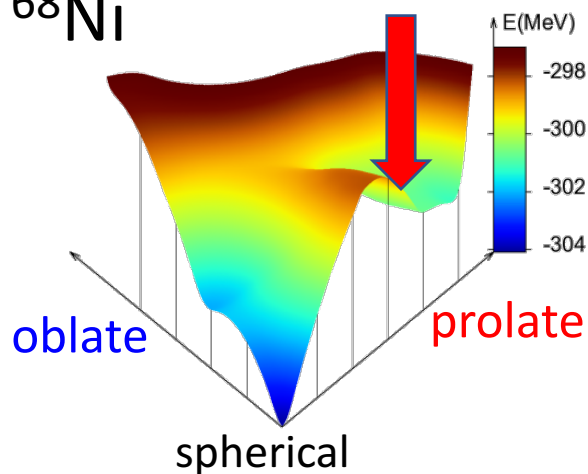


Bohr-model calc. by HFB with **Gogny** force,
Girod, Dessagne, Bernes, Langevin, Pougheon
and Roussel, PRC 37,2600 (1988)

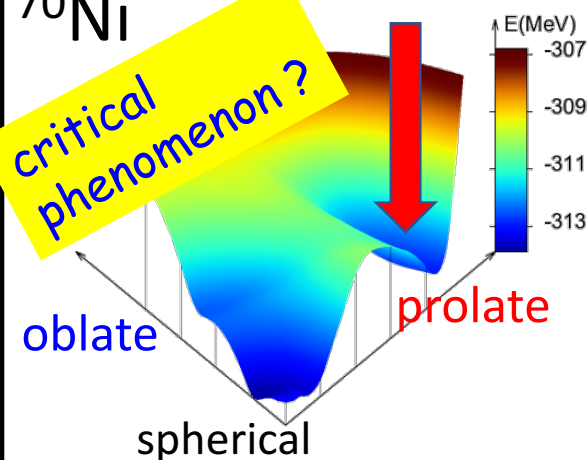


Shape or structure evolution of Ni isotopes

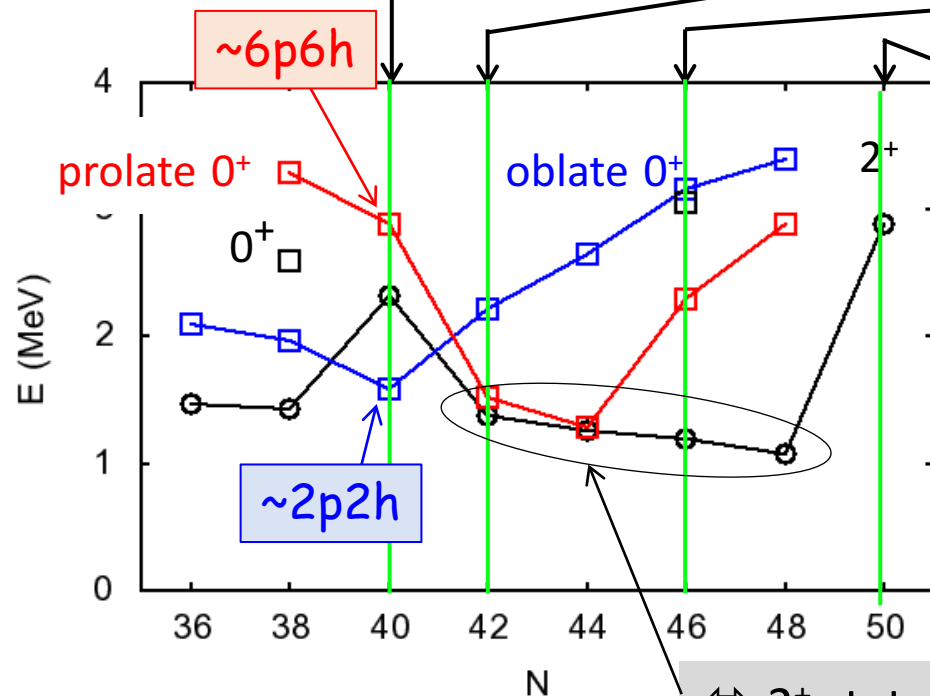
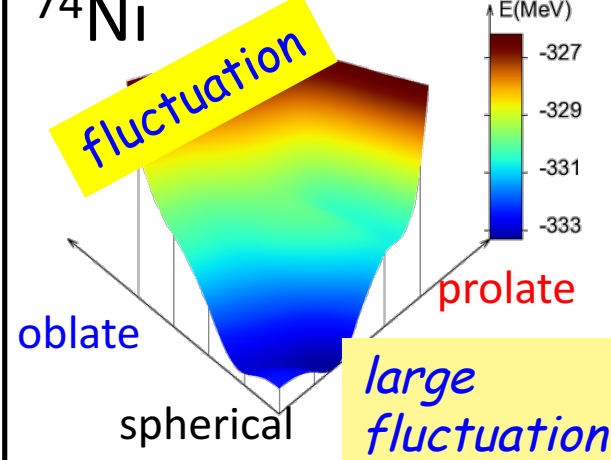
^{68}Ni



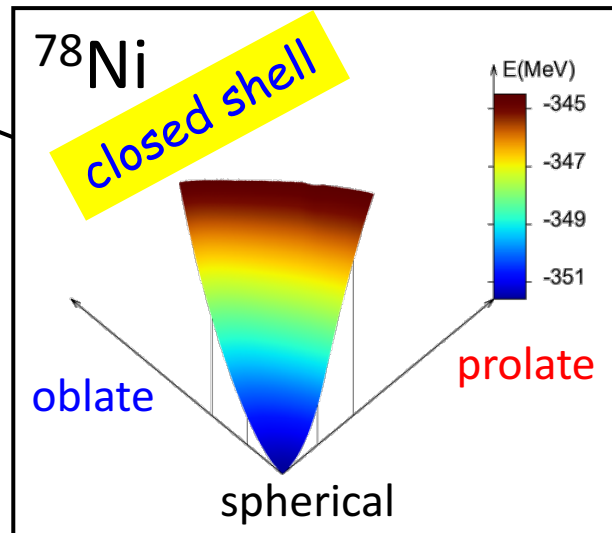
^{70}Ni



^{74}Ni



^{78}Ni



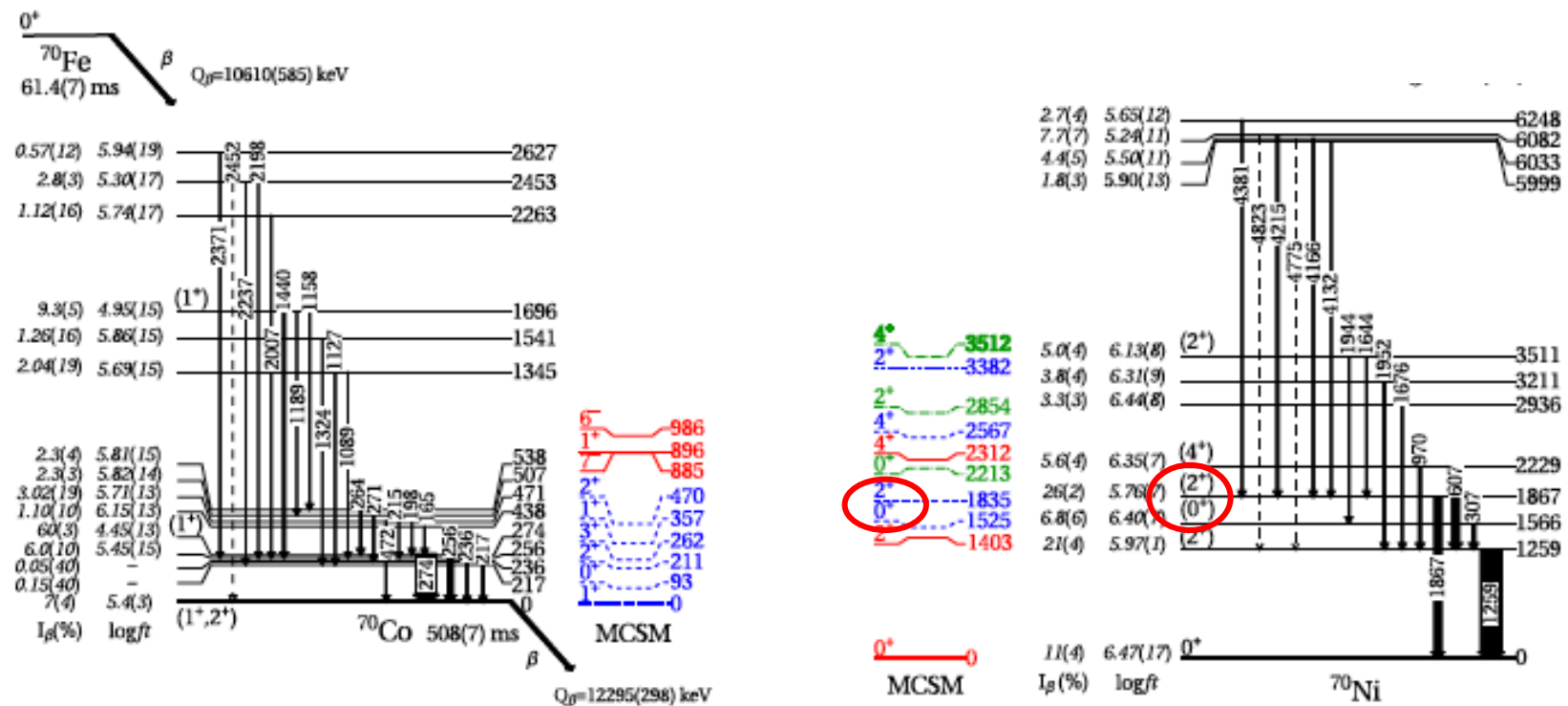
$\Leftrightarrow 2^+$ states in the $g_{9/2}$ seniority scheme

