PRIMORDIAL GRAVITATIONAL WAVES AS A PROMISING TEST FOR INFLATIONARY MODELS

Maria Chiara Guzzetti*, Nicola Bartolo, Michele Liguori and Sabino Matarrese

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GRAVITATIONAL WAVES FROM INFLATION

ANY inflationary model

QUANTUM FLUCTUATIONS

of the gravitational field



gravitational wave POWER SPECTRUM

$$P_T(k) = A_T(k_*) \left(\frac{k}{k_*}\right)^{n_T}$$

GRAVITATIONAL WAVES FROM INFLATION

primordial → **present time**

 $P_T(k) = A_T(k_*) \left(\frac{k}{k_*}\right)^{n_T}$

present time gw spectral energy density

$$\Omega_{\text{GW}}(k, \tau_0) \equiv \frac{1}{\rho_{\text{c}}} \frac{\mathrm{d}\rho_{\text{gw}}}{\mathrm{d}\ln k} = \frac{1}{12} \left(\frac{k}{aH}\right)^2 T(k) P_{\text{T}}(k)$$

GRAVITATIONAL WAVES FROM INFLATION

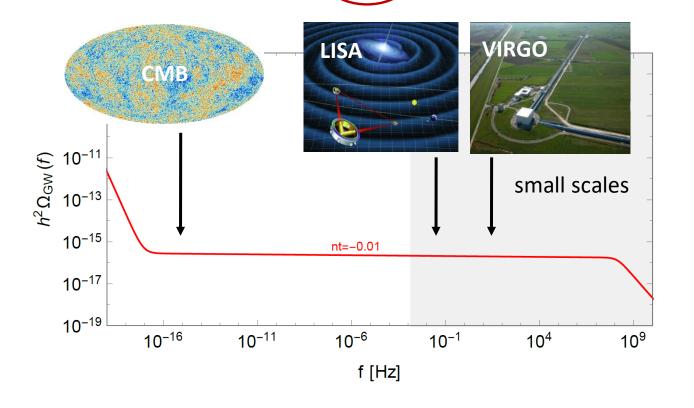
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stochastic gravitational wave **background**



GRAVITATIONAL WAVES FROM INFLATION

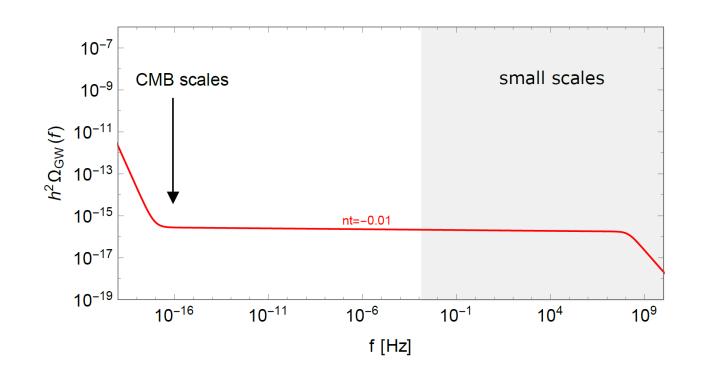
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stochastic gravitational wave **background**



GRAVITATIONAL WAVES FROM INFLATION

primordial \rightarrow present time

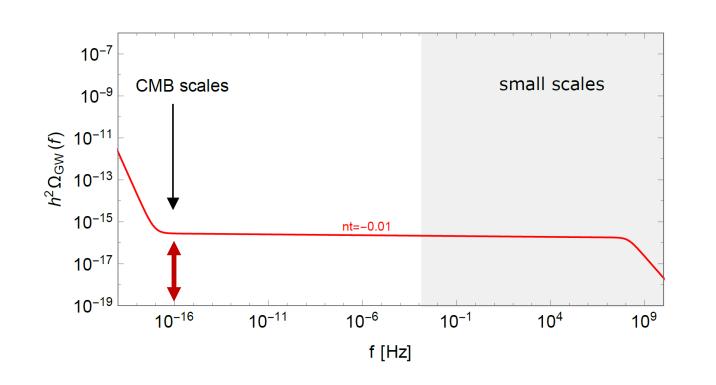
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stochastic gravitational wave **background**