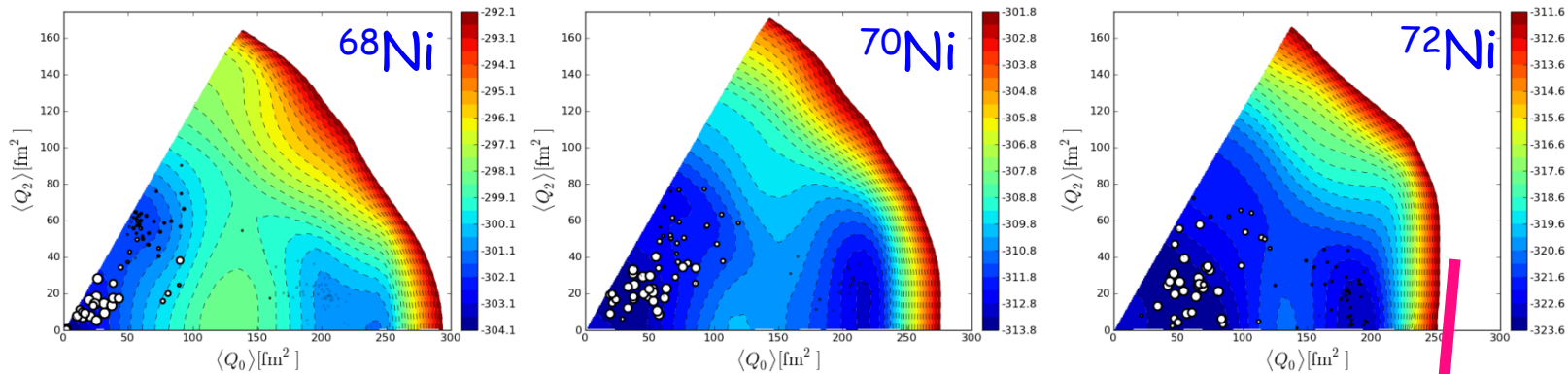
Type II shell evolution in $A = 70$ isobars from the $N \geq 40$ island of inversion

A.I. Morales^{a,b,*}, G. Benzoni^a, H. Watanabe^{c,d}, Y. Tsunoda^e, T. Otsuka^{f,g,h}, S. Nishimura^d, F. Browne^{i,d}, R. Daido^j, P. Doornenbal^d, Y. Fang^j, G. Lorusso^d, Z. Patel^{k,d}, S. Rice^{k,d}, L. Sinclair^{l,d}, P.-A. Söderström^d, T. Sumikama^m, J. Wu^d, Z.Y. Xu^{f,d}, A. Yagi^j, R. Yokoyama^f, H. Baba^d, R. Avigo^{a,b}, F.L. Bello Garroteⁿ, N. Blasi^a, A. Bracco^{a,b}, F. Camera^{a,b}, S. Ceruti^{a,b}, F.C.L. Crespi^{a,b}, G. de Angelis^o, M.-C. Delattre^p, Zs. Dombradi^q, A. Gottardo^o, T. Isobe^d, I. Kojouharov^r, N. Kurz^r, I. Kuti^q, K. Matsui^f, B. Melon^s, D. Mengoni^{t,u}, T. Miyazaki^f, V. Modamio-Hoybjør^o, S. Momiyama^f, D.R. Napoli^o, M. Niikura^f, R. Orlandi^{h,v}, H. Sakurai^{d,f}, E. Sahinⁿ, D. Sohler^q, H. Schaffner^r, R. Taniuchi^f, J. Taprogge^{w,x}, Zs. Vajta^q, J.J. Valiente-Dobón^o, O. Wieland^a, M. Yalcinkaya^y

^a Istituto Nazionale di Fisica Nucleare, Sezione di Milano, Via Celoria 16, 20133 Milano, Italy

^b Dipartimento di Fisica, Università degli Studi di Milano, Via Celoria 16, 20133 Milano, Italy





Energy of prolate state comes down.
Barrier becomes low.

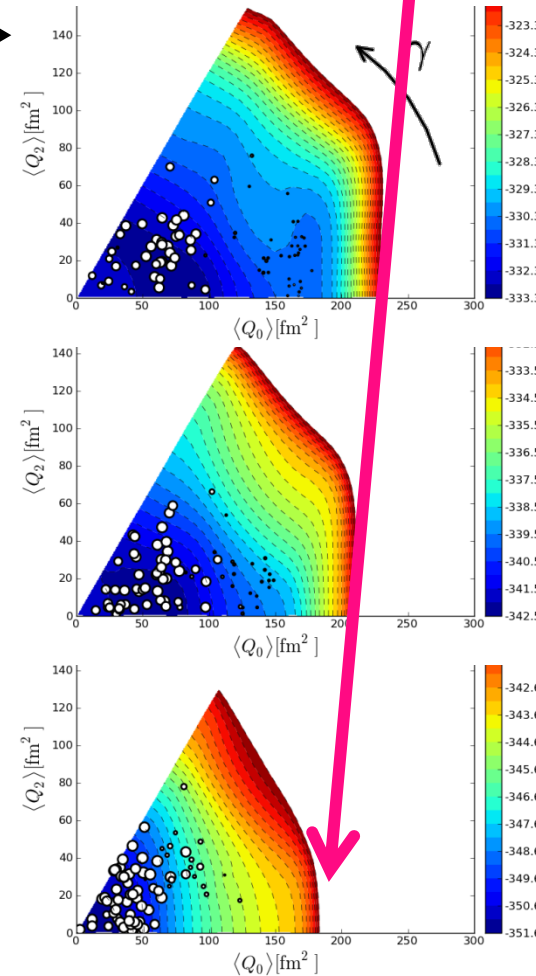
T-plot
for 0^+_1 states of $^{68-78}\text{Ni}$

The ground state is always like
seniority-zero (BCS-type)
spherical state.

^{74}Ni

^{76}Ni

^{78}Ni



γ -soft deformation
(strong fluctuation in the γ direction)

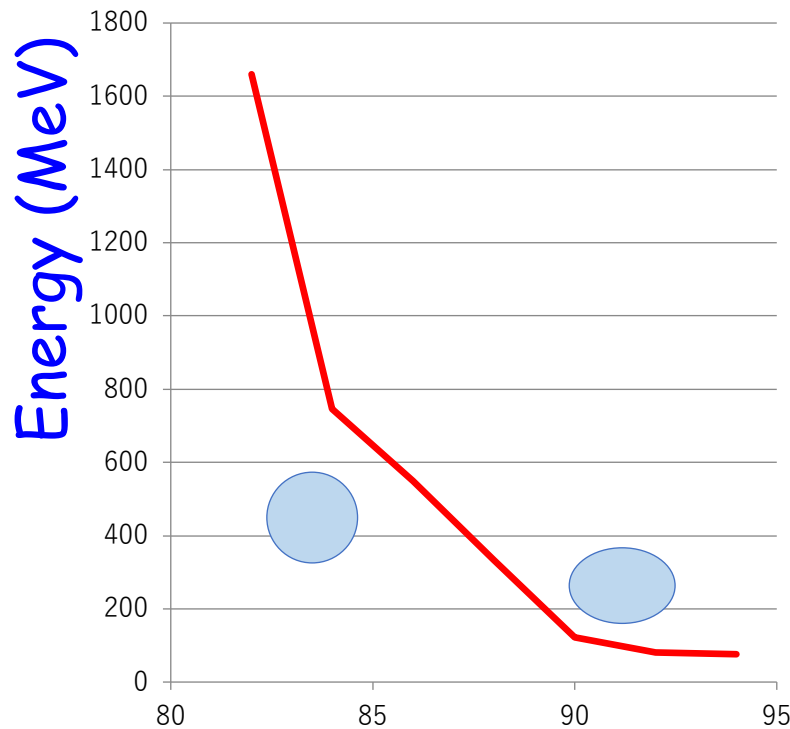
Shape transition and quantum phase transition

Shape change as a function of N (or Z)

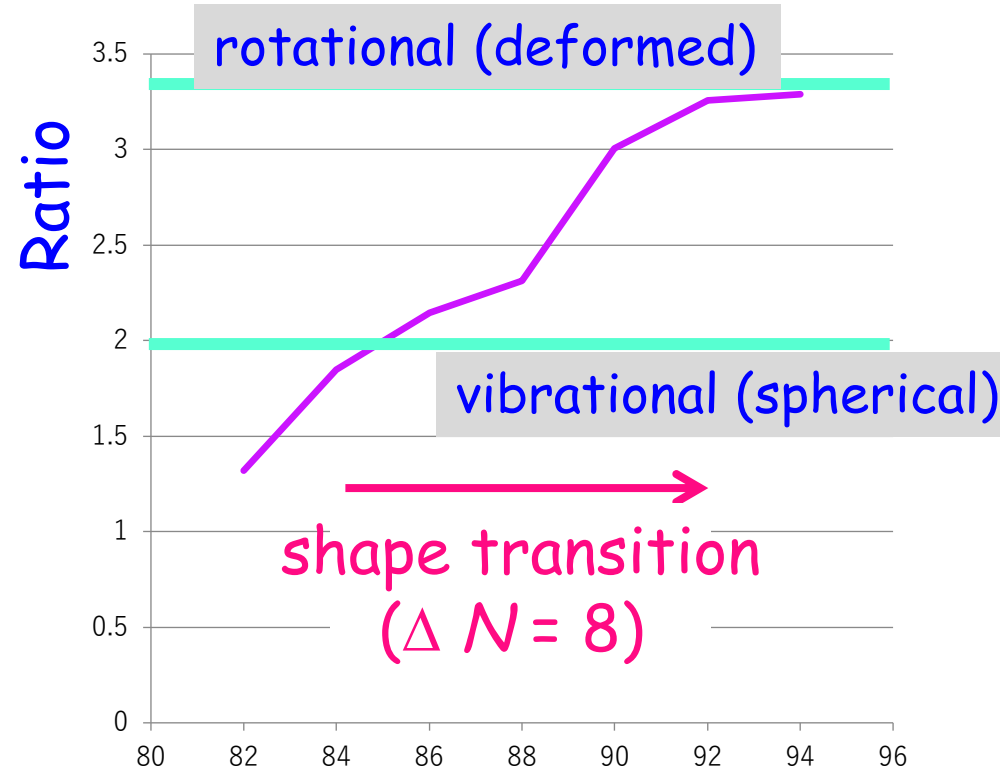
2^+ and 4^+ level properties of Sm isotopes

$Ex(2^+)$:
excitation energy of first 2^+ state

$$R_{4/2} = Ex(4^+) / Ex(2^+)$$



Neutron number, N



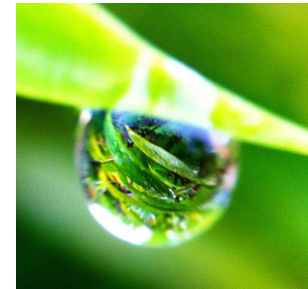
Neutron number, N

Can this be a kind of *Phase Transition*?

Can the shape transition in nuclei be a "Phase Transition" ?



ice



water

Phase Transition :

A **macroscopic** system can change qualitatively from a stable state (e.g. ice for H_2O) to another stable state (e.g., water for H_2O) as a function of a certain parameter (e.g., temperature).

The phase transition implies this kind of phenomena of macroscopic systems consisting of **almost infinite number of molecules**.

Can the shape transition be a "Quantum Phase Transition" ?

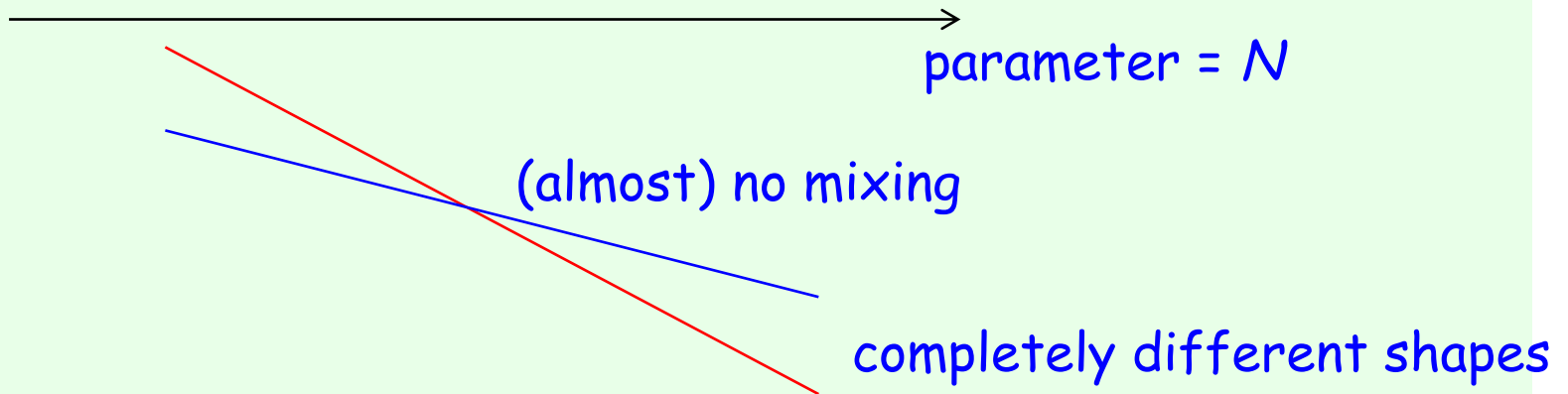
The shape transition occurs rather gradually.

Quantum Phase Transition :

an abrupt change in the ground state of a many-body system by varying a physical parameter at zero temperature.

(cf., Wikipedia)

possible scenario



Sizable mixing occurs usually in finite quantum systems.

Quest for Quantum Phase Transition: Shapes of Zr isotopes by Monte Carlo Shell Model

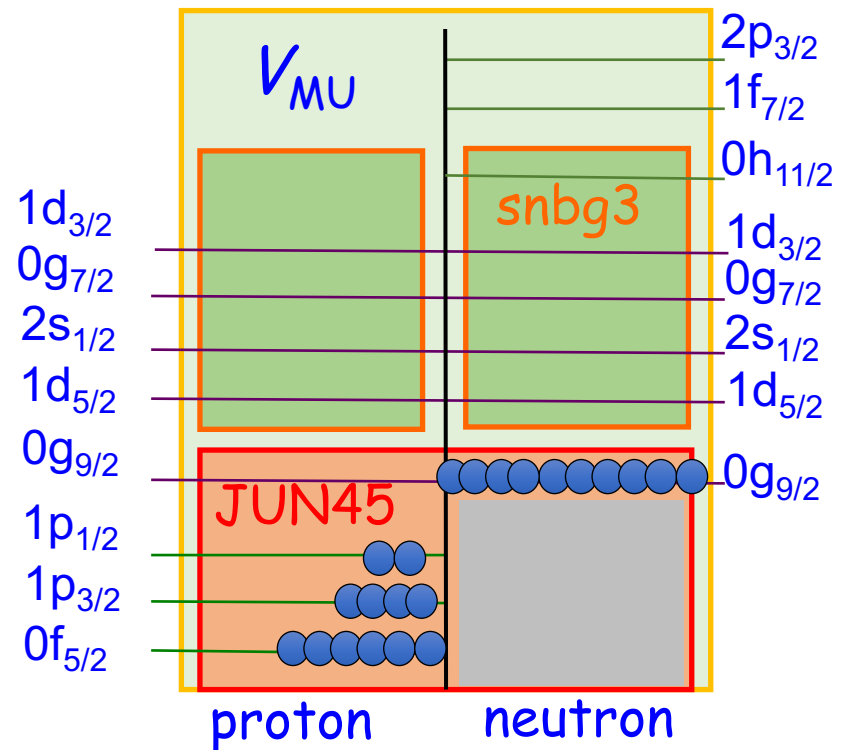
- Effective interaction:
 $\text{JUN45} + \text{snbg3} + V_{\text{MU}}$

known effective interactions

+ minor fit for a part of
T=1 TBME's

Nucleons are excited fully
within this model space
(no truncation)

We performed **Monte Carlo Shell Model (MCSM)** calculations, where the largest case corresponds to the diagonalization of 3.7×10^{23} **dimension** matrix.



Togashi, Tsunoda, TO *et al.* PRL
117, 172502 (2016)

From earlier shell-model works ...

PHYSICAL REVIEW C

VOLUME 20, NUMBER 2

AUGUST 1979

Unified shell-model description of nuclear deformation

P. Federman

Instituto de Física, Universidad Nacional Autónoma de México, Apartado Postal 20-364, México 20, D. F.

S. Pittel

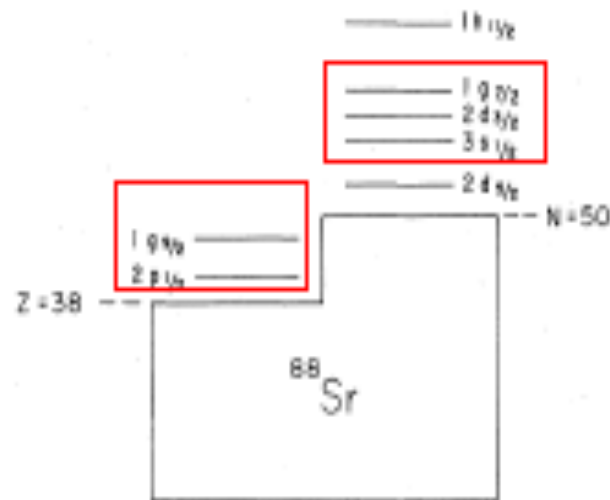


FIG. 3. Single-particle levels appropriate to a description of nuclei in the Zr-Mo region. An ^{88}Sr core is assumed.

PHYSICAL REVIEW C 79, 064310 (2009)

Shell model description of zirconium isotopes

K. Sieja,^{1,2} F. Nowacki,³ K. Langanke,^{2,4} and G. Martínez-Pinedo¹

In this paper, we perform for the first time a SM study of Zr isotopes in an extended model space ($1f_{5/2}$, $2p_{1/2}$, $2p_{3/2}$, $1g_{9/2}$) for protons and ($2d_{5/2}$, $3s_{1/2}$, $2d_{3/2}$, $1g_{7/2}$, $1h_{11/2}$) for neutrons, dubbed hereafter $\pi(r3 - g)$, $\nu(r4 - h)$.

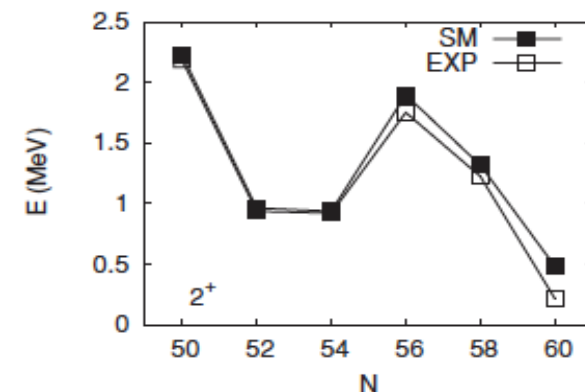
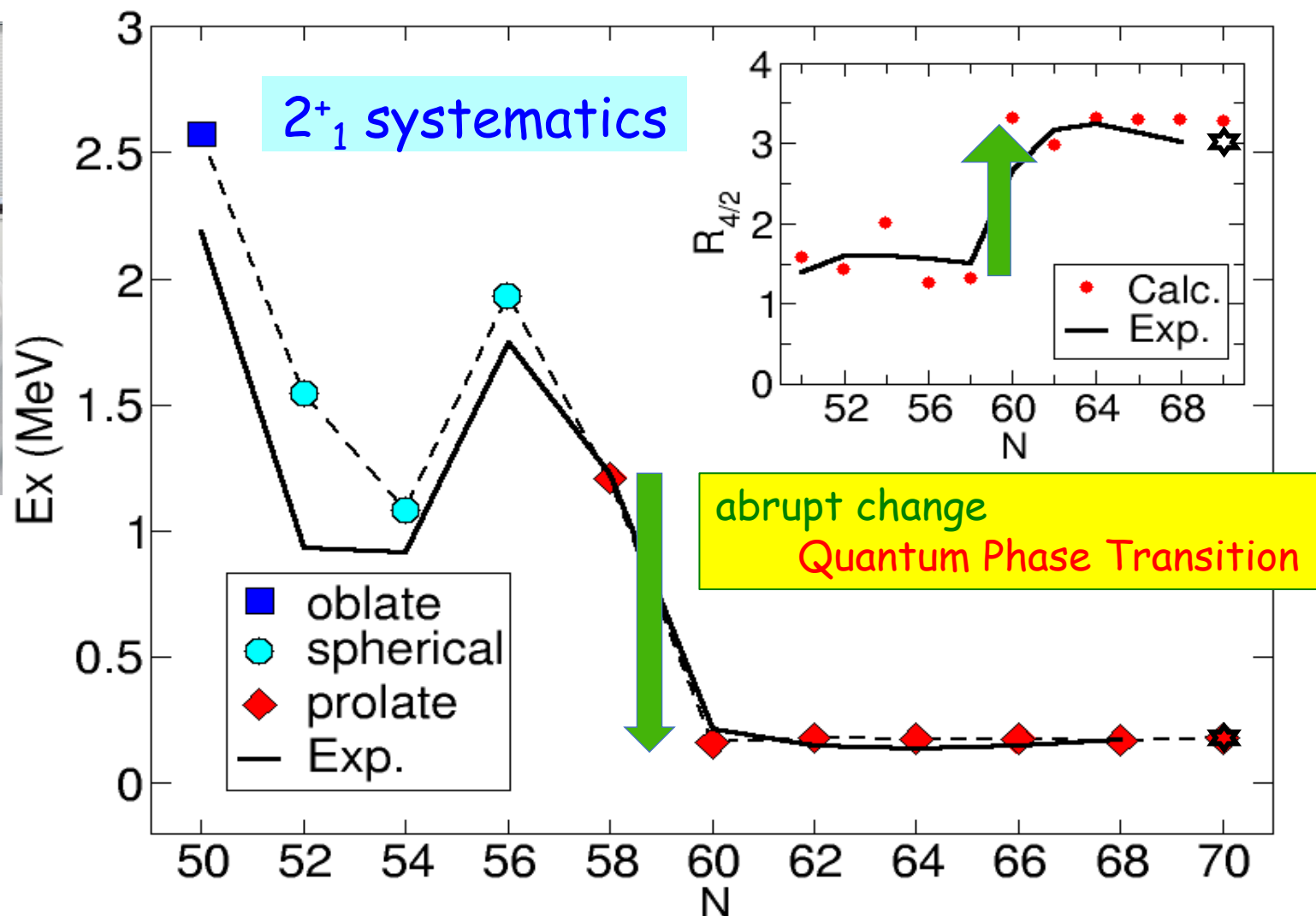