• AMPLITUDE $A_{T}\left(k_{*}\right)$



GRAVITATIONAL WAVES FROM INFLATION

primordial → **present time**

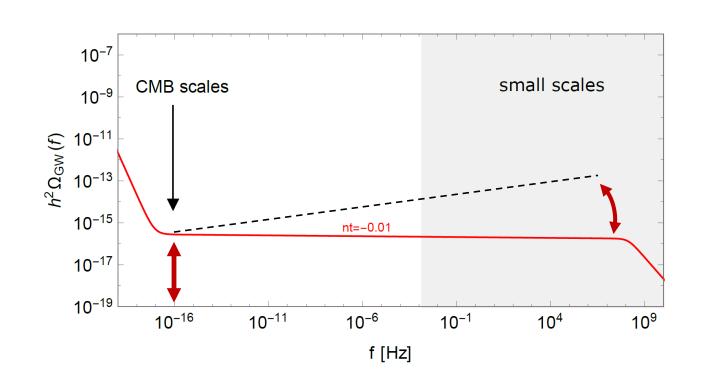
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stochastic gravitational wave **background**

- AMPLITUDE $A_{T}\left(k_{*}\right)$
- SPECTRAL INDEX $\,\eta_t$



FURTHER MECHANISMS OF GW PRODUCTION

ANY inflationary model ——— quantum fluctuations of the gravitational field

POSSIBLE EXTRA PRODUCTION

due to further fields besides the gravitational one

CLASSICAL gw production

$$h_{ij}^{\prime\prime}+2\mathcal{H}h_{ij}^{\prime}-
abla^2h_{ij}=rac{2}{M_{
m pl}^2}\hat{\Pi}_{ij}^{lm}T_{lm}$$
 source term

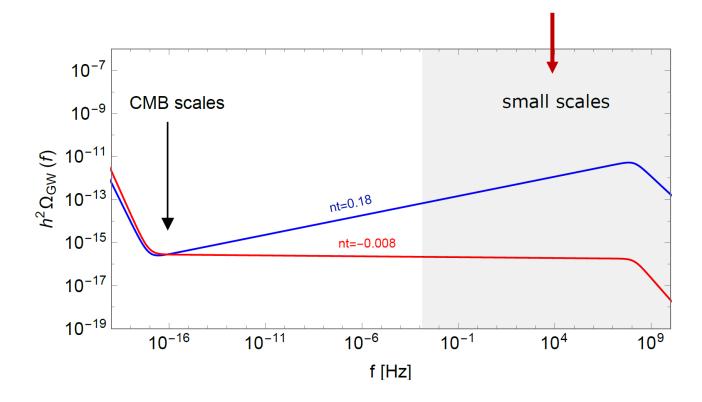
$$\square \qquad \qquad \Omega_{\rm gw} = \Omega_{\rm gw}^{vacuum} + \Omega_{\rm gw}^{sourced}$$

interferometers

INFLATIONARY GRAVITATIONAL WAVES

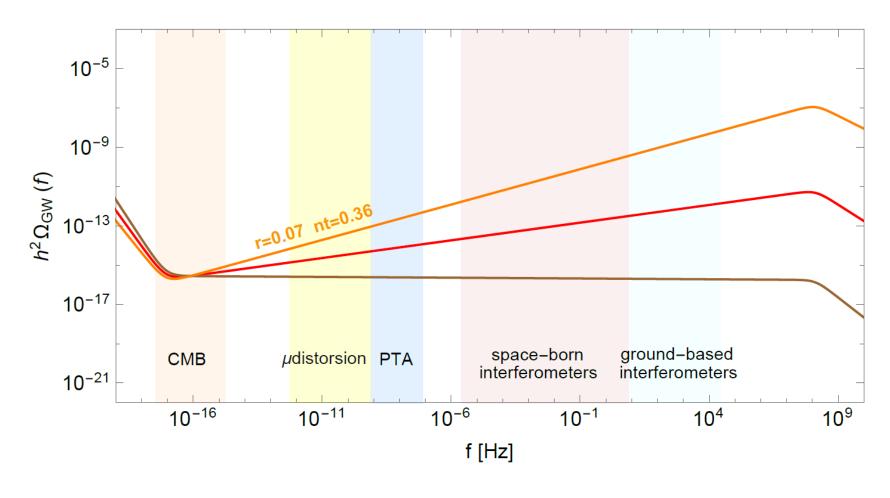
$$\Omega_{\rm gw} = \Omega_{\rm gw}^{vacuum} + \Omega_{\rm gw}^{sourced}$$