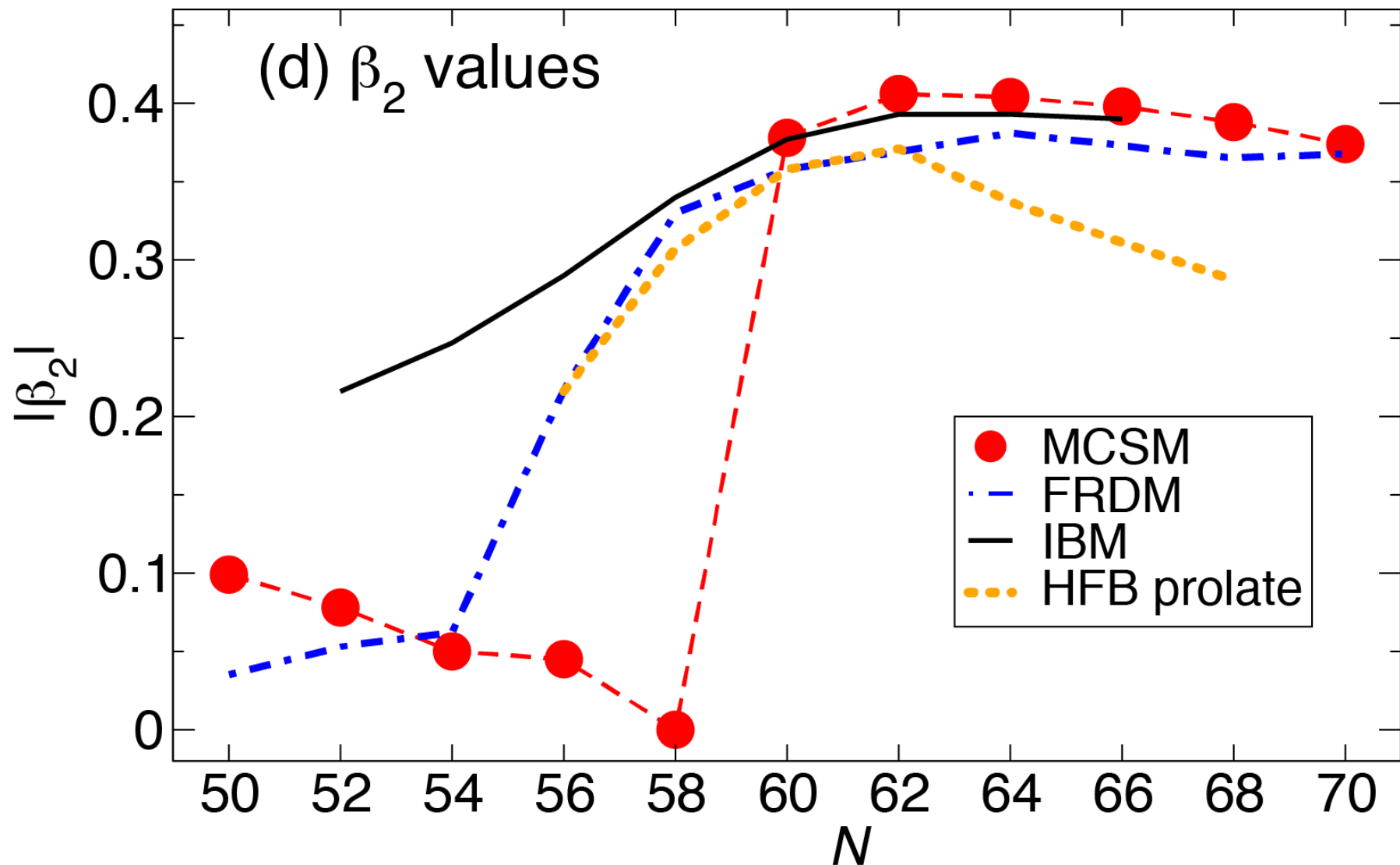
FIG. 12. Systematics of the experimental and theoretical first excited 2^+ states along the zirconium chain.



Quantum Phase Transition in the Shape of Zr isotopes

Tomoaki Togashi,¹ Yusuke Tsunoda,¹ Takaharu Otsuka,^{1,2,3,4} and Noritaka Shimizu¹

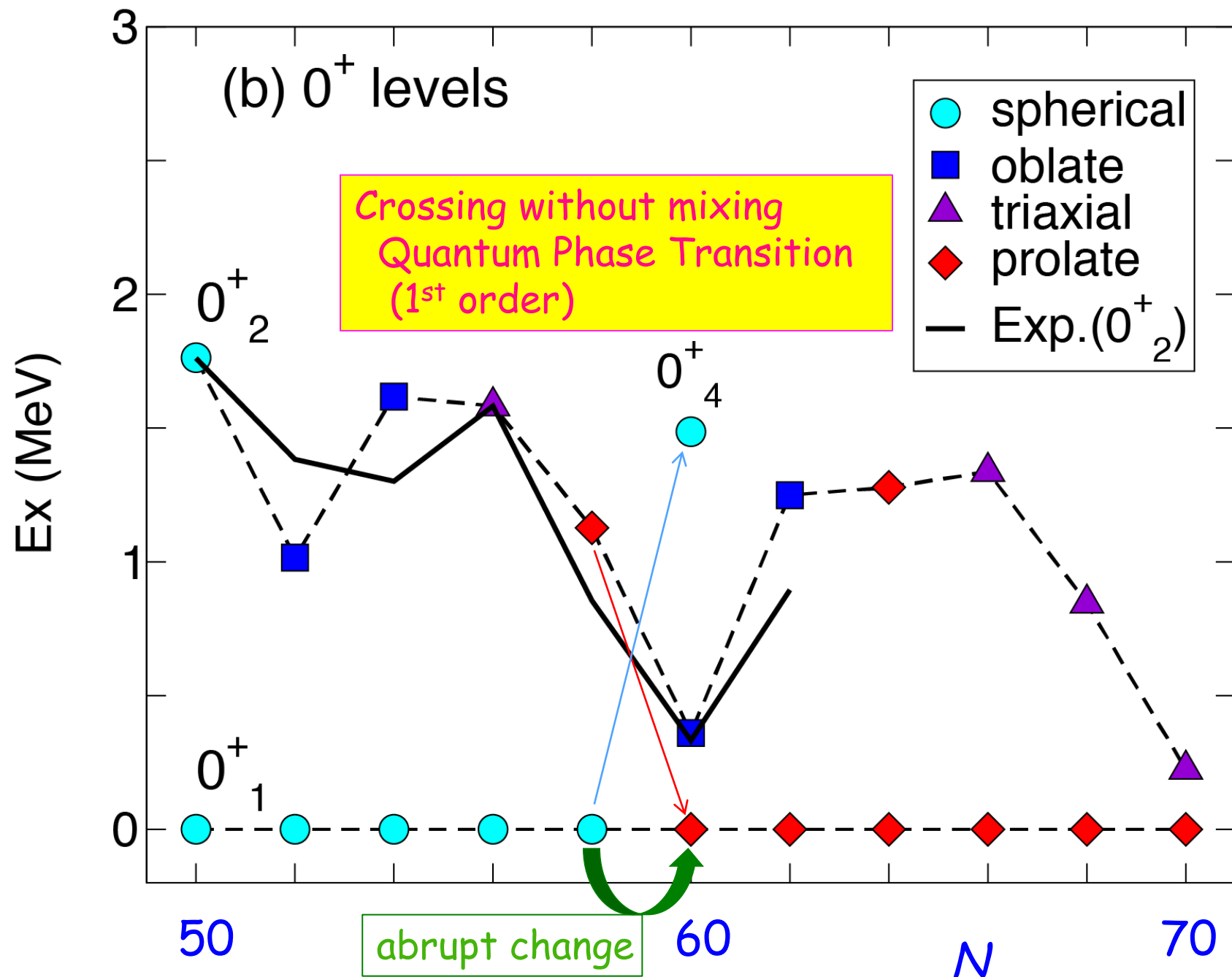




FRDM: S. Moeller et al. At. Data Nucl. Data Tables 59, 185 (1995).

IBM: M. Boyukata et al. J. Phys. G 37, 105102 (2010).

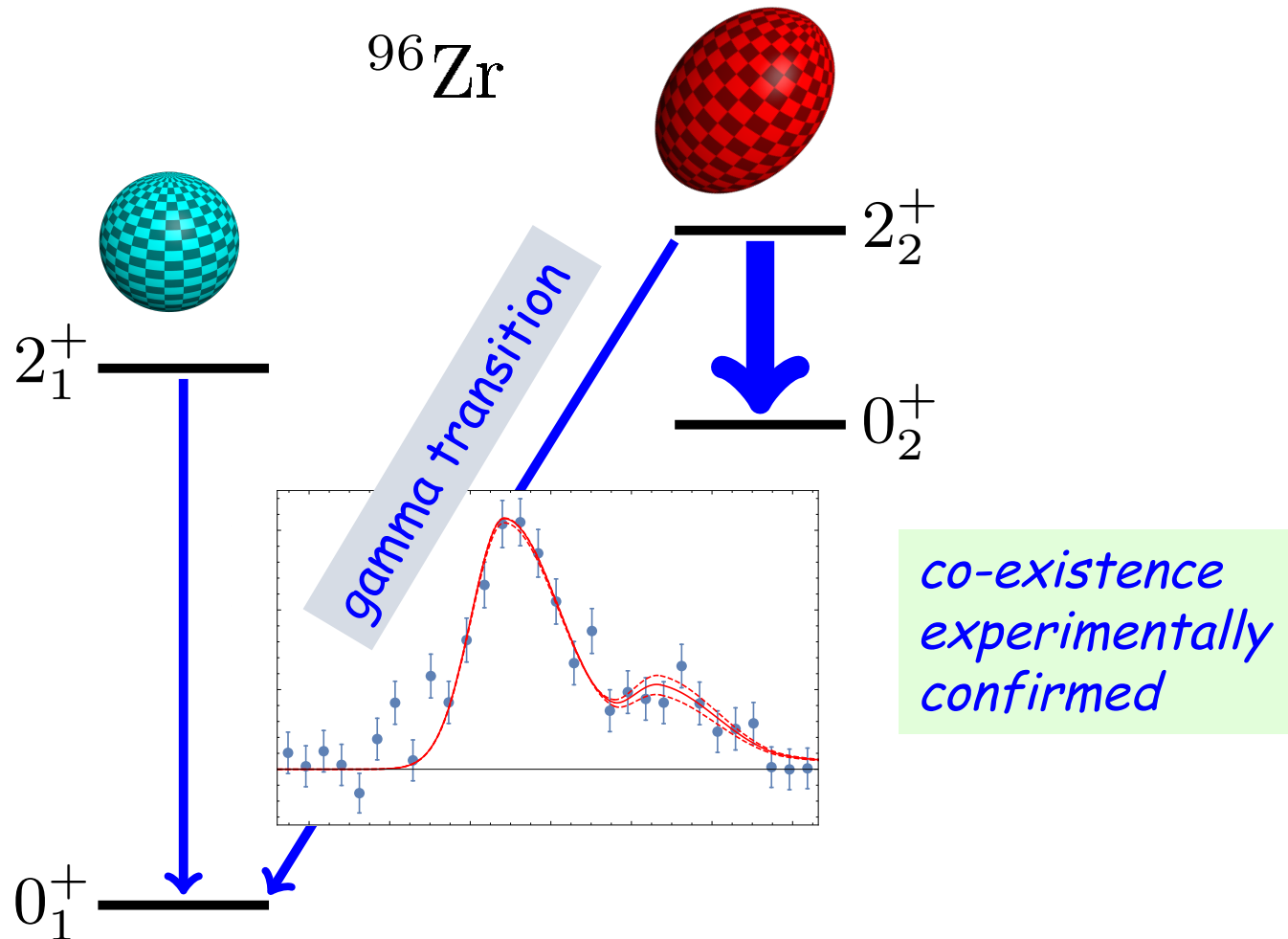
HFB: R. Rodriguez-Guzman et al. Phys. Lett. B 691, 202 (2010).⁹⁸



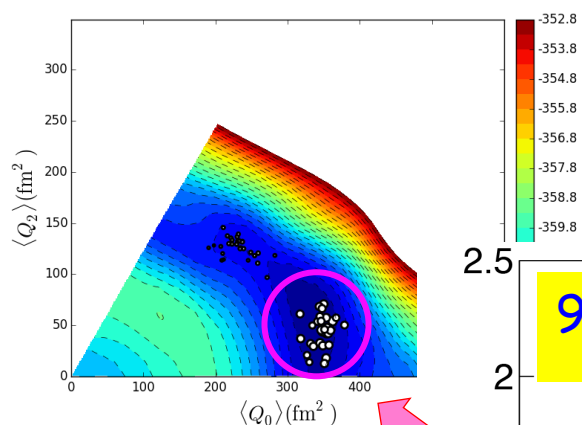
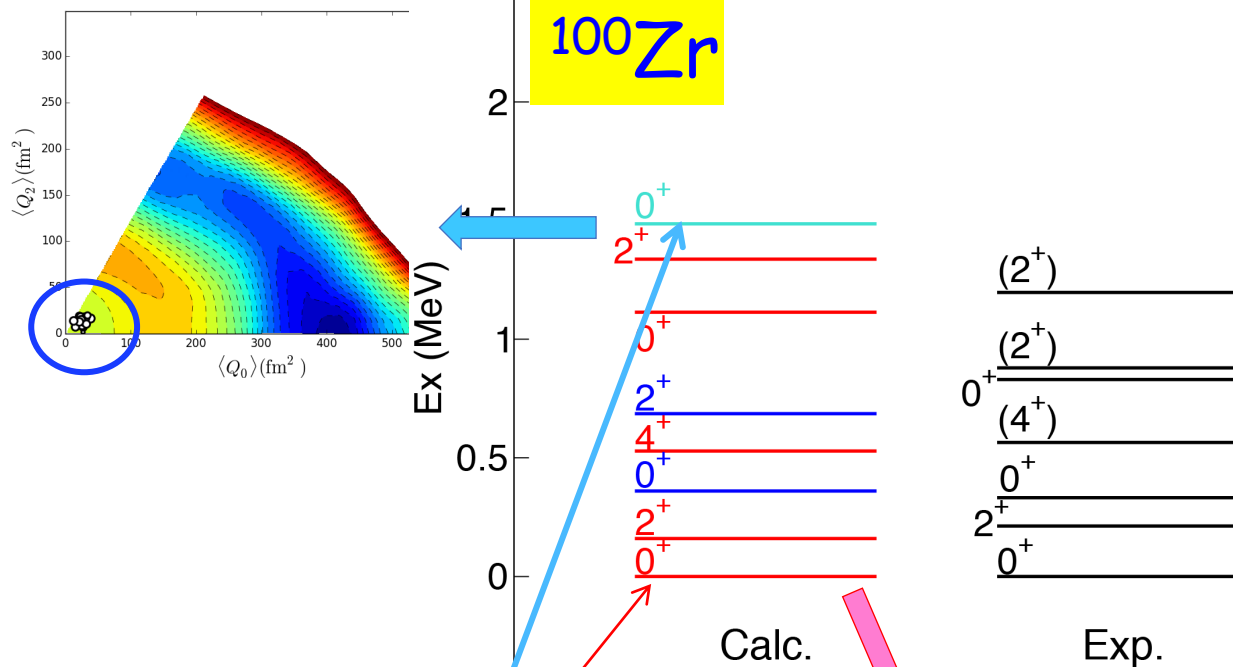


First Measurement of Collectivity of Coexisting Shapes Based on Type II Shell Evolution: The Case of ^{96}Zr

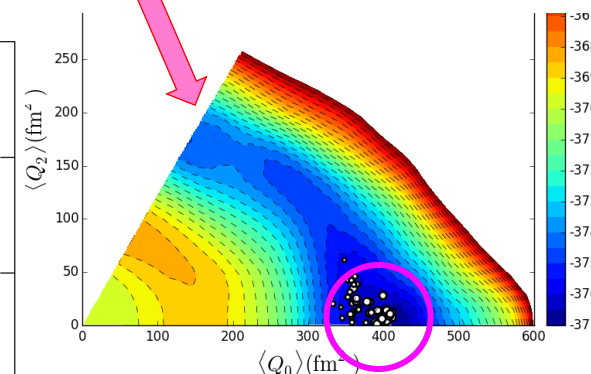
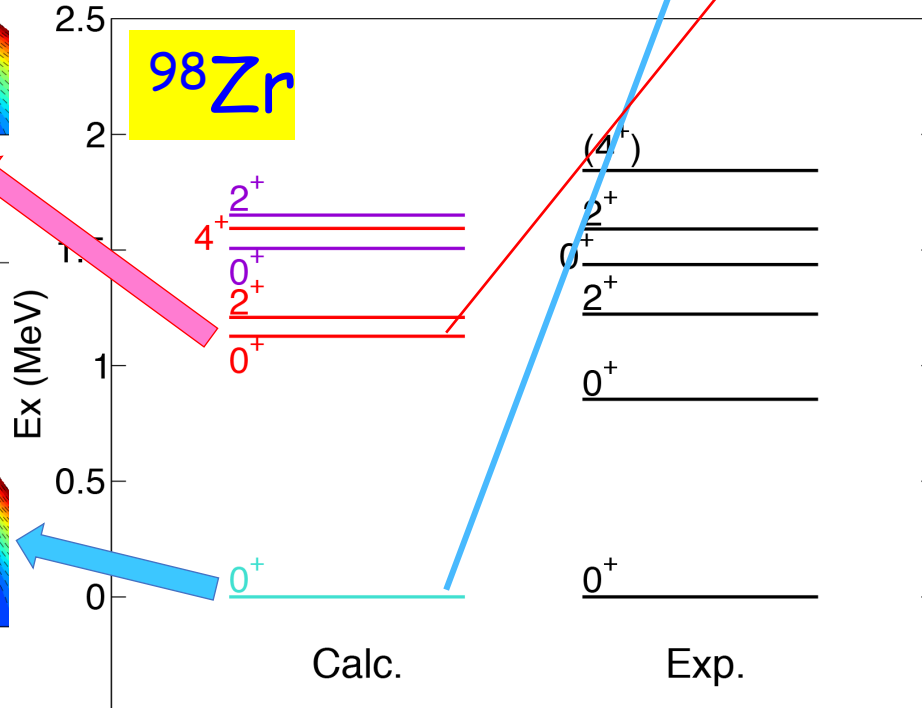
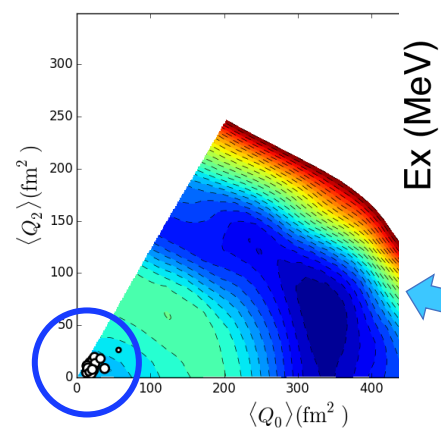
C. Kremer,¹ S. Aslanidou,¹ S. Bassauer,¹ M. Hilcker,¹ A. Krugmann,¹ P. von Neumann-Cosel,¹
T. Otsuka,^{2,3,4,5} N. Pietralla,¹ V. Yu. Ponomarev,¹ N. Shimizu,³ M. Singer,¹ G. Steinhilber,¹
T. Togashi,³ Y. Tsunoda,³ V. Werner,¹ and M. Zweidinger¹