GW PRODUCTION	Discriminant	Specific discriminant	Examples of specific models
Vacuum oscillations quantum fluctuations of the gravitational field stretched by the accelerated expansion	theory of gravity	General Relativity	single-field slow-roll
		Golden Testering	all other models in GR
			G-Inflation
		MG/EFT approach	Potential-driven G-Inflation
			EFT approach
Classical production second-order GW get virtal by the presence on a surface term in GW equation of motion	source term	a juk maton fluctuations	all models
		fluctuations of extra scalar fields	inflaton+spectator fields
			curvaton
		gauge particle production	pseudoscalar inflaton+gauge field
			scalar infl.+pseudoscalar+gauge
		scalar particle production	scalar inflaton+ scalar field
		particle production during preheating	chaotic inflation
		r Francisco de la constante de la constan	hybrid inflation

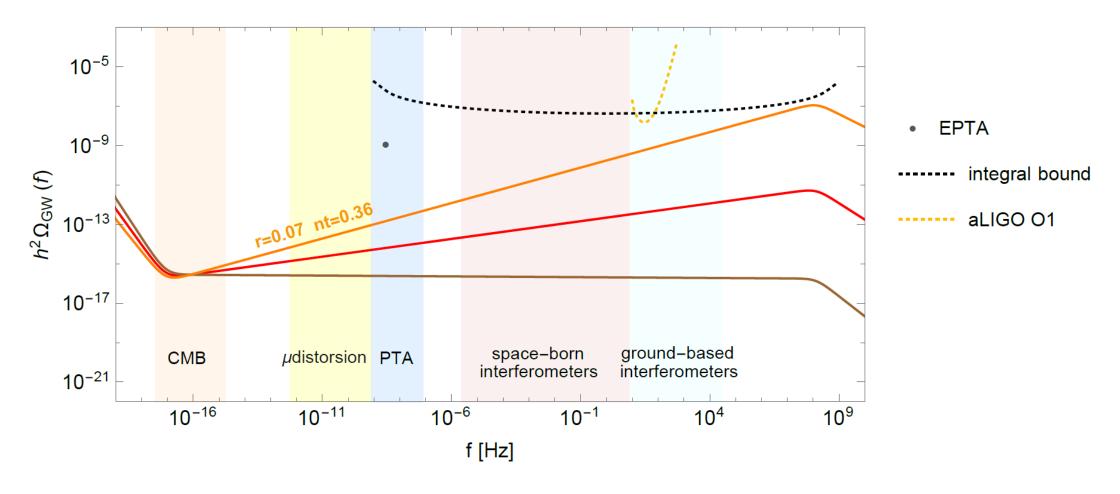


CURRENT BOUNDS AND OBSERVATIONAL PROSPECTS

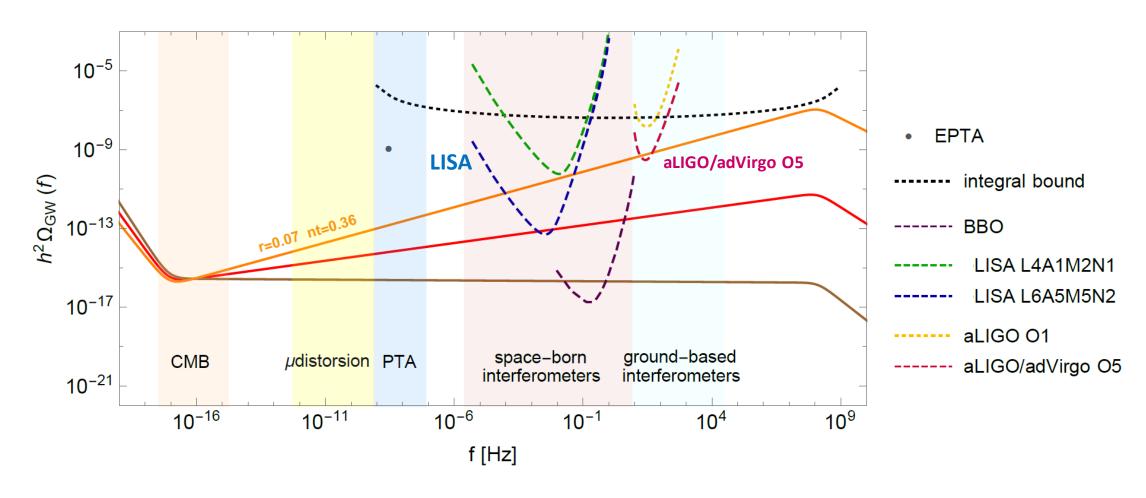
CURRENT BOUNDS



CURRENT BOUNDS