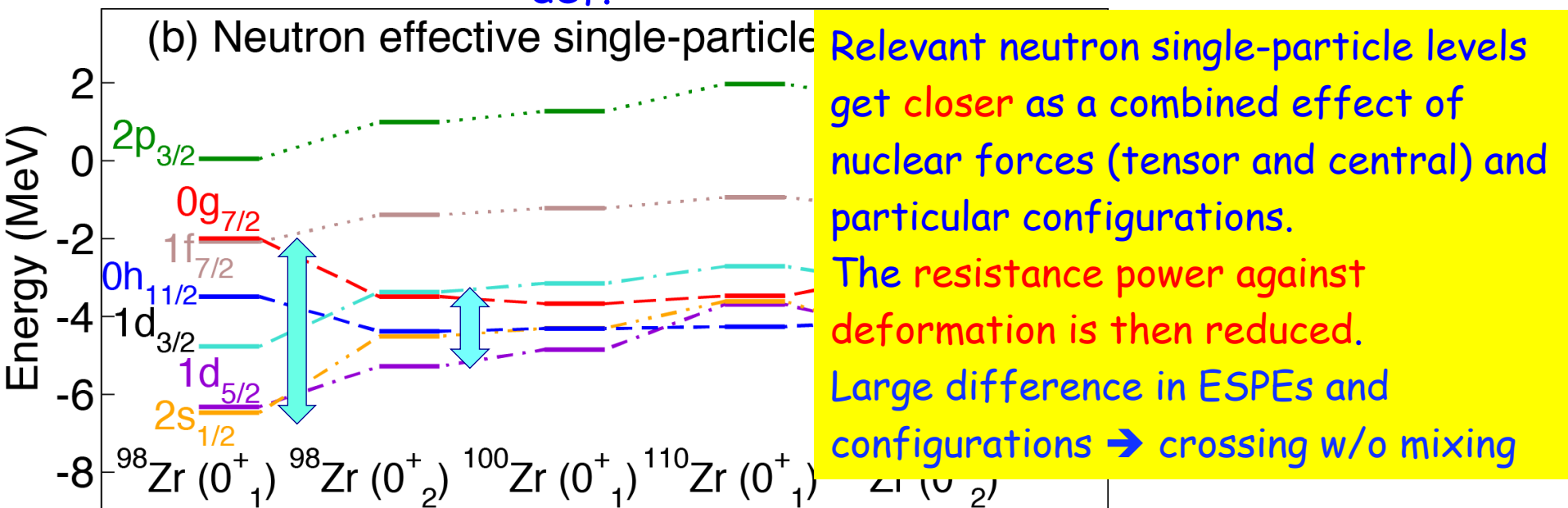
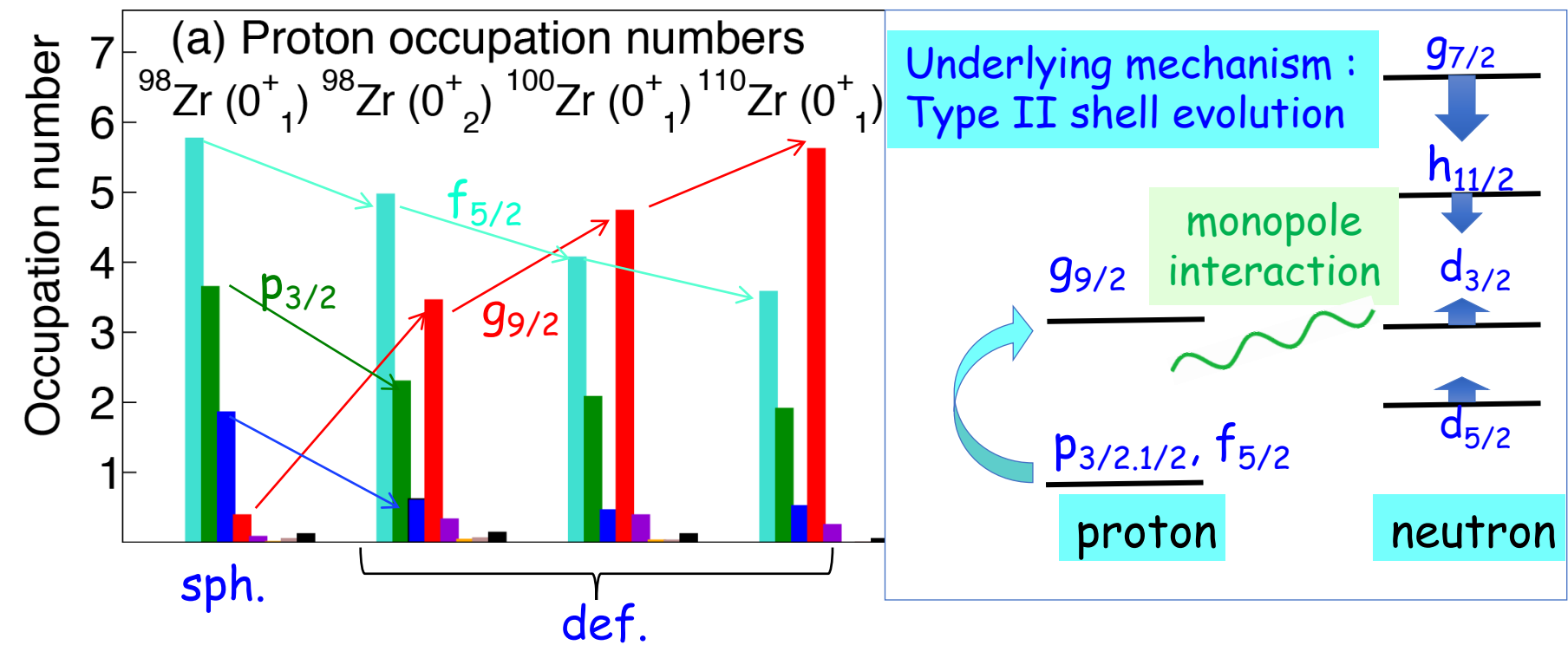


Quantum Phase
Transition
(1st order)
due to crossing
without mixing



^{98}Zr

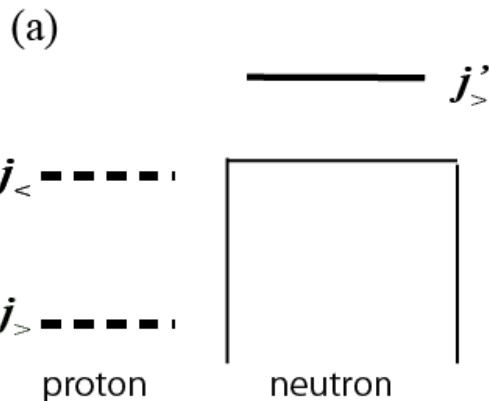




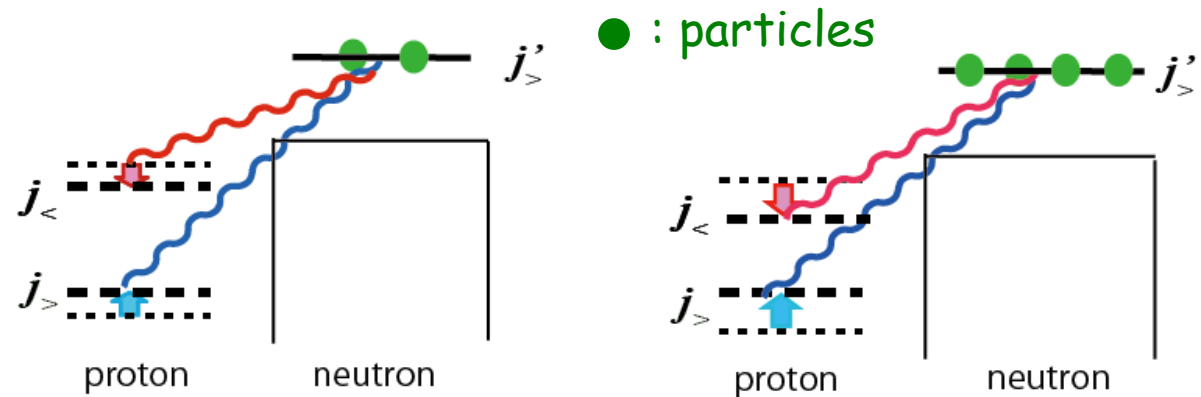
Underlying mechanism of the appearance of low-lying deformed states : Type II Shell Evolution

TO and Y. Tsunoda, J. Phys. G: Nucl. Part. Phys. 43 (2016) 024009

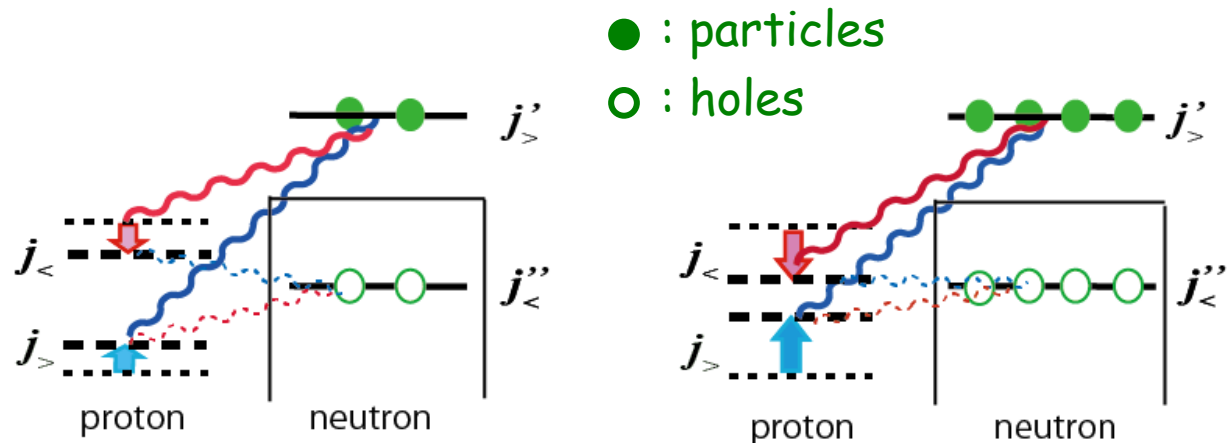
Monopole effects on the shell structure from the tensor interaction



Type I Shell Evolution : different isotopes



Type II Shell Evolution : within the same nucleus



Type II shell evolution is a simplest and visible case of

Quantum Self Organization

$$\text{deformation} = \frac{\text{quadrupole force}}{\text{resistance power}}$$

resistance power ← pairing force

↑
single-particle energies

Atomic nuclei can “organize” their single-particle energies by taking particular configurations of protons and neutrons optimized for each eigenstate, thanks to orbit-dependences of monopole components of nuclear forces (e.g., tensor force).
→ an enhancement of Jahn-Teller effect.

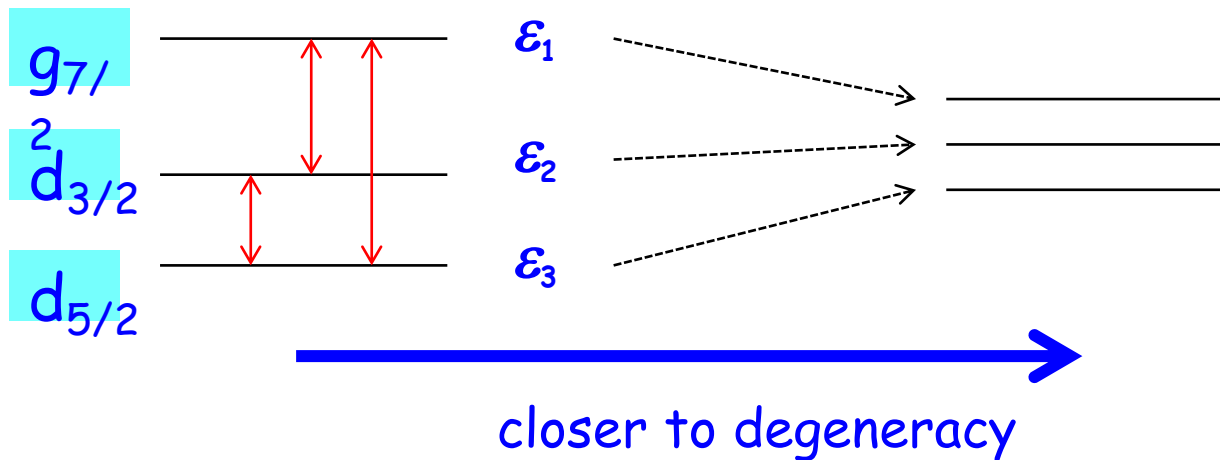
Reminder : Jahn -Teller effect for nuclear deformation

(Self-consistent) quadrupole deformed field $\propto Y_{2,0}(\theta, \phi)$
mixes the orbits below

$$\Psi(J_z=1/2) = c_1 |g_{7/2}; j_z=1/2\rangle + c_2 |d_{3/2}; j_z=1/2\rangle + c_3 |d_{5/2}; j_z=1/2\rangle$$

stronger mixing = larger quadrupole deformation

Mixing depends not only on the strength of the $Y_{2,0}(\theta, \phi)$ field, but also the spherical single-particle energies $\epsilon_1, \epsilon_2, \epsilon_3$, etc.

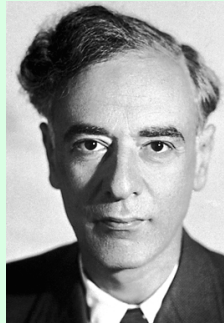


larger deformation for the same deformed field

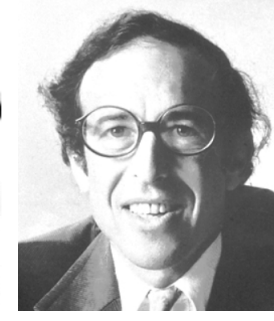
Let's shed light on an old problem

*Atomic nucleus is a quantum
Fermi liquid :*

*The nucleus is composed
of almost **free nucleons**
interacting weakly via
residual forces
in a (solid) (mean) potential
like a **solid vase**.*

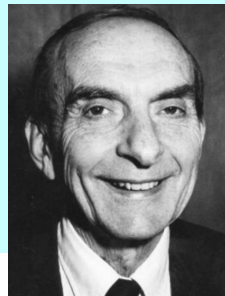


Surface deformation produced
by additional **deformed mean
field** → *Nilsson model*
(deformed "vase",
monopole effect missing)



SVEN GÖSTA NILSSON

“how single-particle states can coexist
with collective modes” *conceived* also by
Gerry Brown as an open unresolved
problem (T. Schaefer,
Fermi Liquid theory: A brief
survey in memory of Gerald
E. Brown, NPA 2014)



Renewed picture:

*Single-particle levels can be
re-organized to enhance
collective modes.*

*SP states and coll. mode are
not enemies but friends !*

Nilsson model Hamiltonian

"Nuclear structure II" by Bohr and Mottelson

deformed nuclei, is obtained by a simple modification of the harmonic oscillator (Nilsson, 1955; Gustafson *et al.*, 1967),

$$H = \frac{\mathbf{p}^2}{2M} + \frac{1}{2}M(\omega_3^2 x_3^2 + \omega_{\perp}^2 (x_1^2 + x_2^2)) + v_{ll}\hbar\omega_0(l^2 - \langle l^2 \rangle_N) + v_{ls}\hbar\omega_0(\mathbf{l} \cdot \mathbf{s})$$

quadrupole deformed field

spherical field (5-10)

$$\langle l^2 \rangle_N = \frac{1}{2}N(N+3)$$

Figure	Region	$-v_{ls}$	$-v_{ll}$
5-1	N and $Z < 20$	0.16	0
5-2	$50 < Z < 82$	0.127	0.0382
5-3	$82 < N < 126$	0.127	0.0268
5-4	$82 < Z < 126$	0.115	0.0375
5-5	$126 < N$	0.127	0.0206

Table 5-1 Parameters used in the single-particle potentials of Figs. 5-1 to 5-5.

Spin-orbit force

$$\left. \begin{array}{ll} A=68 & 1.28 \\ A=100 & 1.12 \\ A=186 & 0.91 \end{array} \right\} (\mathbf{l} \cdot \mathbf{s})$$

$$\text{deformation} = \frac{\text{quadrupole force}}{\text{resistance power}}$$

resistance power \leftarrow pairing force

 single-particle energies

Analogy to electric current,

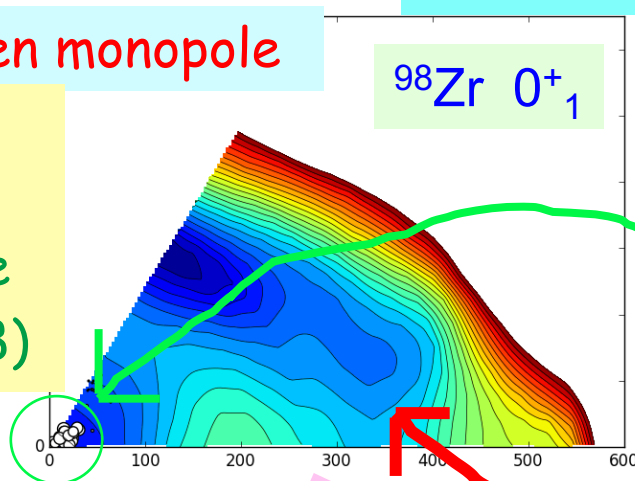
$$\text{current} = \frac{\text{voltage}}{\text{resistance}}$$

Anatomy of this effect : ^{98}Zr spherical 0^+_1 and deformed 0^+_2

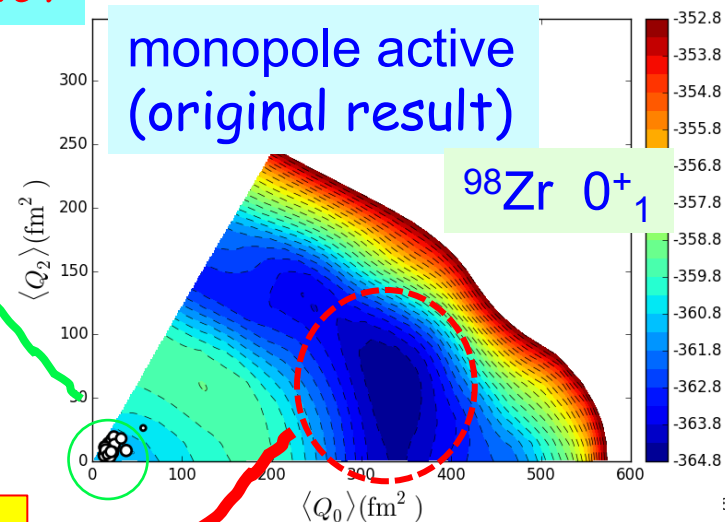
PES with T-plot

Frozen monopole

spherical
ground state
do not change
(overlap ~ 0.98)

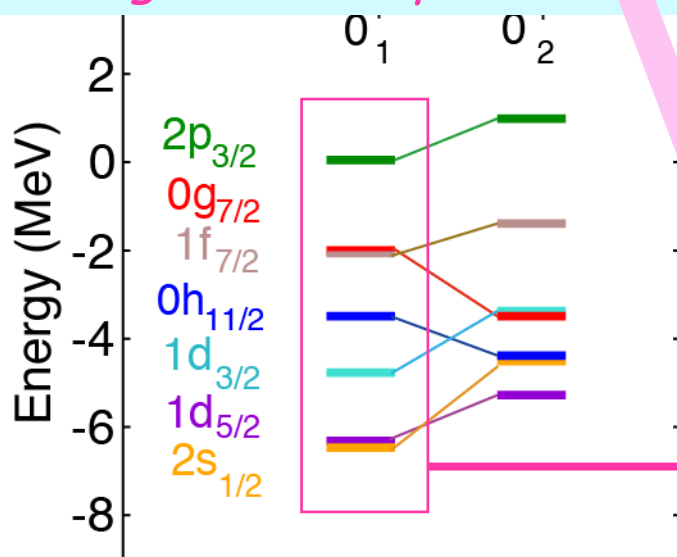


monopole active (original result)