CURRENT BOUNDS AND OBSERVATIONAL PROSPECTS



PROSPECTS FOR INFLATIONARY PHYSICS

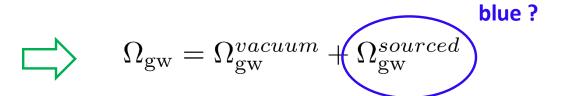
- constraints on a specific inflationary model
- test of the inflationary consistency relation

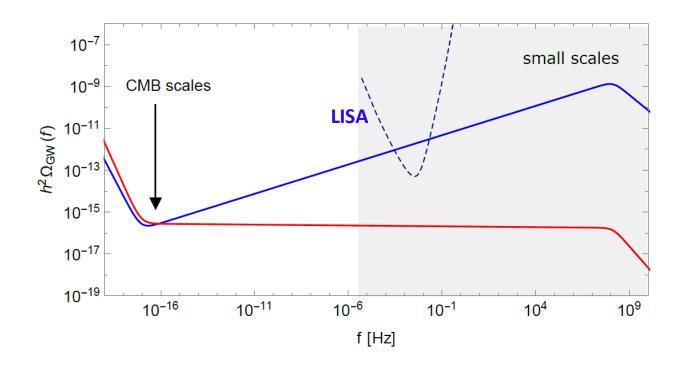
INFLATION WITH SPECTATOR FIELD

inflaton + spectator field $\,\sigma\,$

INFLATION WITH SPECTATOR FIELD

inflaton + spectator field $\,\sigma\,$





INFLATION WITH SPECTATOR FIELD

inflaton + spectator field σ



$$\Omega_{\mathrm{gw}} = \Omega_{\mathrm{gw}}^{vacuum} + \Omega_{\mathrm{gw}}^{sourced}$$
 blue ?

$$\begin{array}{ccc} C_S & & \text{gw} \\ \sigma \text{ speed of sound} & & \text{amplitude} \end{array}$$

$$s=\dot{c}_s/Hc_s$$
 \longrightarrow gw spectral index evolution

INFLATION WITH SPECTATOR FIELD

inflaton + spectator field σ

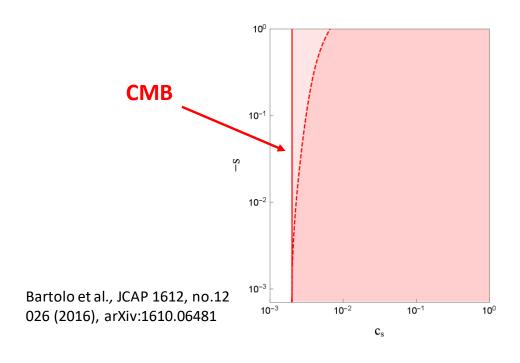


$$\Omega_{\rm gw} = \Omega_{\rm gw}^{vacuum} + \Omega_{\rm gw}^{sourced}$$
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$$s = \dot{c}_s/Hc_s \,\, \longrightarrow \,\, \mathop{\rm gw}_{\rm spectral \, index}$$