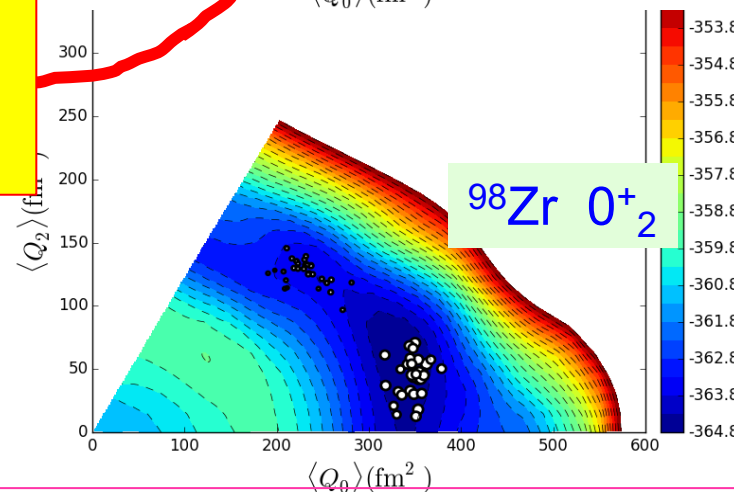


Effective SPE

- configuration dependent

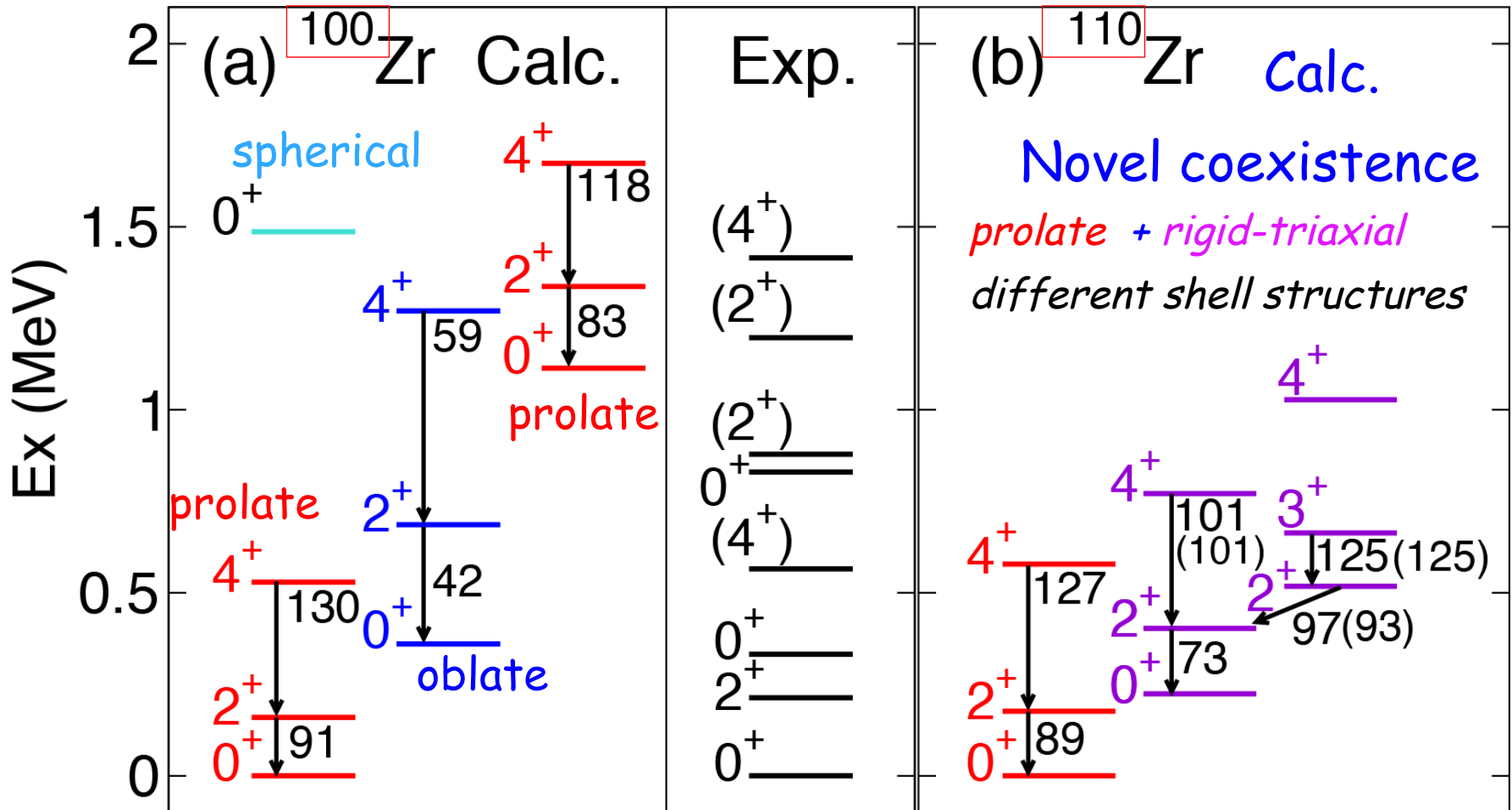


prolate
minimum
is gone



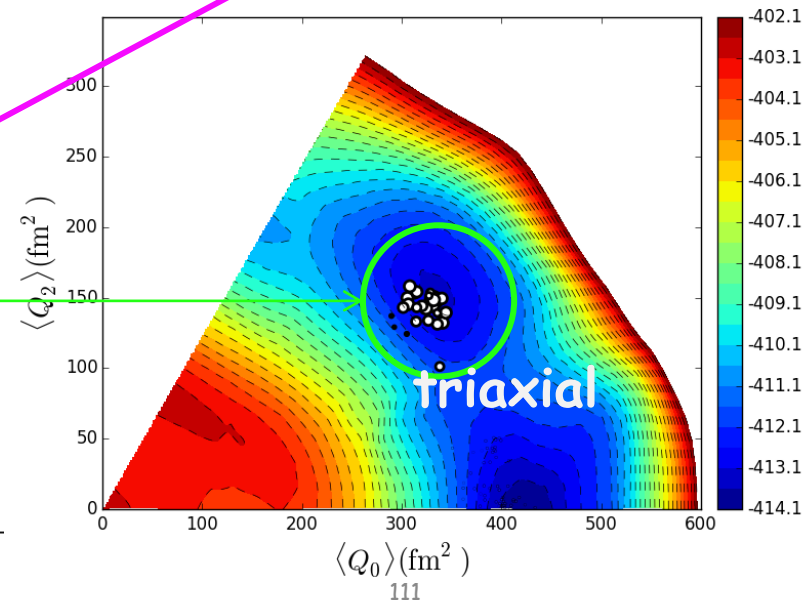
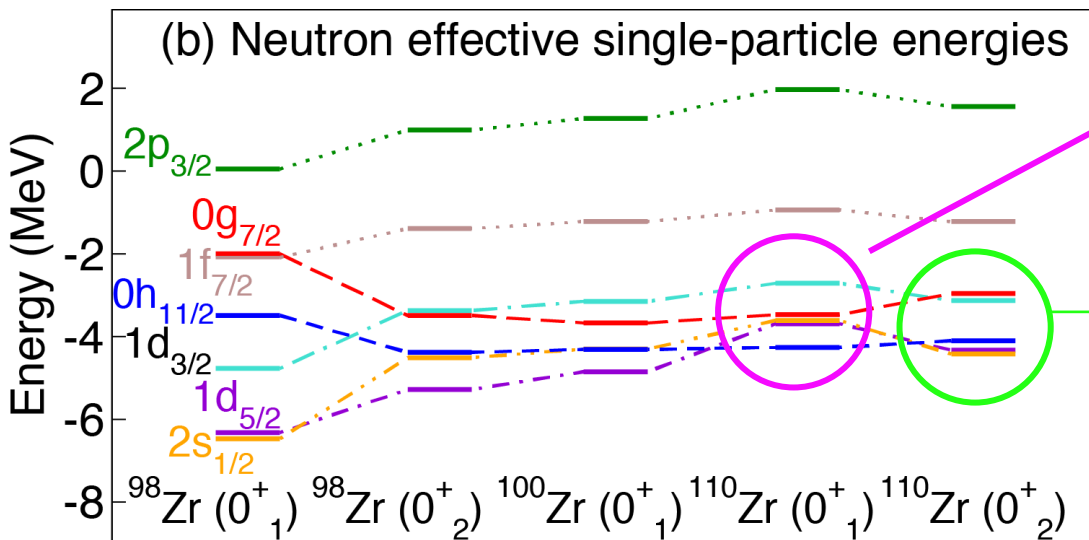
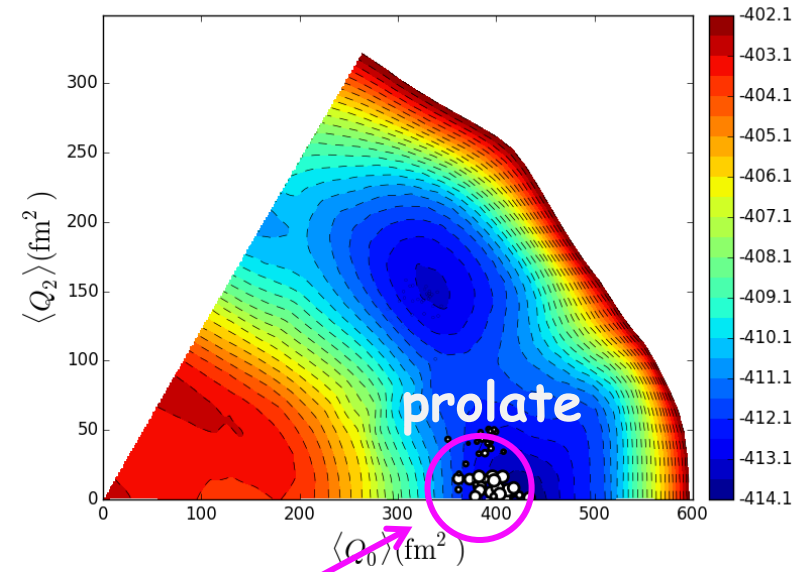
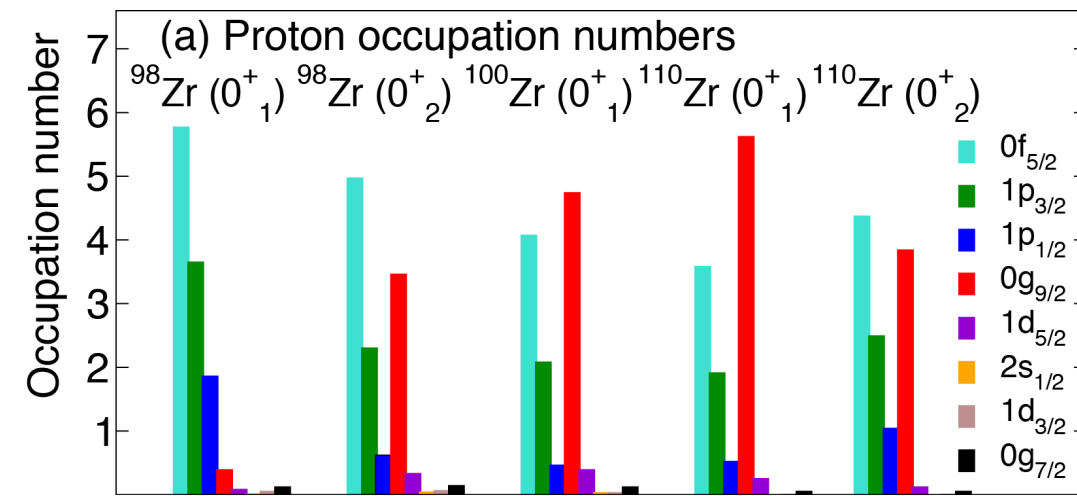
Use them as constant SPEs
independent of configurations, putting
monopole int. aside
→ Frozen monopole treatment

Prolate - rigid-triaxial shape coexistence



() : Rigid-triaxial rotor with $\gamma=28$ degrees
normalized at $2^+_2 \rightarrow 0^+_2$

different shell structures ~ like "different nuclei"



After mastering the shell model,
three (possible) pillars combined for future

computation

Monte Carlo
Shell Model
(MCSM)

(almost)
unlimited
dimensionality

massive
parallel
computers

Hamiltonian

pf
pfg9d5 (A3DA) (Ni)
8+8 on ^{56}Ni core (Zr)
8+8 on ^{80}Zr core (Sn)
8+10 on ^{132}Sn core
(Sm)

...
island of stability
+
 χ EFT based
(multi-)shell int.

many-body dynamics

Shell evolution
(Type I & II)

Quantum Phase
Transition

Shape coexistence

Quantum
Self-organization

Remark on Fermi liquid picture of nuclei

Naïve Fermi liquid picture (a la Landau) is revised, as atomic nuclei are not necessarily like simple solid vases containing almost free nucleons.

Nuclear forces are rich enough to optimize single-particle energies for each eigenstate (especially in the cases of collective-mode states), as referred to as **quantum self-organization**.

The **quantum self-organization** produces sizable effects with

- (i) **two quantum fluids** (protons and neutrons),
- (ii) **two major forces** : *e.g.*, quadrupole interaction to **drive collective mode**
monopole interaction to **control resistance**

Type II shell evolution is one of the most visible cases of the **quantum self-organization**, with massive p-h excitations across the shell gap. Quantum phase transition, shape coexistence, various deformation, fission, ... are related to the **quantum self-organization**.

The beauty of the collective modes is enhanced. Time-dependent version for reactions is of great interest (beyond thermalisation *etc.*).

The microscopic foundation of the IBM (Interacting Boson Model) is also related, for instance, regarding the origin of the Majorana interaction.

Be ambitious and deepen nuclear physics !

ご清聴
ありがとうございました