speed of sound evolution

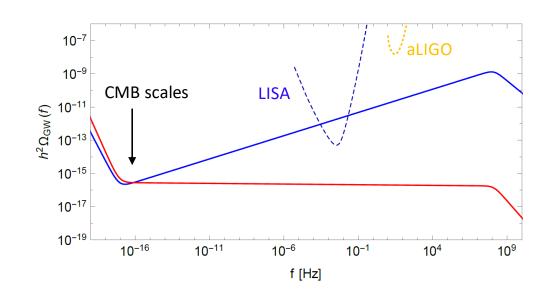


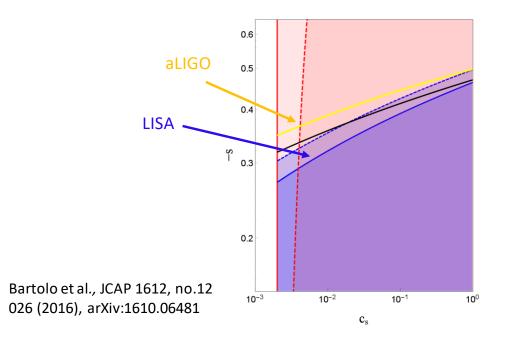
INFLATION WITH SPECTATOR FIELD

inflaton + spectator field σ



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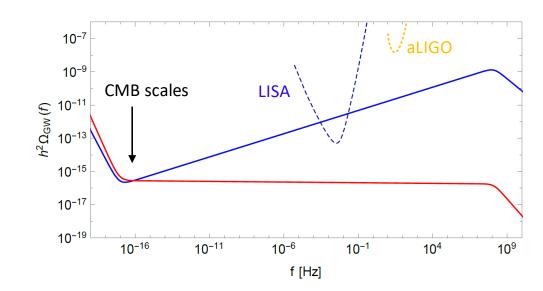


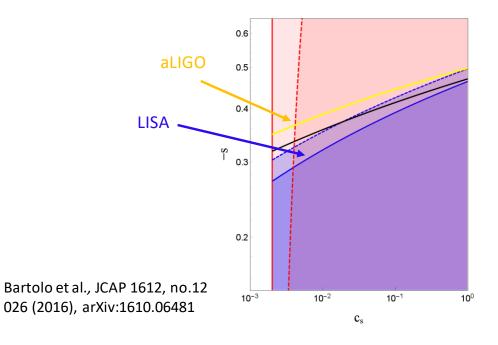
INFLATION WITH SPECTATOR FIELD

inflaton + spectator field σ



$$\Omega_{\mathrm{gw}} = \Omega_{\mathrm{gw}}^{vacuum} + \Omega_{\mathrm{gw}}^{sourced}$$
 blue ?





experiments at small scales improve constraints on specific inflationary models, even in case of a non-detection

single-field slow-roll inflation

(vacuum fluctuations)

$$r = -8n_T$$

single-field slow-roll inflation

(vacuum fluctuations)

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other inflationary models

VIOLATION?

	Model	Tensor power-spectrum	Tensor spectral inde	x	Consistency relation
GW signal due to vacuum fluctuations	Standard infl.	$P_{\rm T} = \frac{8}{M_{\rm pl}^2} \left(\frac{H}{2\pi}\right)^2$	$n_{\mathrm{T}} = -2\epsilon$	red	$r = -8n_{\mathrm{T}}$
	EFT inflation ^(a)	$P_{\rm T} = \frac{8}{c_{\rm T} M_{\rm pl}^2} \left(\frac{H}{2\pi}\right)^2$	$n_{\rm T} = -2\epsilon + \frac{2}{3} \frac{m_{\rm T}^2}{\alpha H^2} \left(1 + \frac{4}{3} \epsilon \right)$	r/b	-
	EFT inflation ^(b)	$P_{\mathrm{T}} = \frac{8}{c_{\mathrm{T}} M_{\mathrm{pl}}^2} \frac{2^{\frac{-p}{1+p}}}{\pi} \Gamma^2 \left(\frac{1}{2(1+p)}\right) \left(\frac{H}{2\pi}\right)^2$	$n_{\mathrm{T}} = \frac{p}{1+p}$	blue	violation
	Gen. G-Infl.	$P_{\mathrm{T}} = \frac{8}{M_{\mathrm{pl}}^2} \gamma_{\mathrm{T}} \frac{\mathscr{G}_{\mathrm{T}}^{1/2}}{\mathscr{F}_{\mathrm{T}}^{3/2}} \left(\frac{H}{2\pi}\right)^2$	$n_{\rm T} = 3 - 2\nu_{\rm T}$	r/b	-
	Potdriv. G-Infl.	$P_{\mathrm{T}} = rac{8}{M_{\mathrm{pl}}^2} \left(rac{H}{2\pi} ight)^2$	$n_{\mathrm{T}} = -2\epsilon$	r/b	$r \simeq -\frac{32\sqrt{6}}{9}n_{\mathrm{T}}$
Extra GW signal due to a source term	Particle prod.	$P_{\rm T}^+ = 8.6 \times 10^{-7} \frac{4H^2}{M_{\rm pl}^2} \left(\frac{H}{2\pi}\right)^2 \frac{e^{4\pi\xi}}{\xi^6}$	-	blue	violation
	Spectator field	$P_{ m T} \simeq 3 rac{H^4}{c_{ m S}^{18/5} M_{ m pl}^4}$	$n_{\mathrm{T}} \simeq 2\left(\frac{2m^2}{3H^2} - 2\epsilon\right) - \frac{18}{5} \frac{\dot{c}_{\mathrm{S}}}{Hc_{\mathrm{S}}}$	r/b	violation

single-field slow-roll inflation

(vacuum fluctuations)

$$r = -8n_T$$

test
for single-field
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other inflationary models

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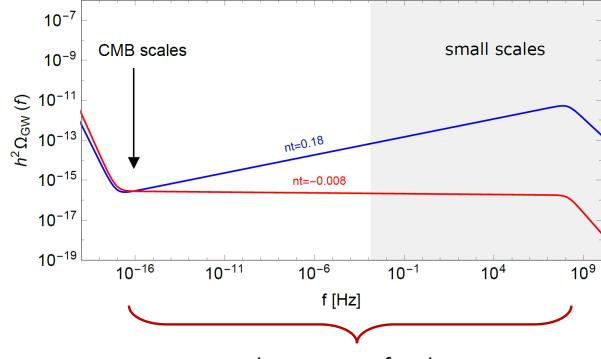
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experiments at small scales
are crucial in order to exploit
the long lever arm between CMB scales and
laser interferometers scales



large range of scales

single-field slow-roll inflation

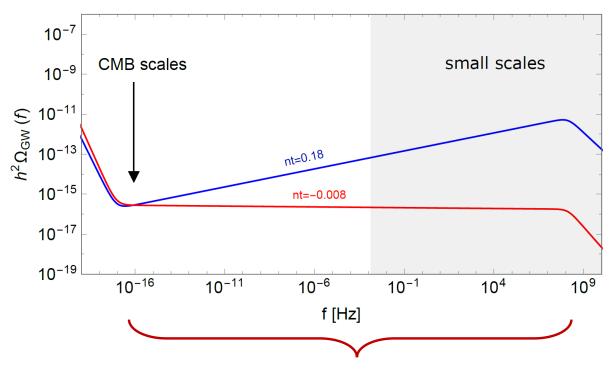
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the long lever arm between CMB scales and
laser interferometers scales

experiments at small scales improve the capabilities of testing the single-field slow-roll inflationary model



large range of scales

CONCLUSIONS

INFLATIONARY GW → interesting signal in order to investigate the inflationary physics

exploiting experiments of direct detection at small scales:

- improvement of constraints on power-spectrum parameters
- improvement of constraints of specific inflationary models
- significant role in testing the inflationary consistency relation

